

Searching for Low-Surface-Brightness Galaxies with the Hyper Suprime-Cam Survey

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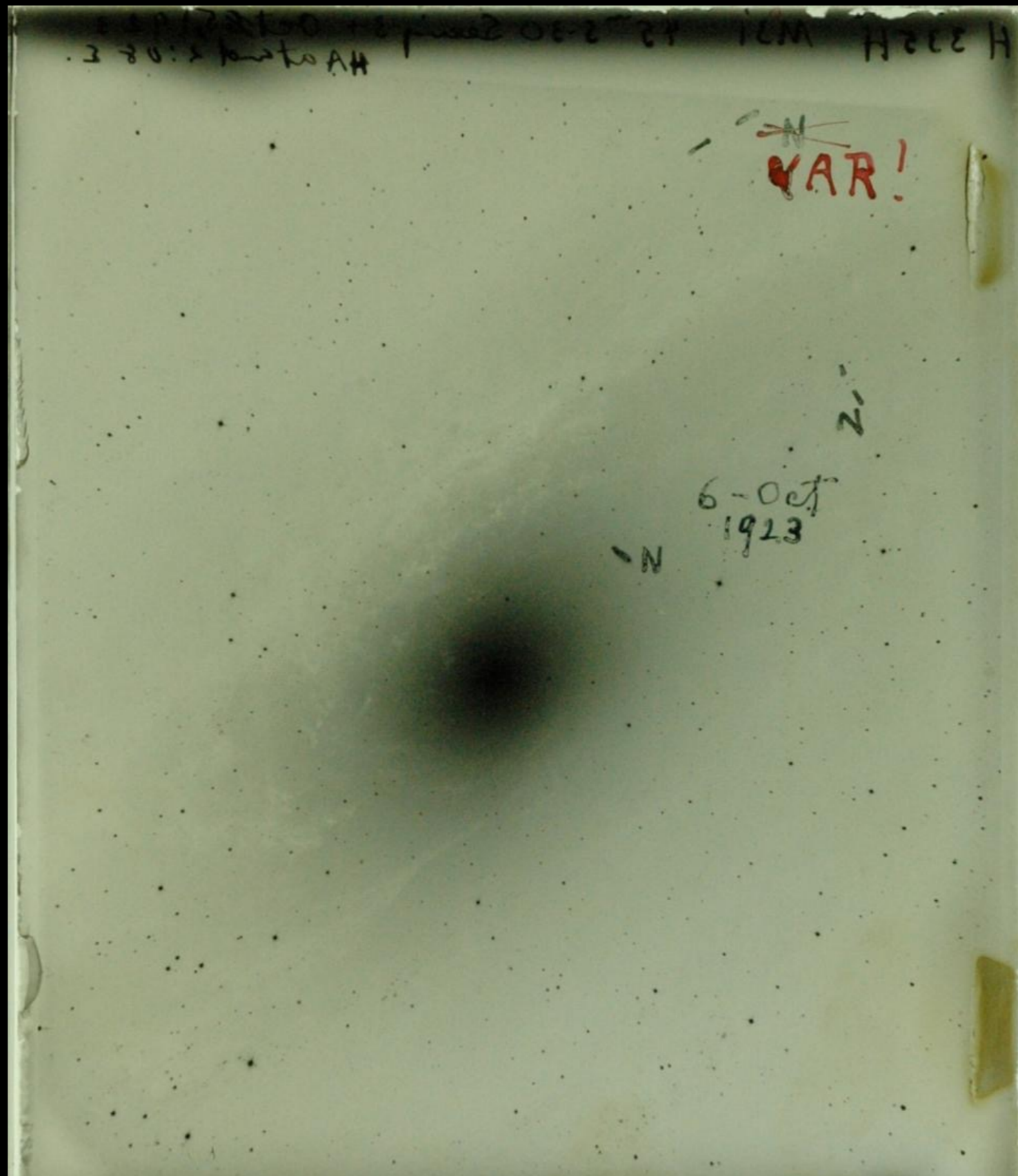
jgreco@astro.princeton.edu



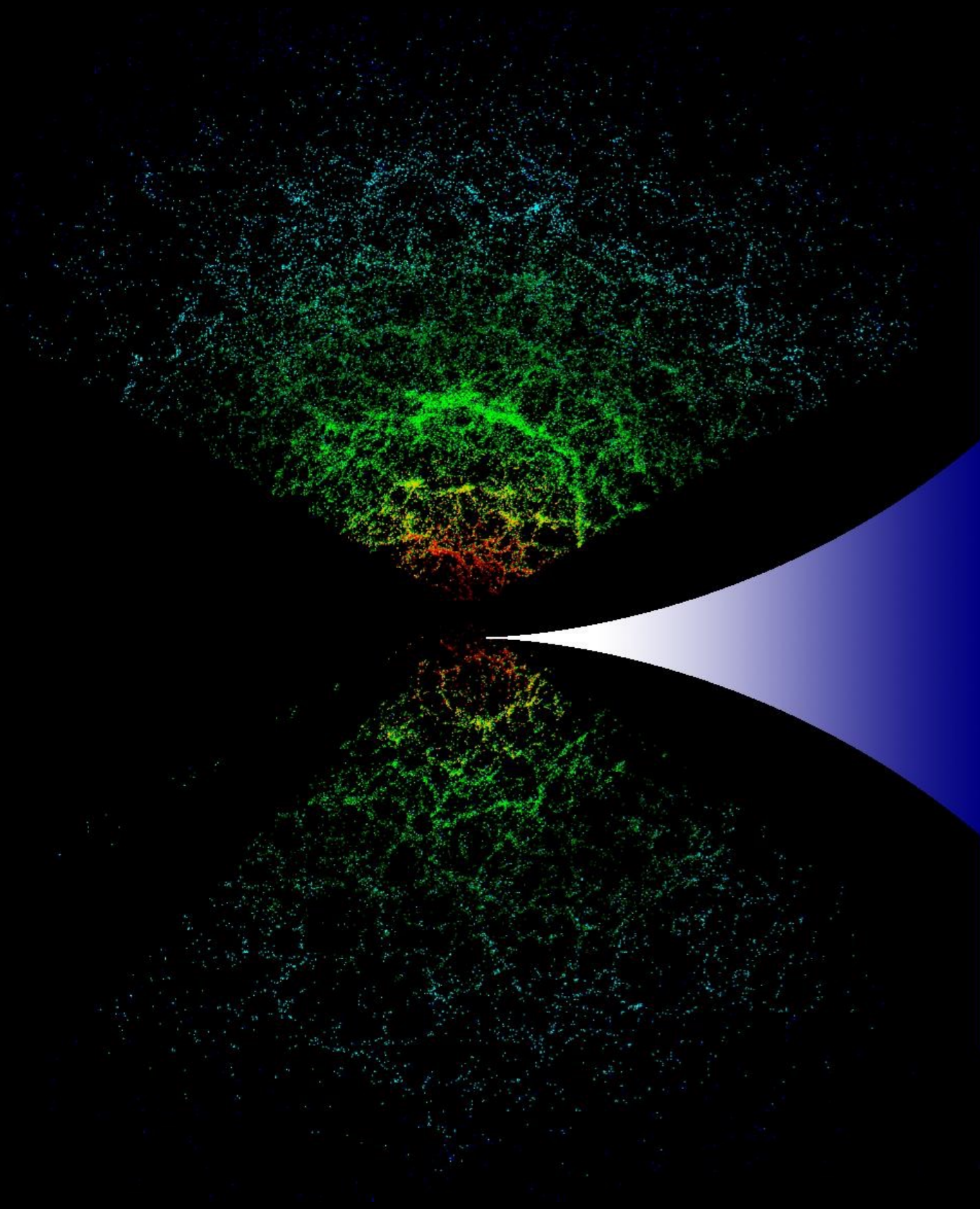
<https://github.com/johnnygreco>



Pre 1923: The **Realm** of the **Nebulae**



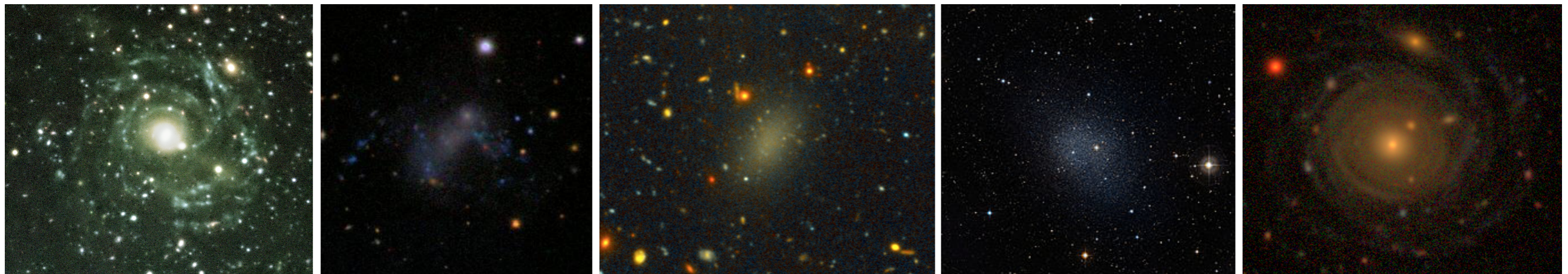
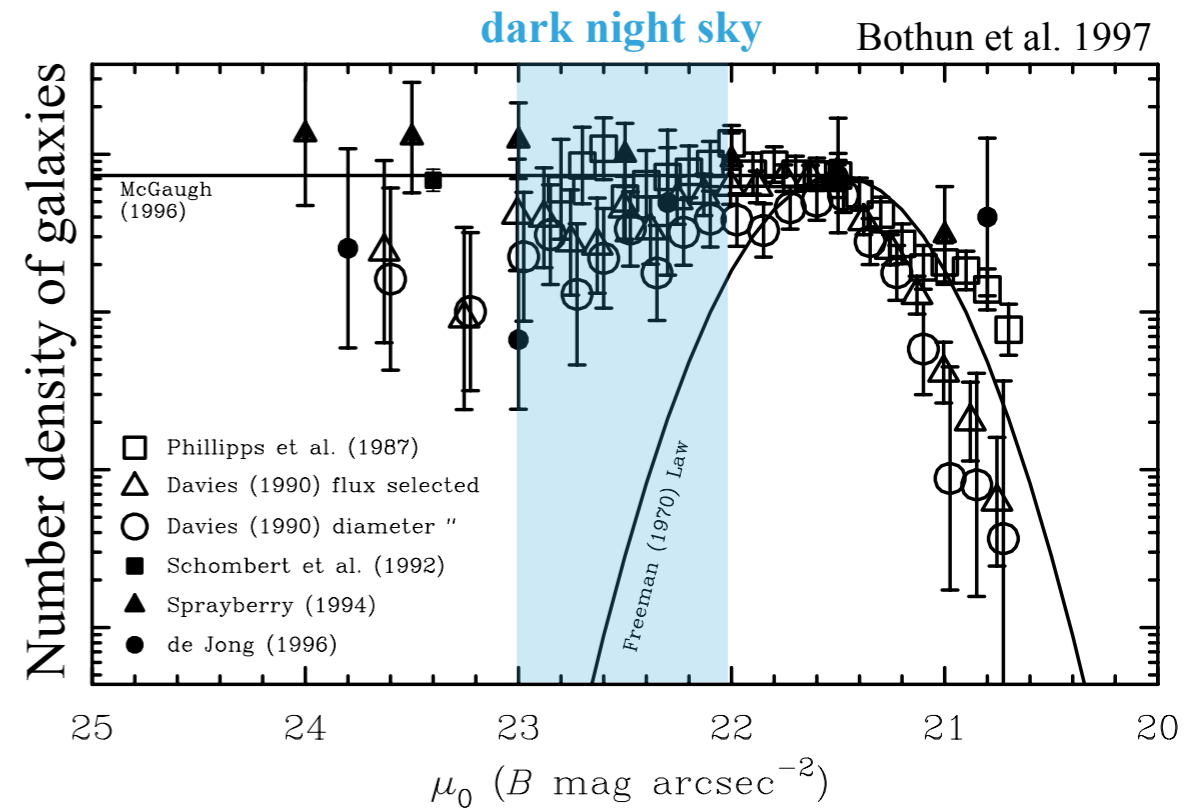
80 years later...



The **Hidden** Galaxy Population

Low-surface-brightness galaxies (LSBGs)

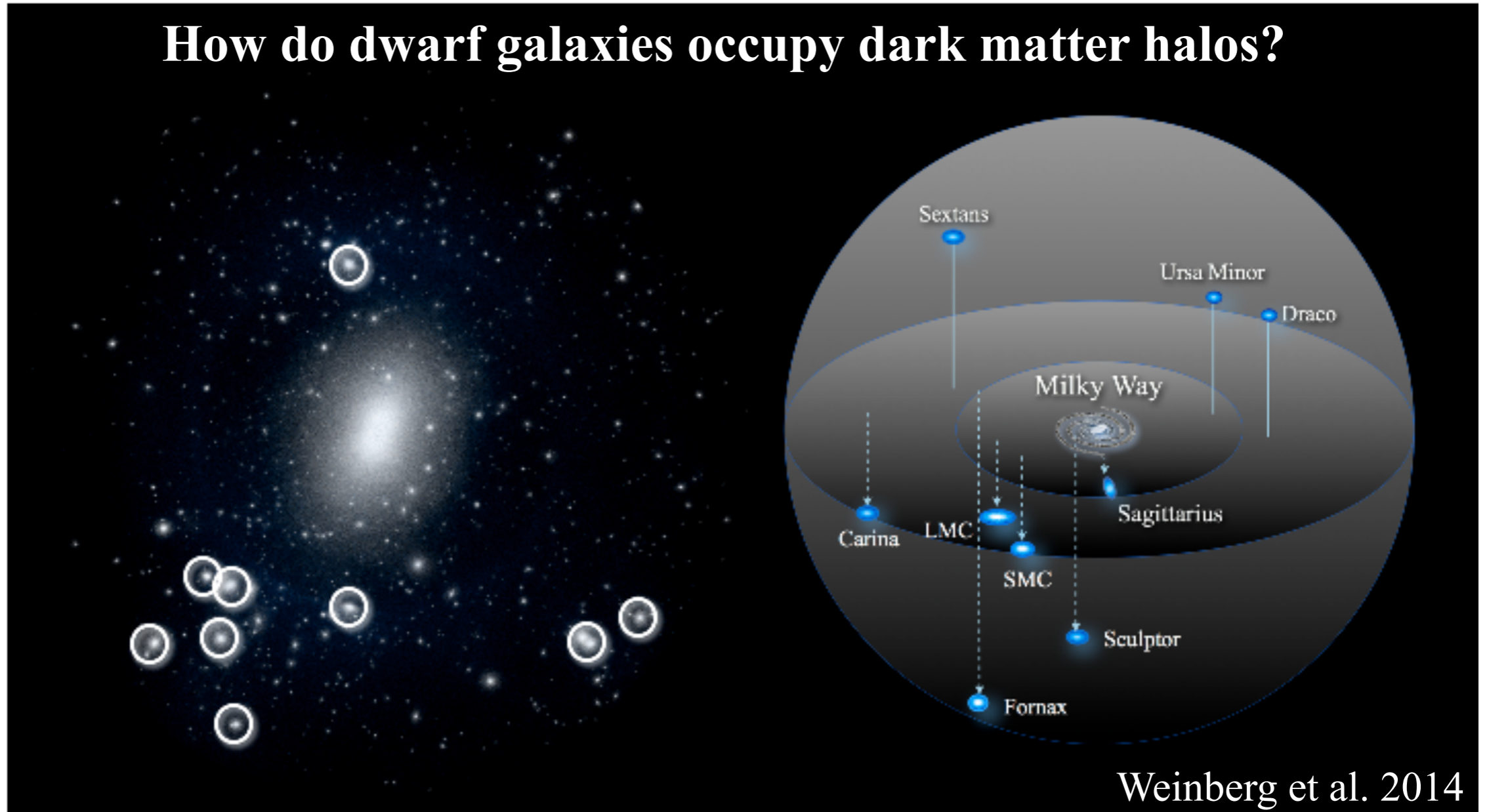
- Surface brightness **fainter than night sky**
- Span **all galaxy types** and environments
- Underrepresented in previous optical surveys



The **Hidden** Galaxy Population

LSBGs as a testing ground for LCDM...

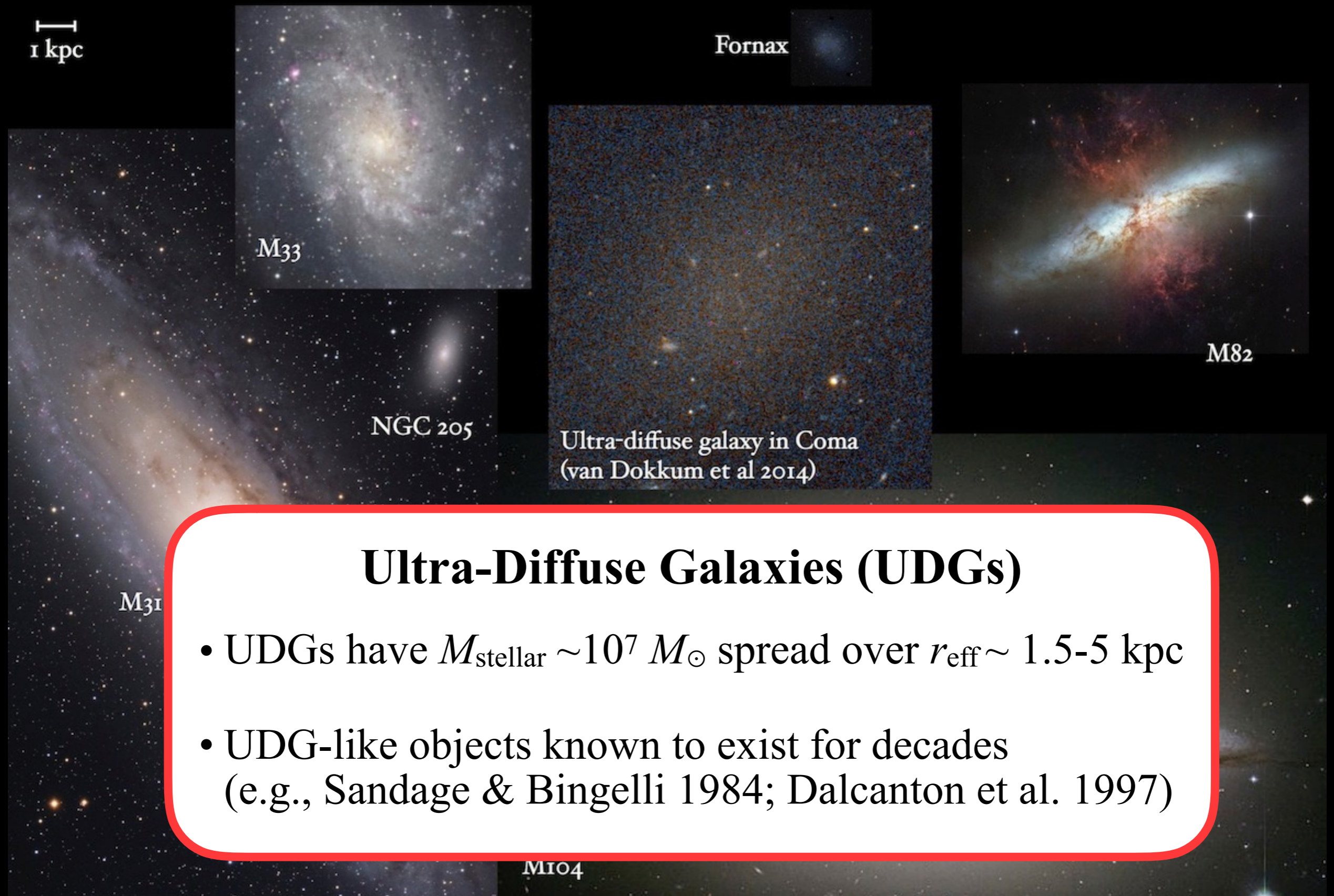
How do dwarf galaxies occupy dark matter halos?



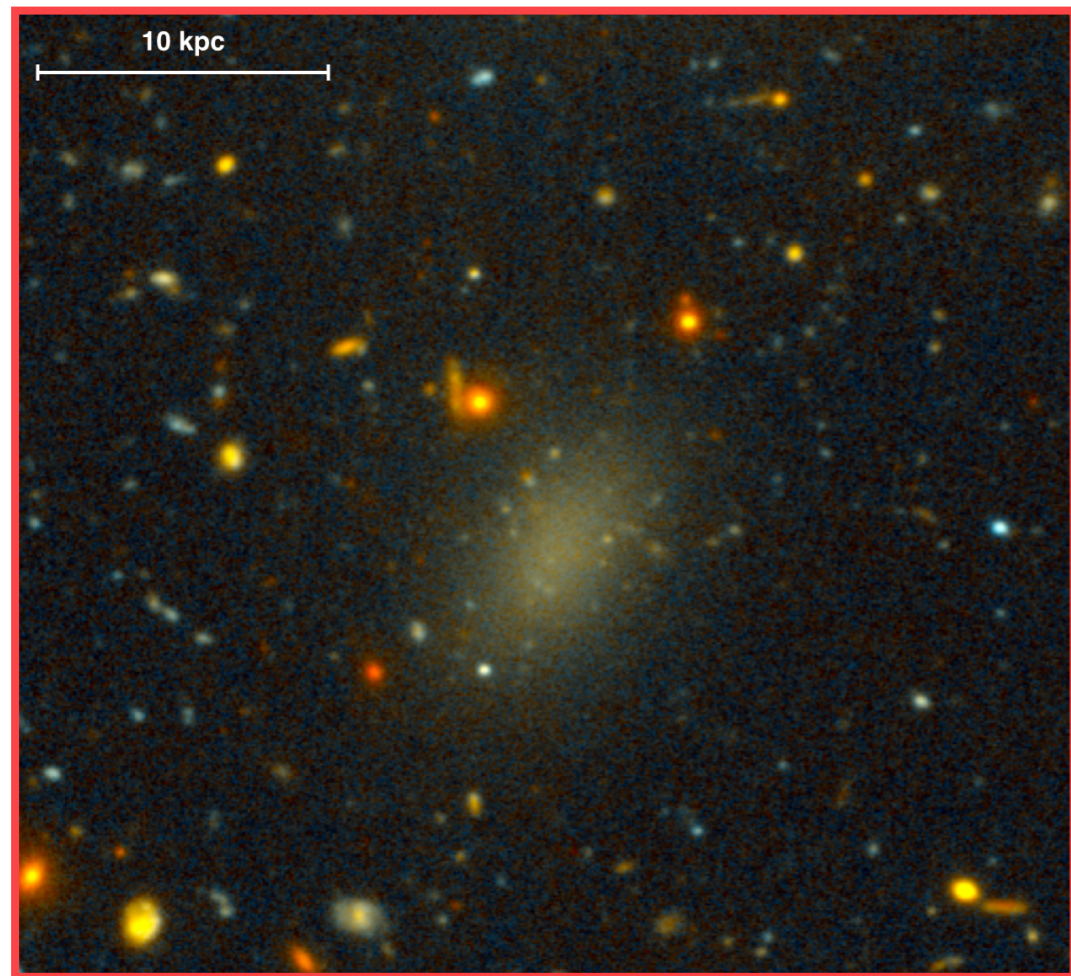
The Latest Craze: Ultra-Diffuse Galaxies (UDGs)



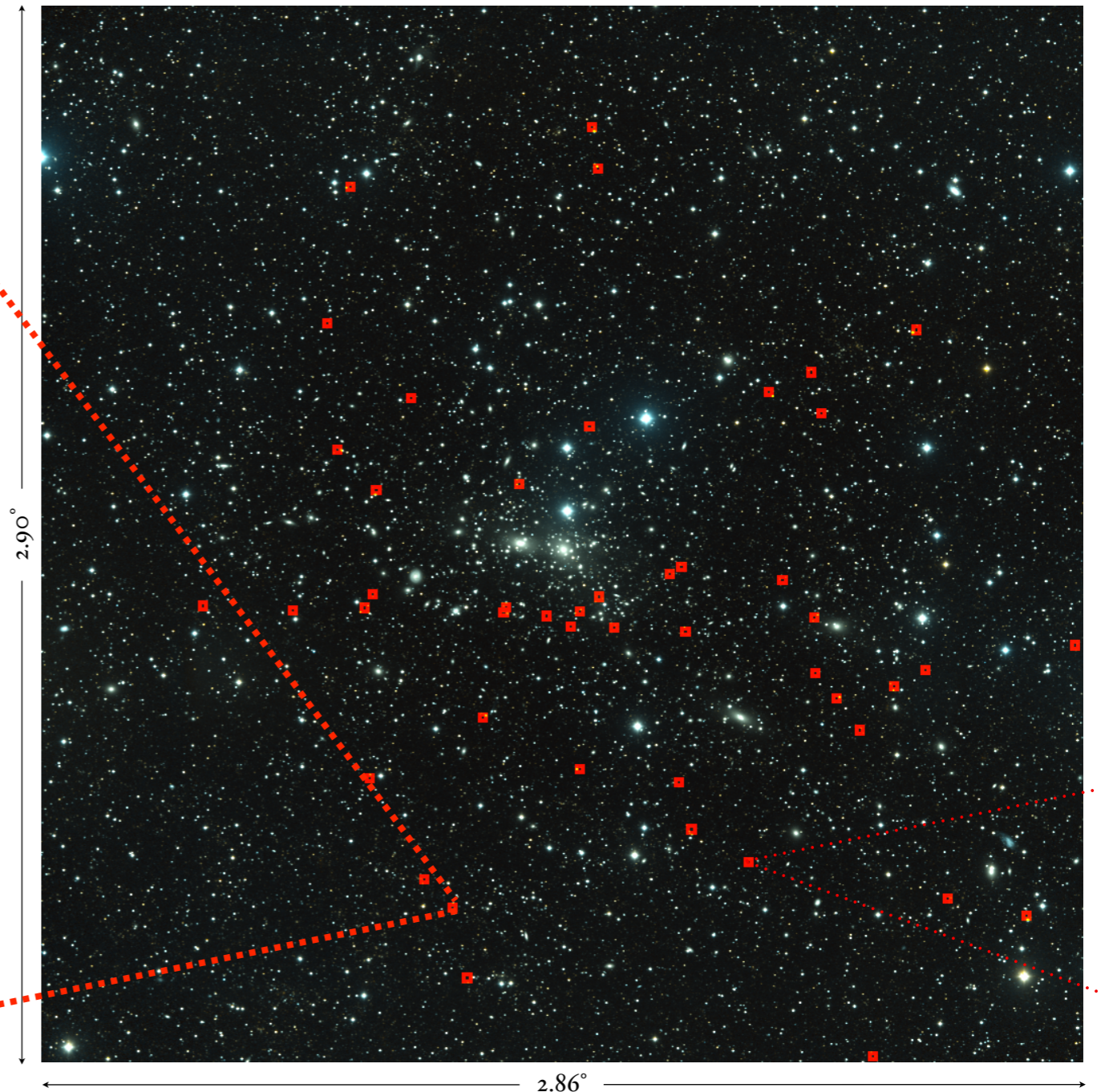
The Latest Craze: Ultra-Diffuse Galaxies (UDGs)



~1000 UDG candidates in Coma!



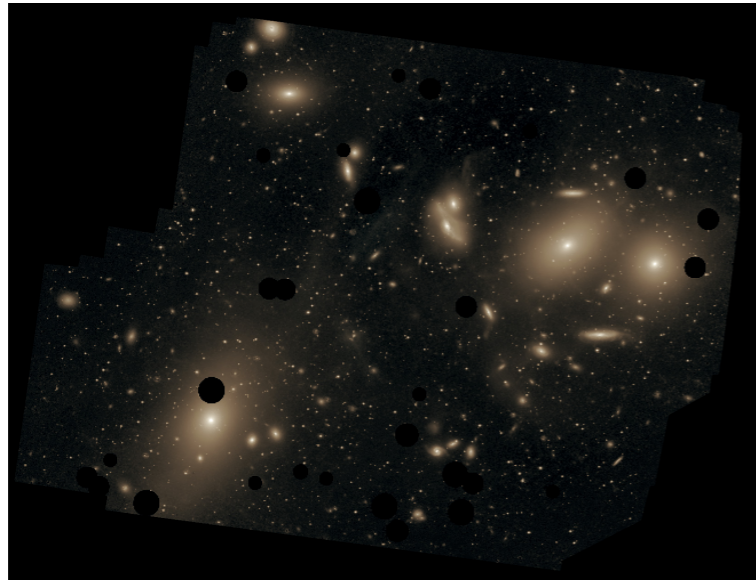
Dragonfly 44



Coma Cluster

UDGs common in rich clusters

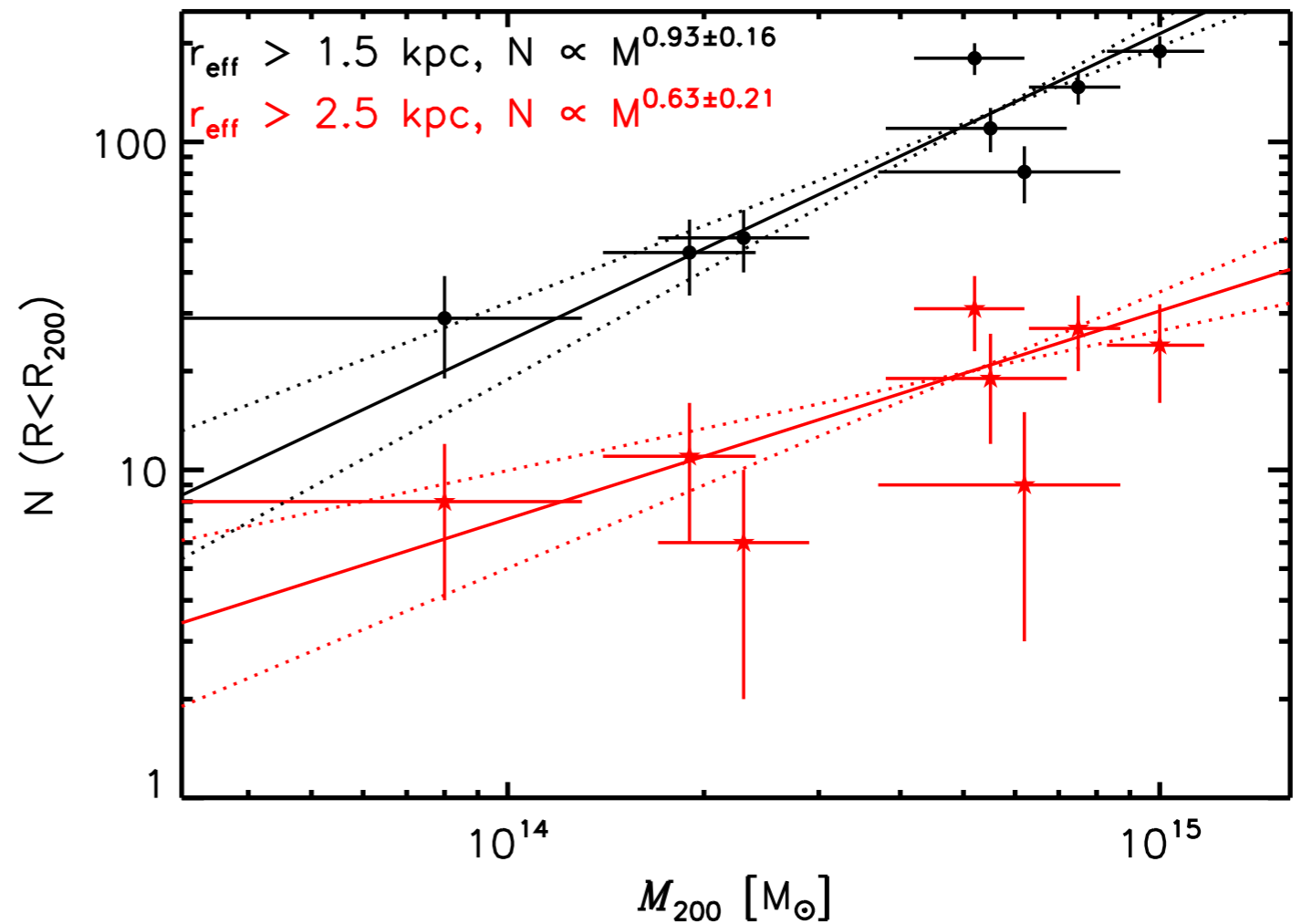
New UDGs also found in....



Virgo
(Mihos et al. 2015)



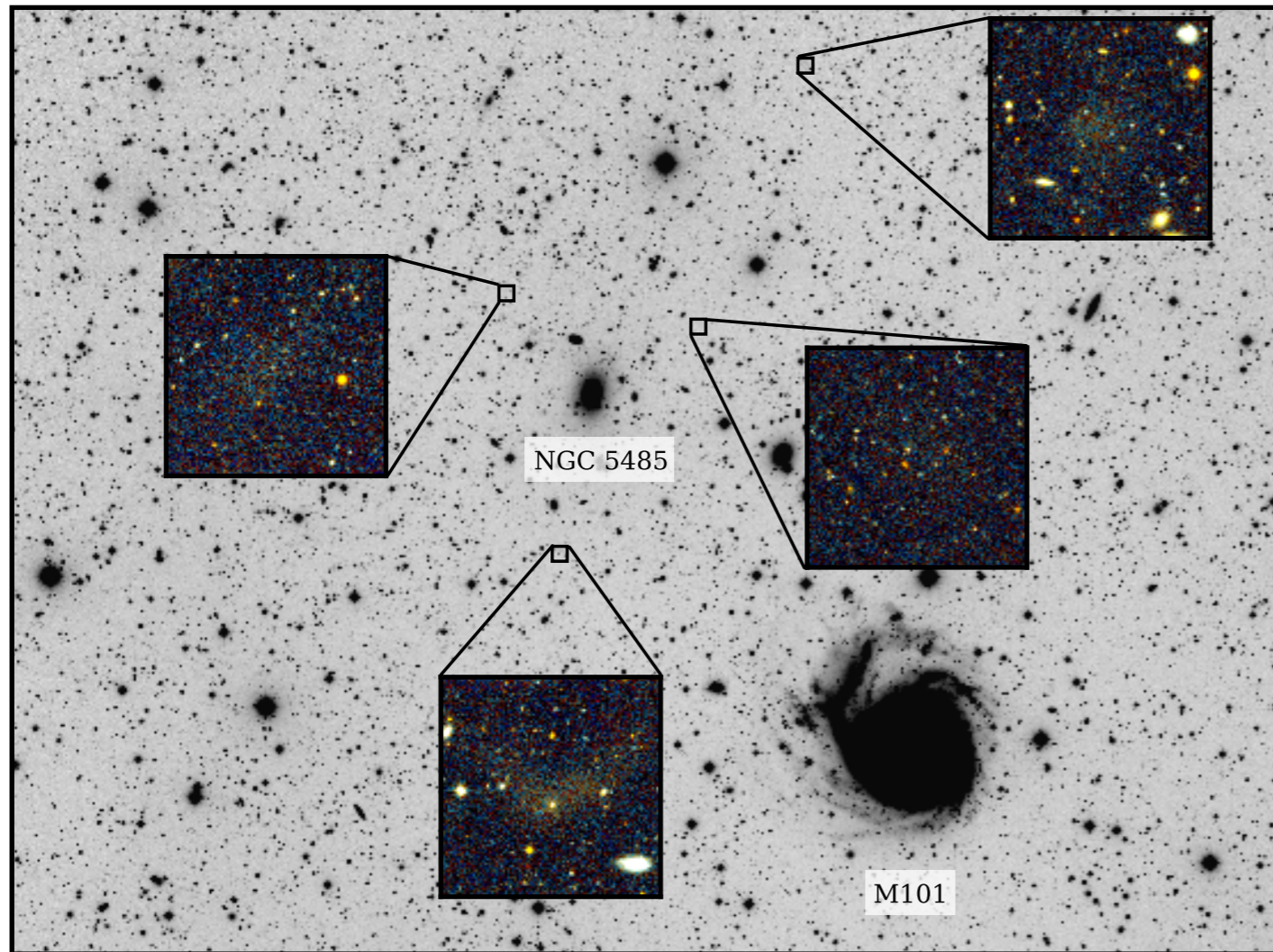
Fornax
(Munoz et al. 2015)



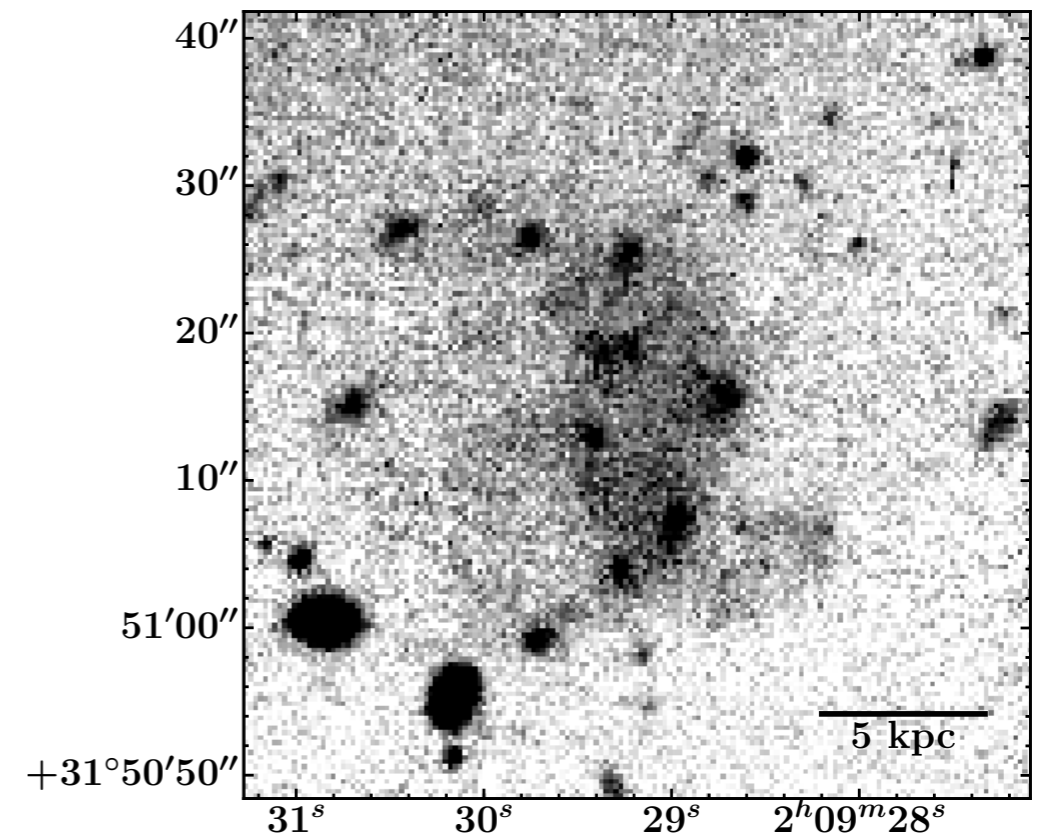
... and 8 low-z clusters:
(van der Burg et al. 2016)

UDGs in **groups** and the **field**

New UDGs also found in....



Small groups
(Merritt et al. 2016)



The field
(Leisman et al. 2017)

Need **deep** + **wide** blind search

Need **deep** + **wide** blind search



Ultra-LSB sources in SDSS

Need **deep** + **wide** blind search

A **new view** with the **Hyper Suprime-Cam (HSC)**

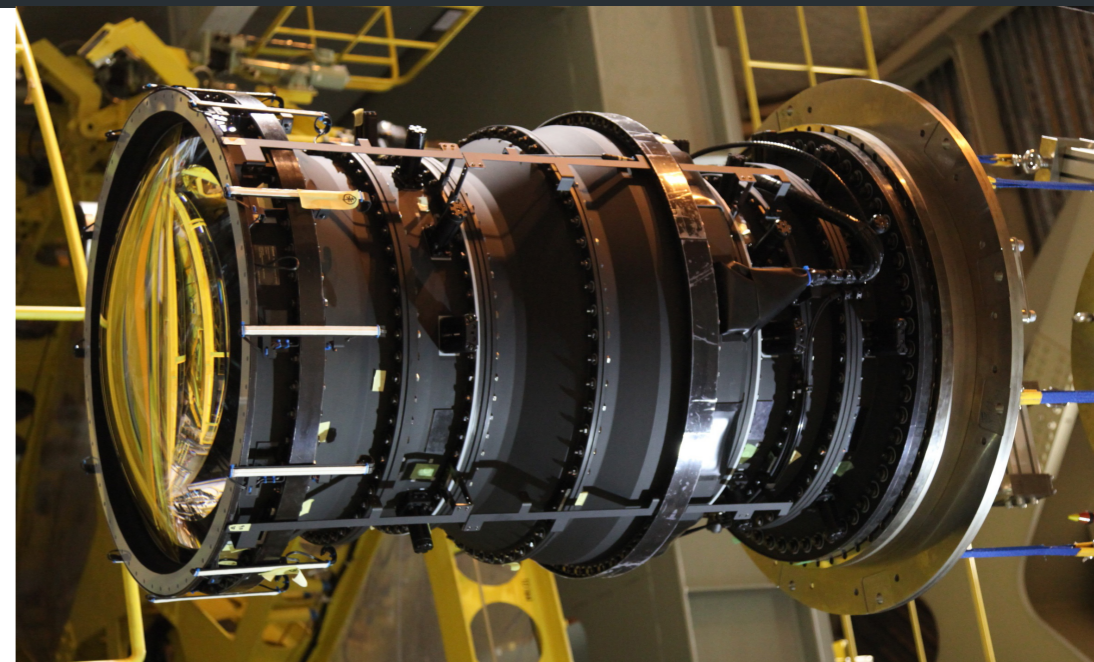


Ultra-LSB sources in SDSS

The **Hyper Suprime-Cam** Subaru Strategic Program



HSC Collaboration Meeting Kavli IPMU, Kashiwa, Japan (2016)



Hyper Suprime-Cam



Typical Apparent
Diameter of the
Moon (0.5 degrees)



Suprime-Cam

First Light Release
January 1999



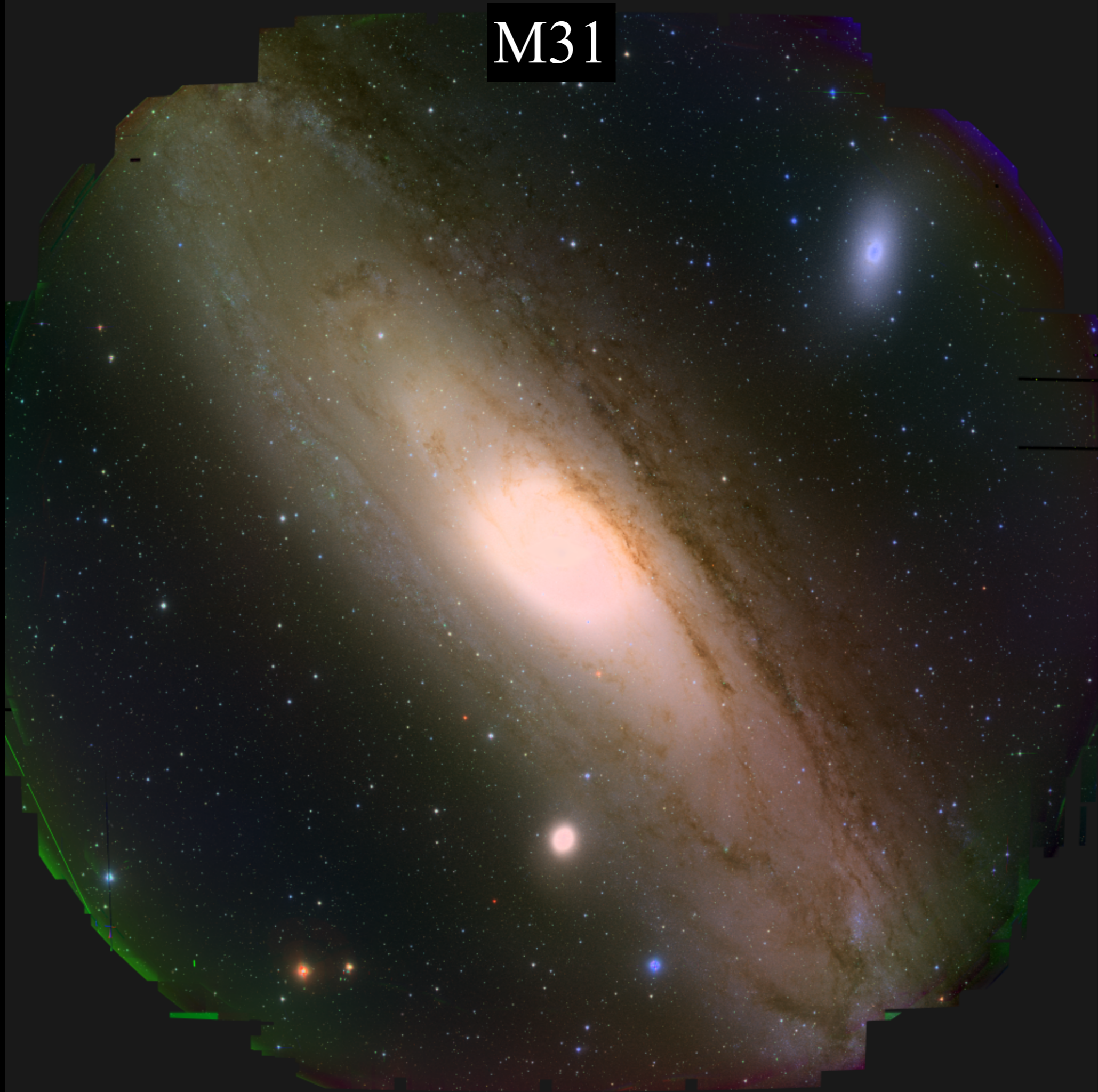
Suprime-Cam

Image Release
September 2001

Hyper Suprime-Cam

Image Release
July 2013

M31



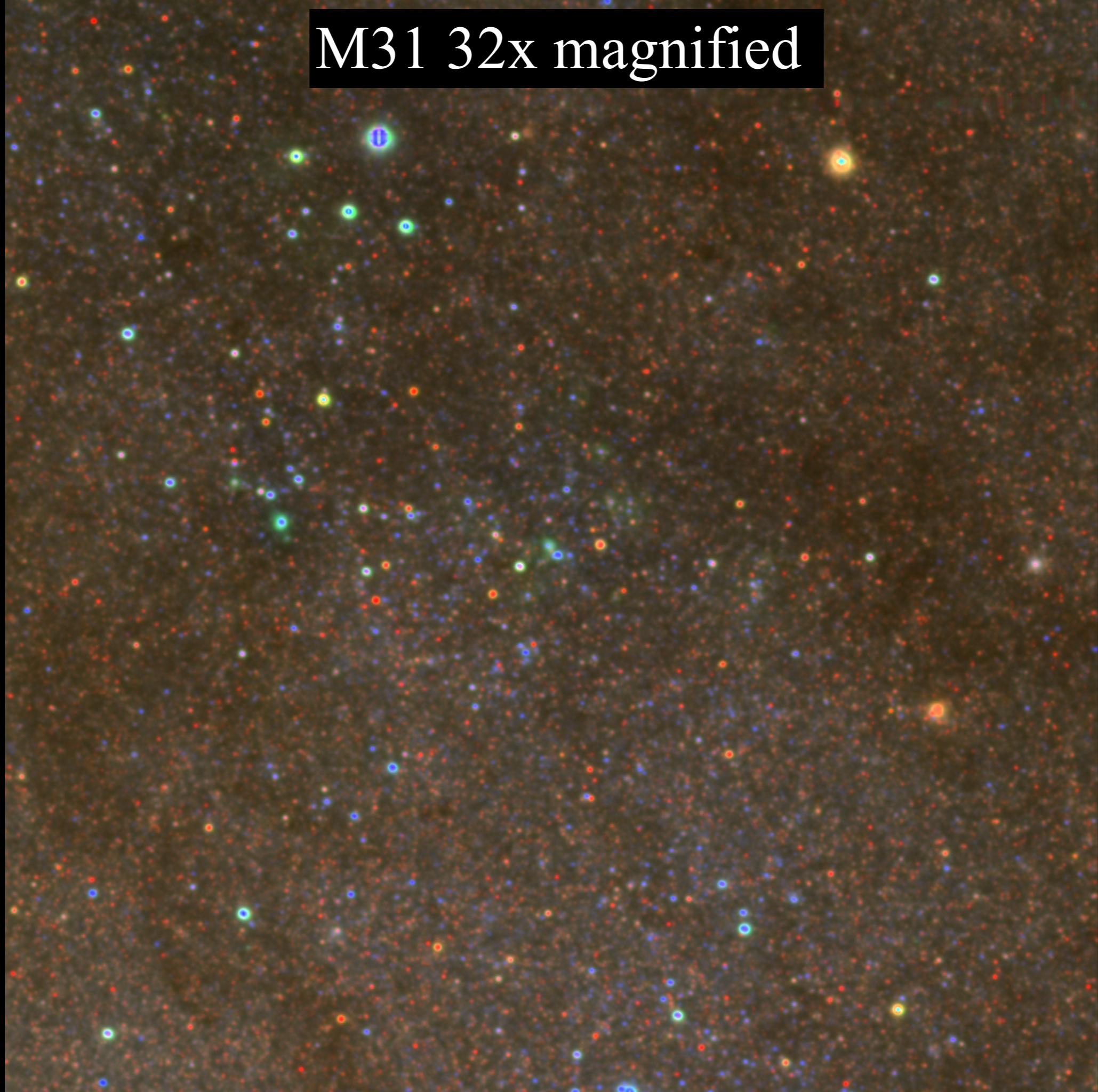
M31 2x magnified



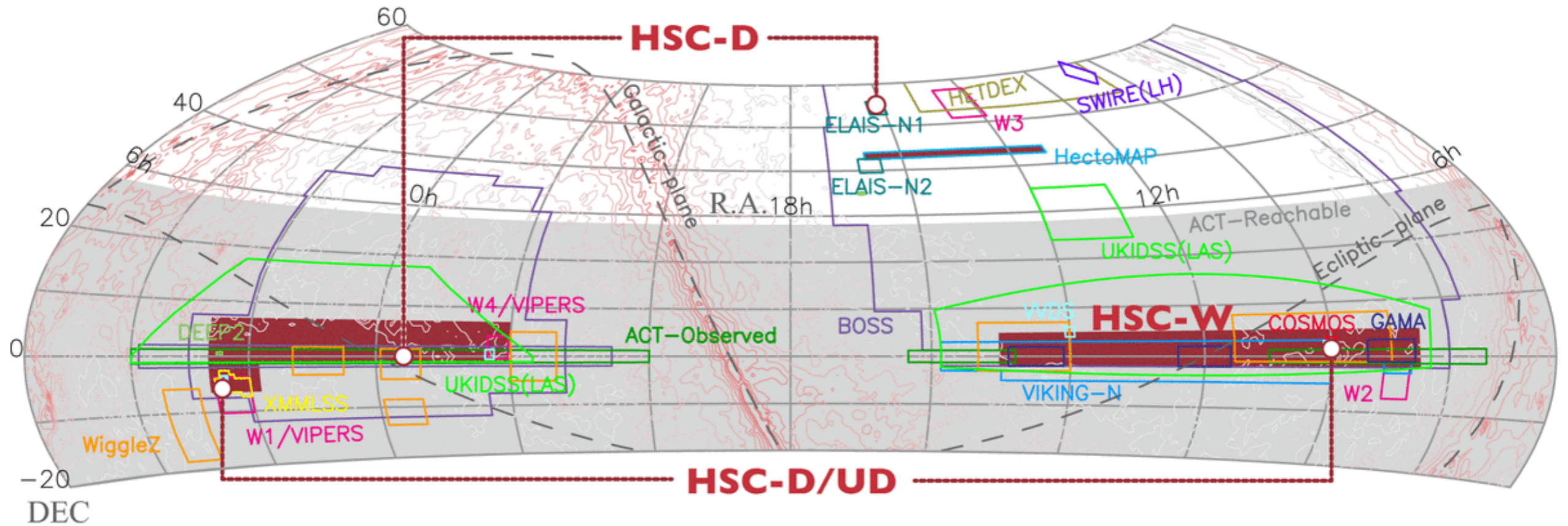
M31 8x magnified



M31 32x magnified



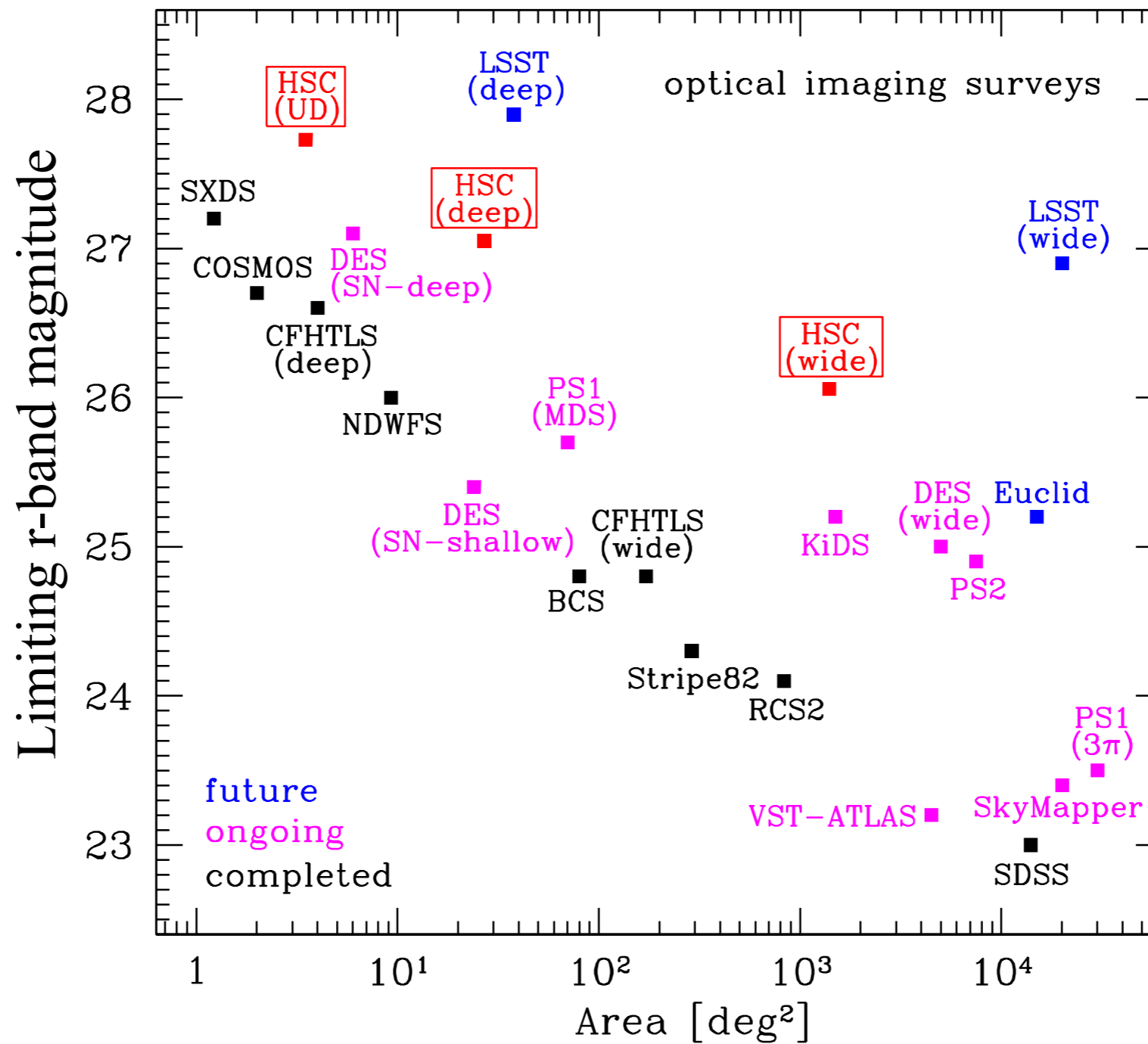
Hyper Suprime-Cam Survey



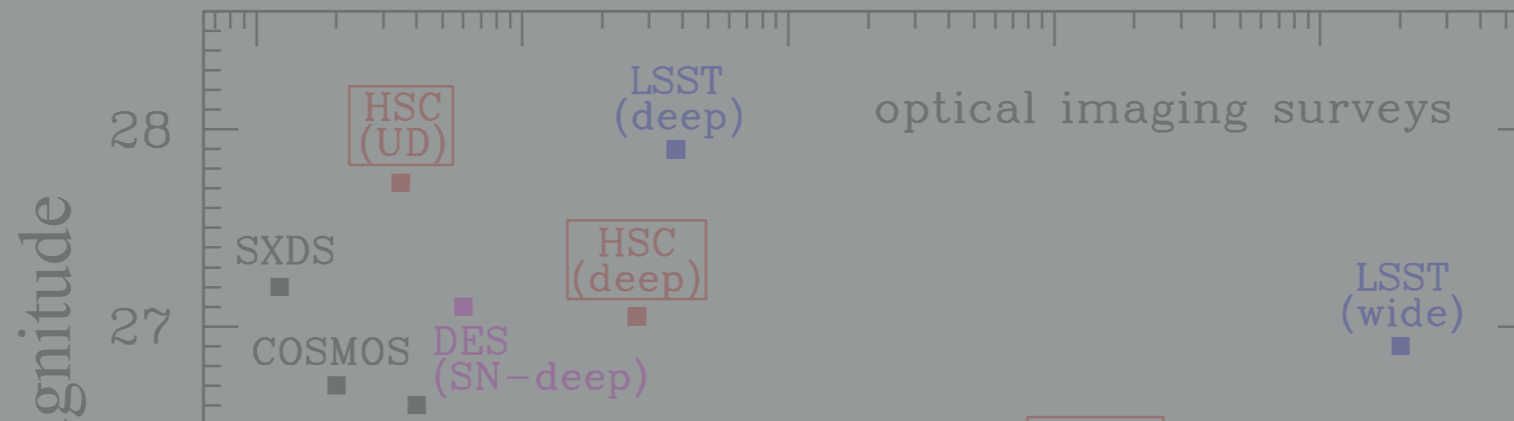
5 years, 300 Nights

	Wide	Deep	Ultra-Deep
Area	1400 deg ²	27 deg ²	3.5 deg ²
Filters	<i>grizy</i>	<i>grizy</i> +2NBs	<i>grizy</i> +2NBs
Depth (<i>i</i> -band)	25.9	26.8	27.4

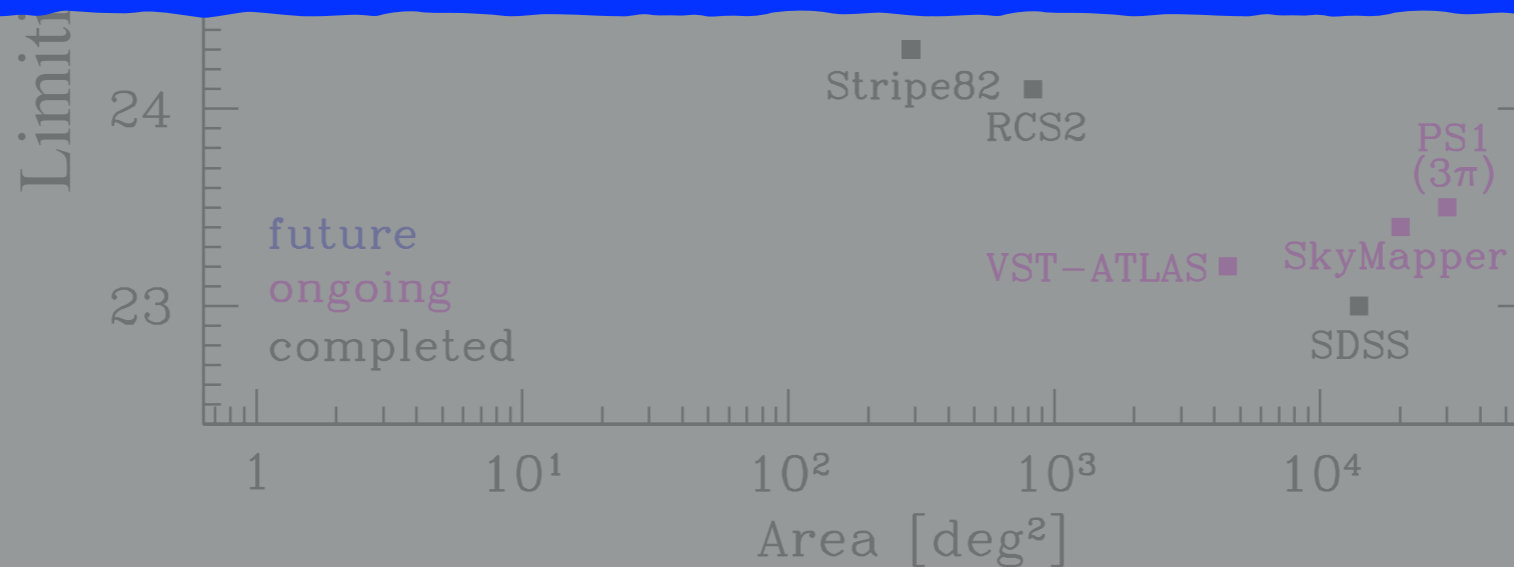
Hyper Suprime-Cam Survey



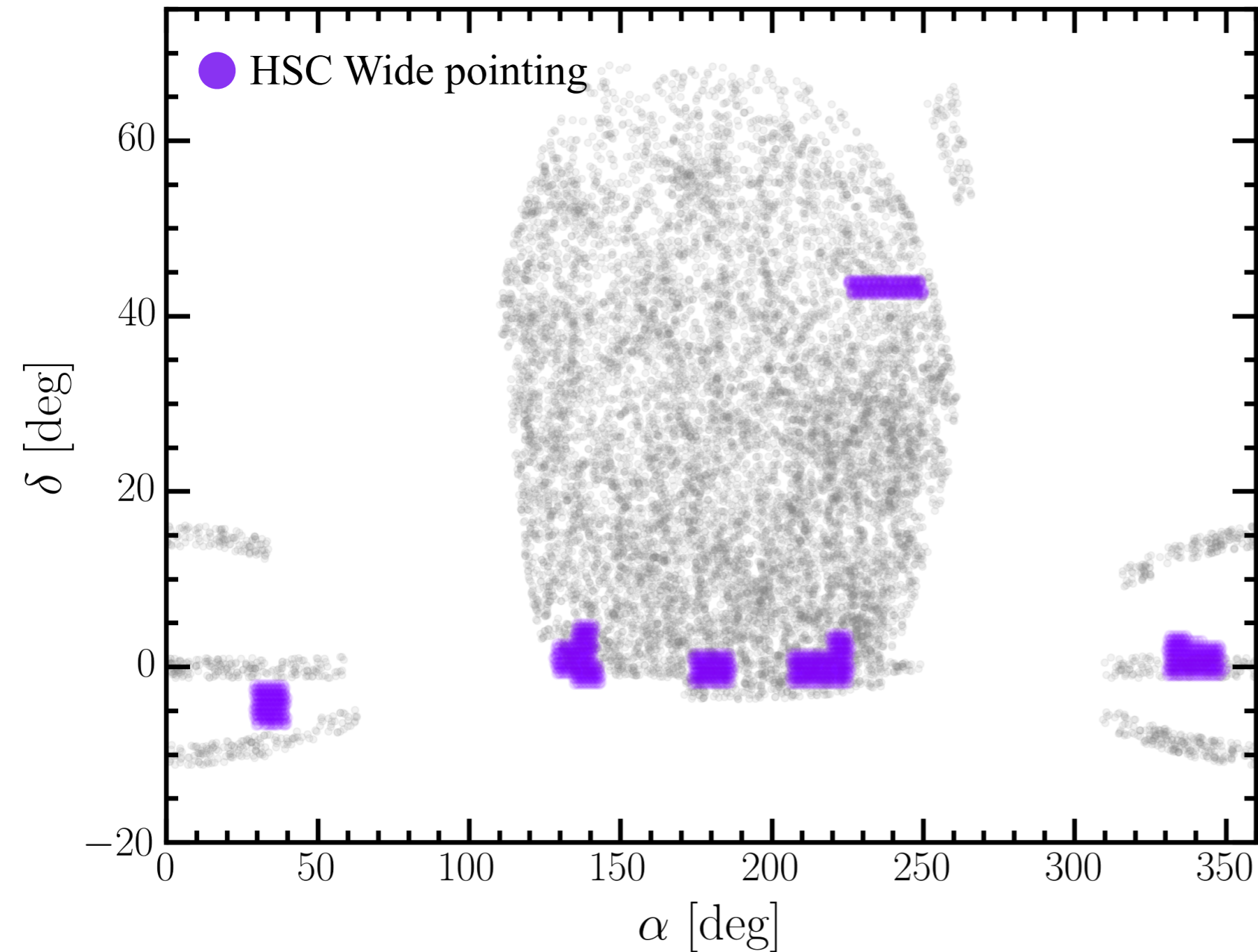
Hyper Suprime-Cam **Survey**



Special issue of PASJ with ~40 HSC papers
Public DR1 now available (100 deg²) !



Our search for LSBGs

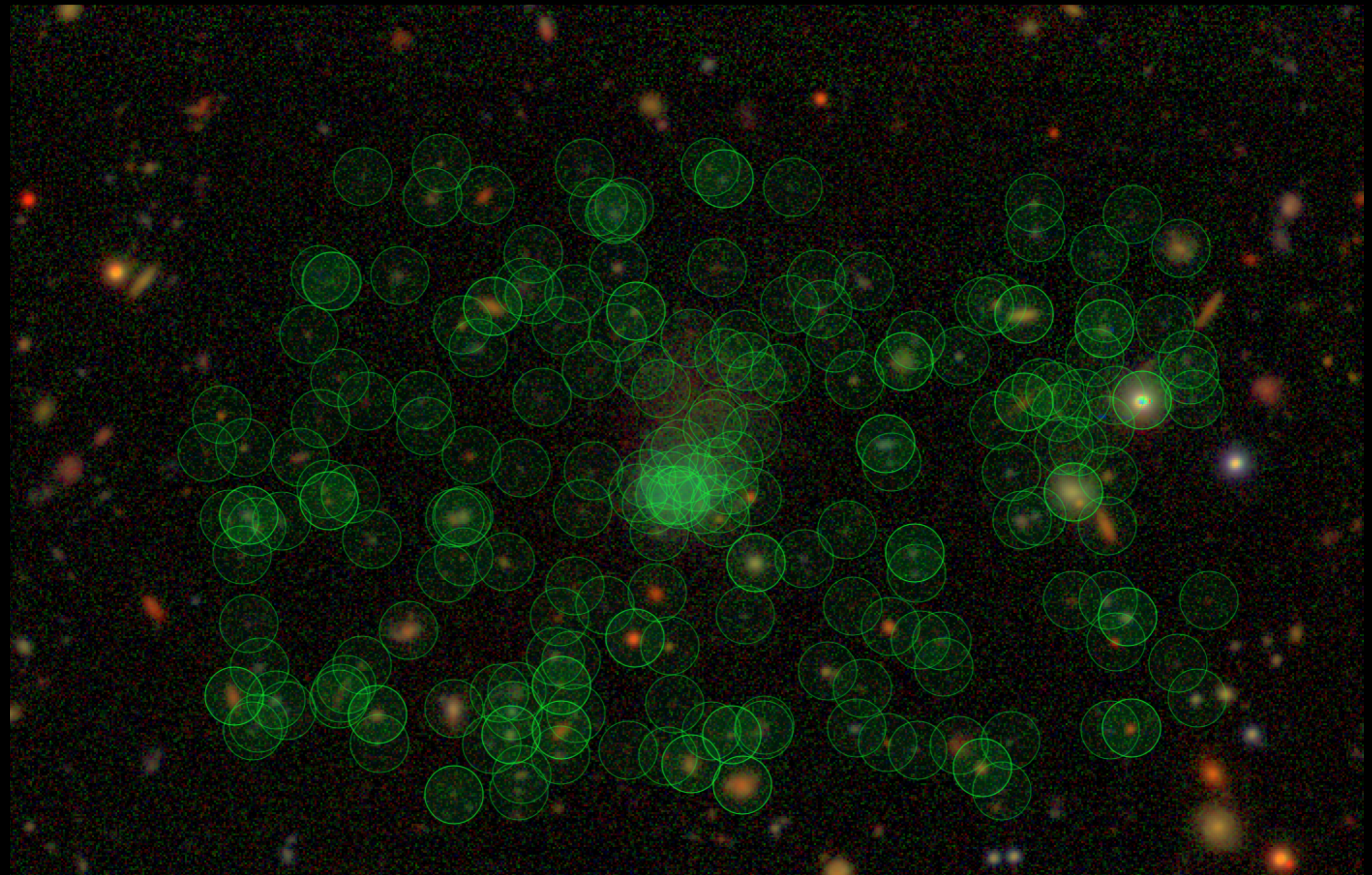


- ◆ Carry out search in HSC **Wide** layer
- ◆ **~200 deg²** with full Wider layer depth in *gri*

Search using **HSC** catalog?



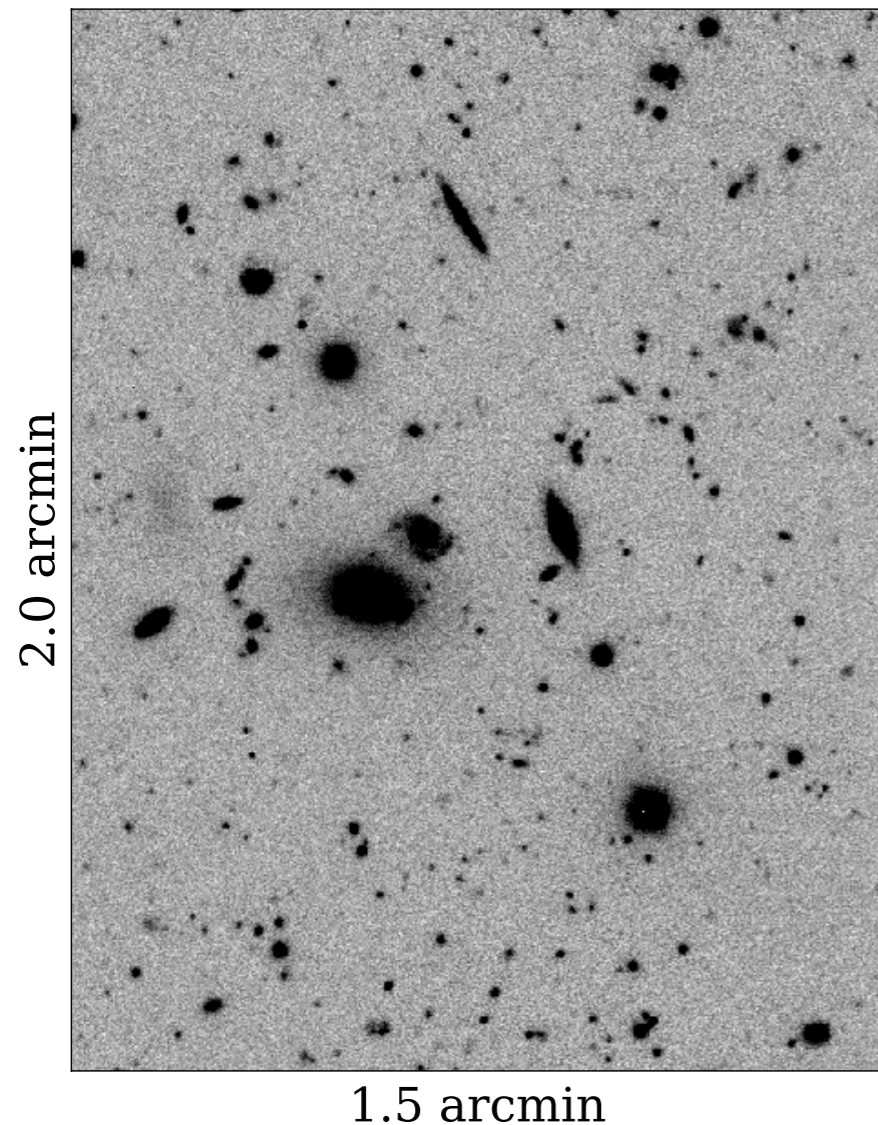
Galaxy Shredding



LSBG Detection Pipeline

LSBG Detection Pipeline: Source Extraction

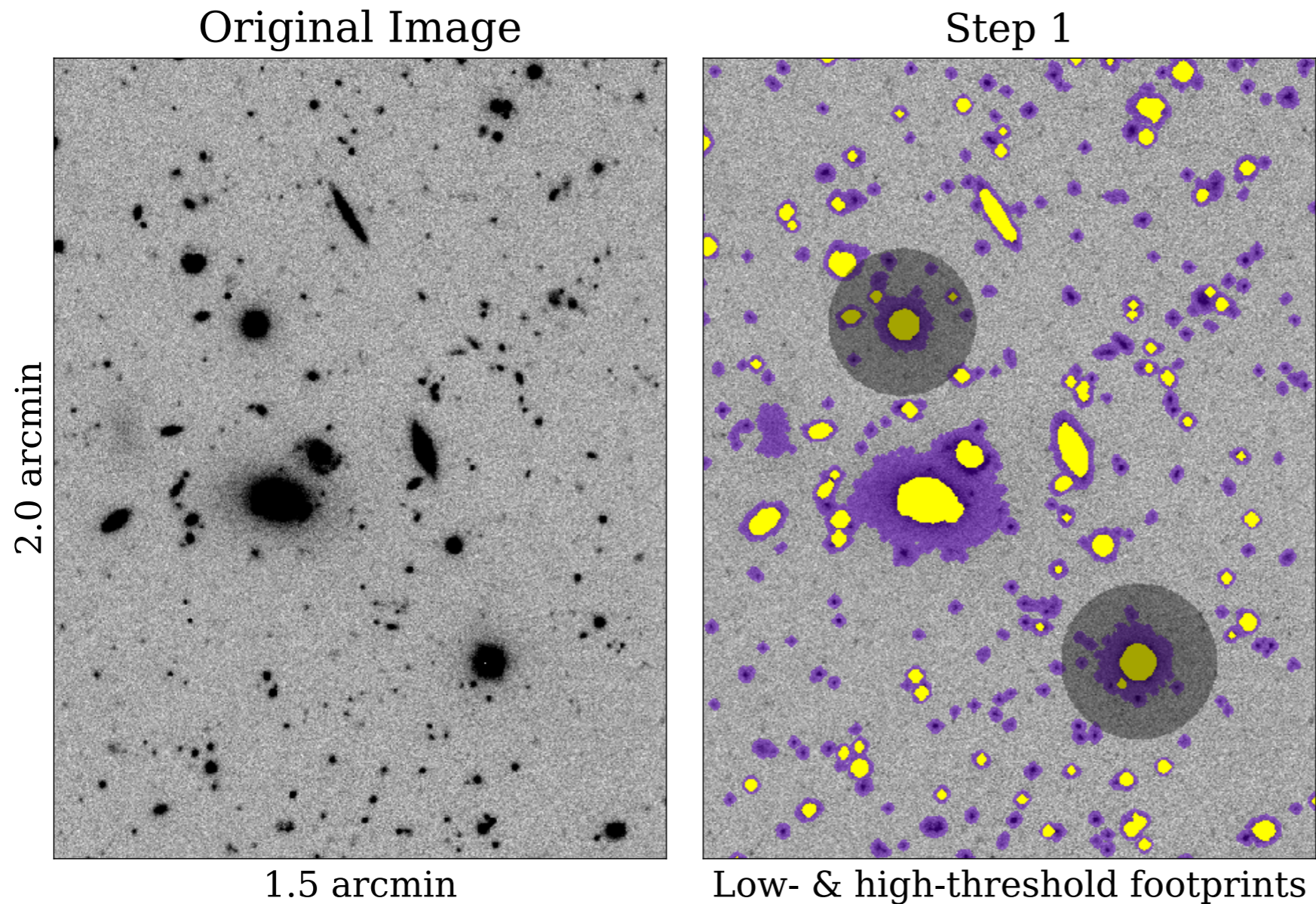
Original Image



Source extraction in two steps:

- Initial image processing with LSST codebase: <http://dm.lsst.org>
- Final source detection with *SExtractor*

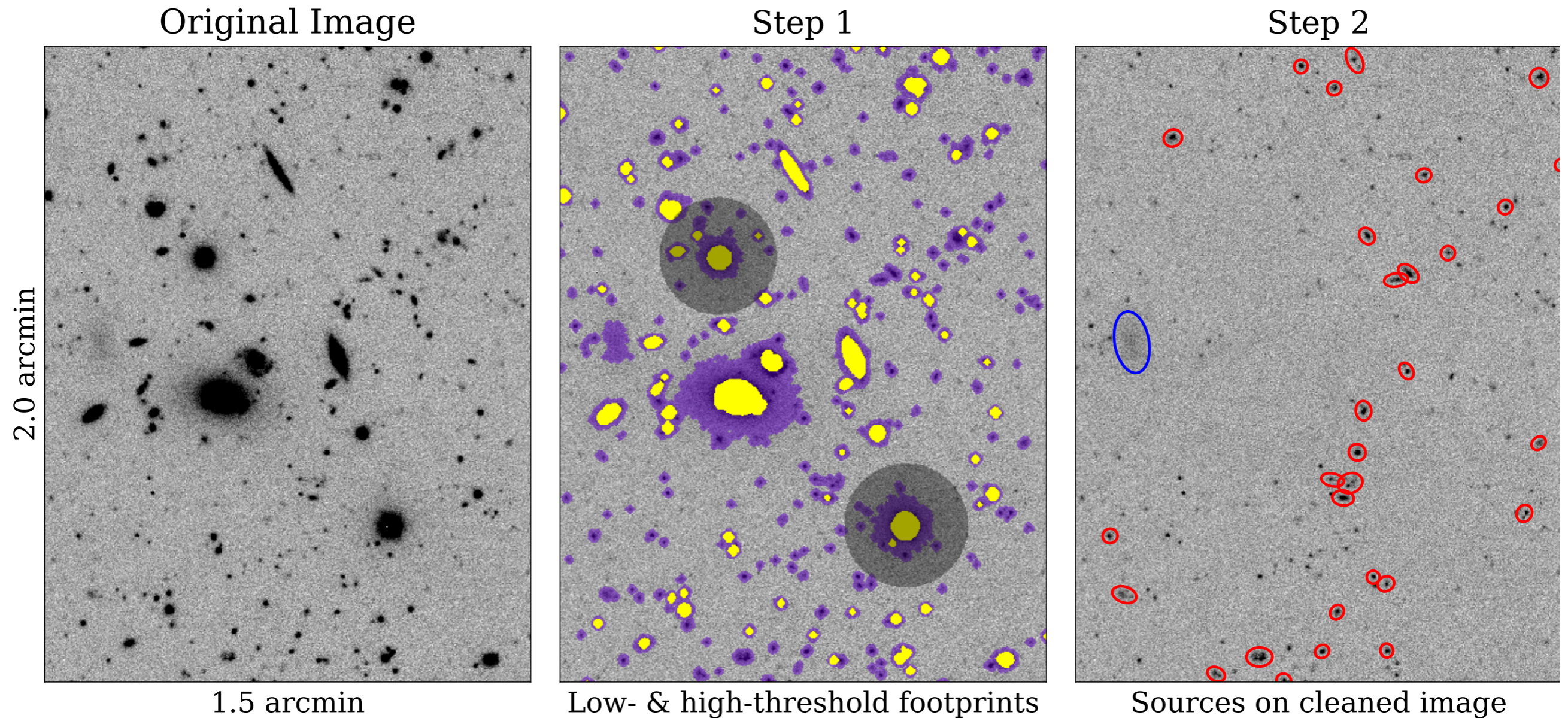
LSBG Detection Pipeline: Source Extraction



Source extraction in two steps:

- Initial image processing with LSST codebase: <http://dm.lsst.org>
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LSBG Detection Pipeline: Source Extraction



Source extraction in two steps:

- Initial image processing with LSST codebase: <http://dm.lsst.org>
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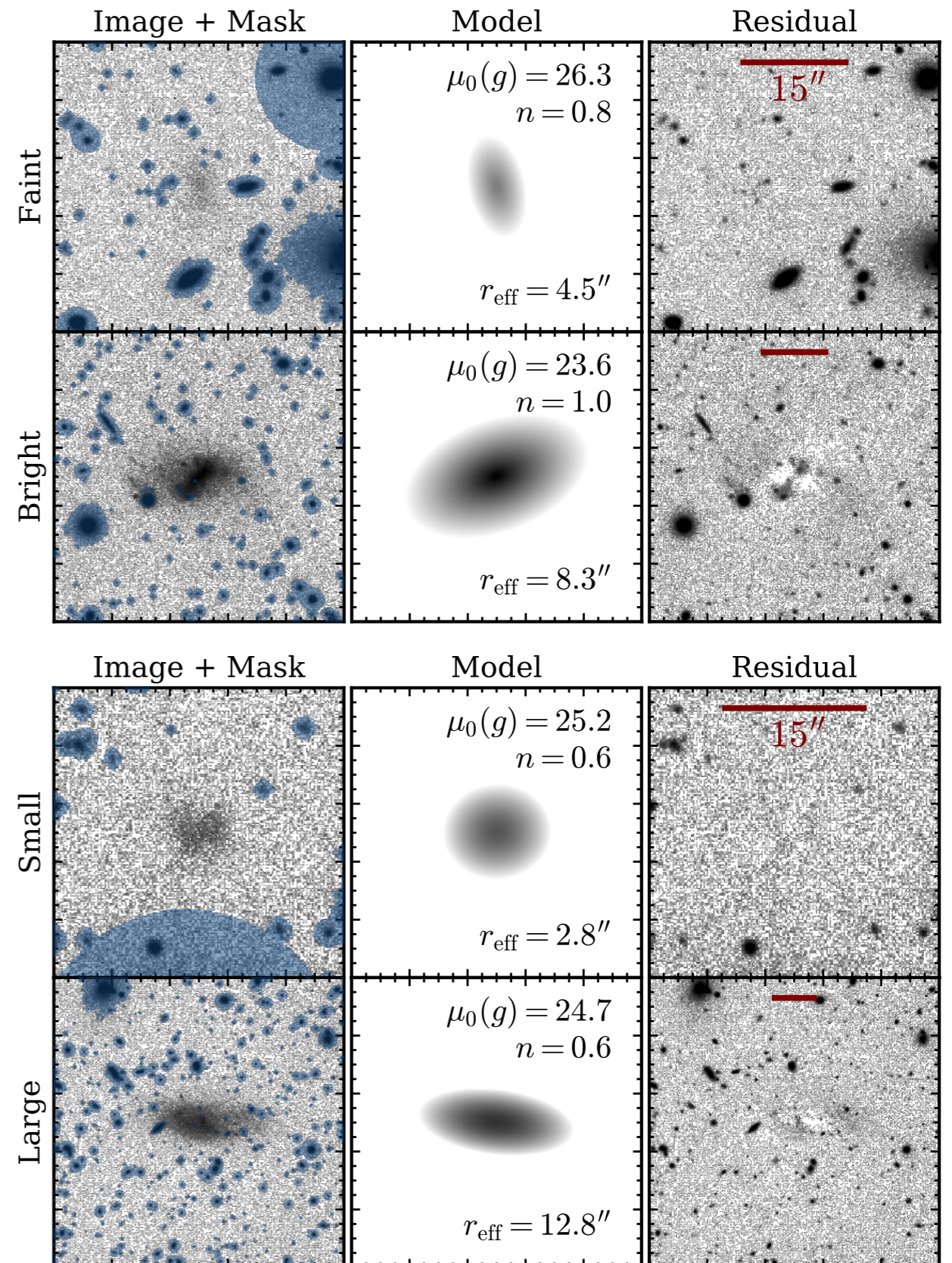
LSBG Detection Pipeline: Galaxy Modeling

Our LSBG definition:

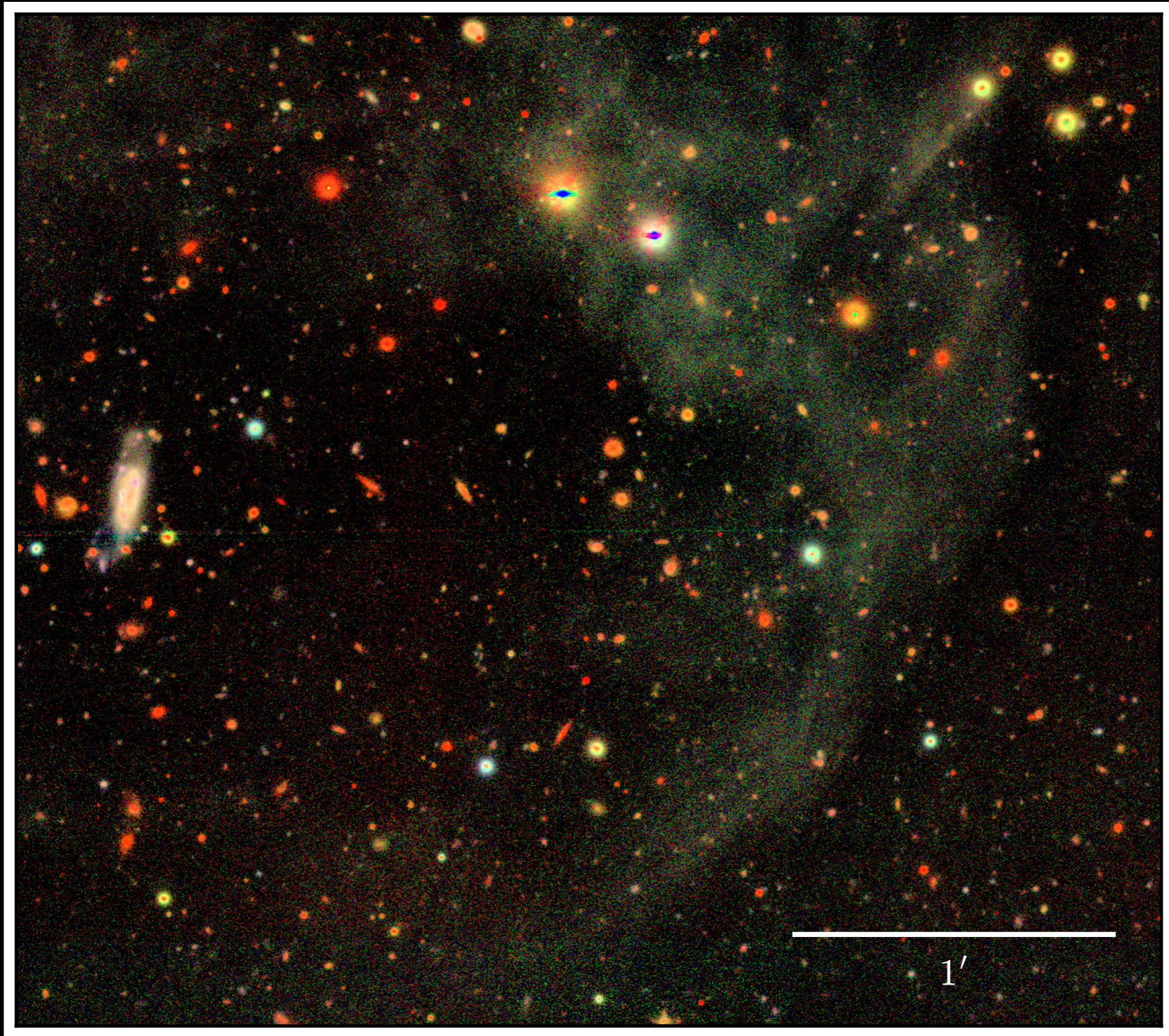
$$\bar{\mu}_{\text{eff}}(g) > 24.3 \text{ mag arcsec}^{-2}$$

$$r_{\text{eff}} > 2.5''$$

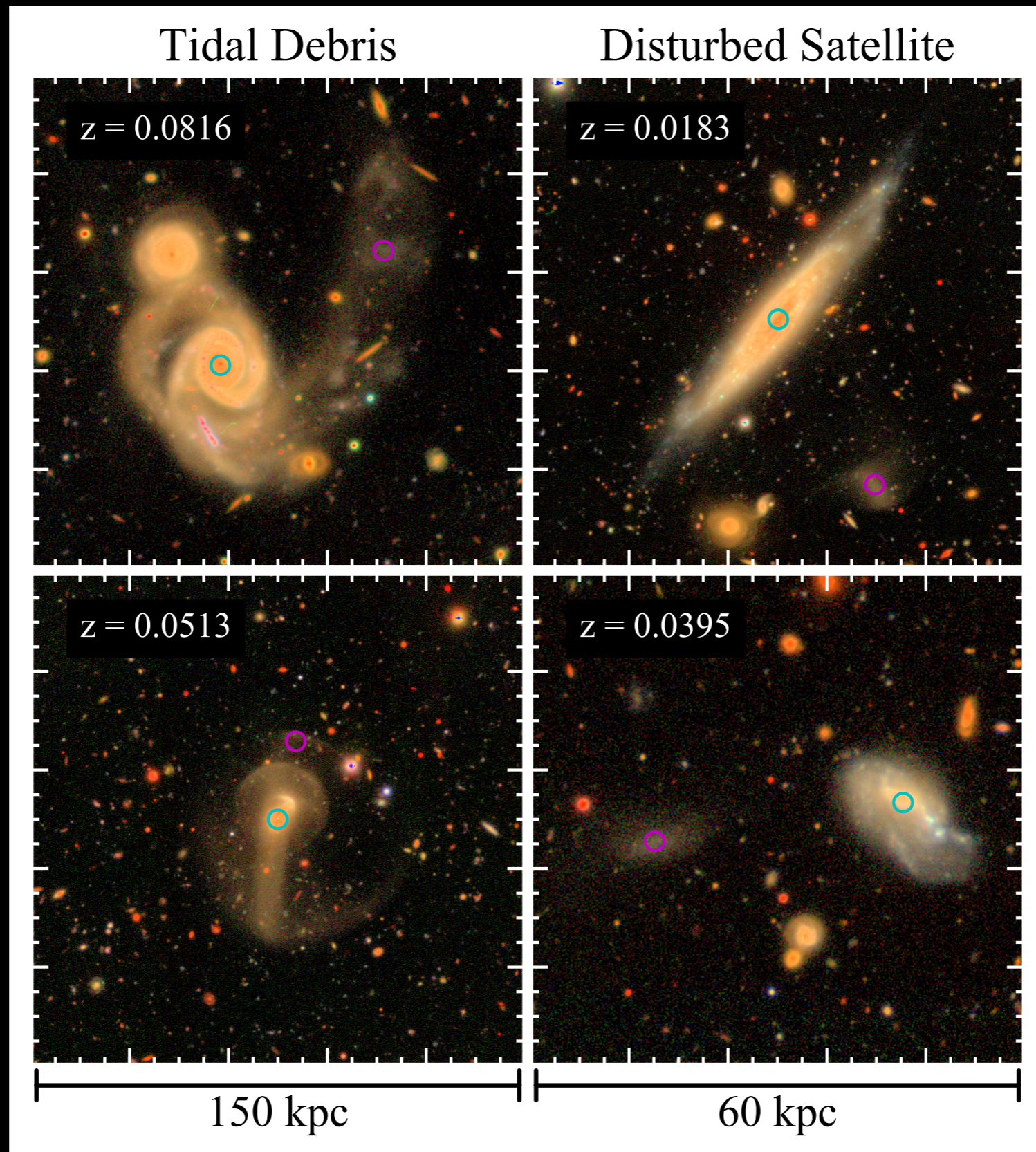
- Model LSBG candidates as single-component Sersic functions
- Make selection on best-fit parameters
- Visually inspect remaining candidates
- Final sample size: **781 LSBGs**



LSB False Positives: Galactic Cirrus



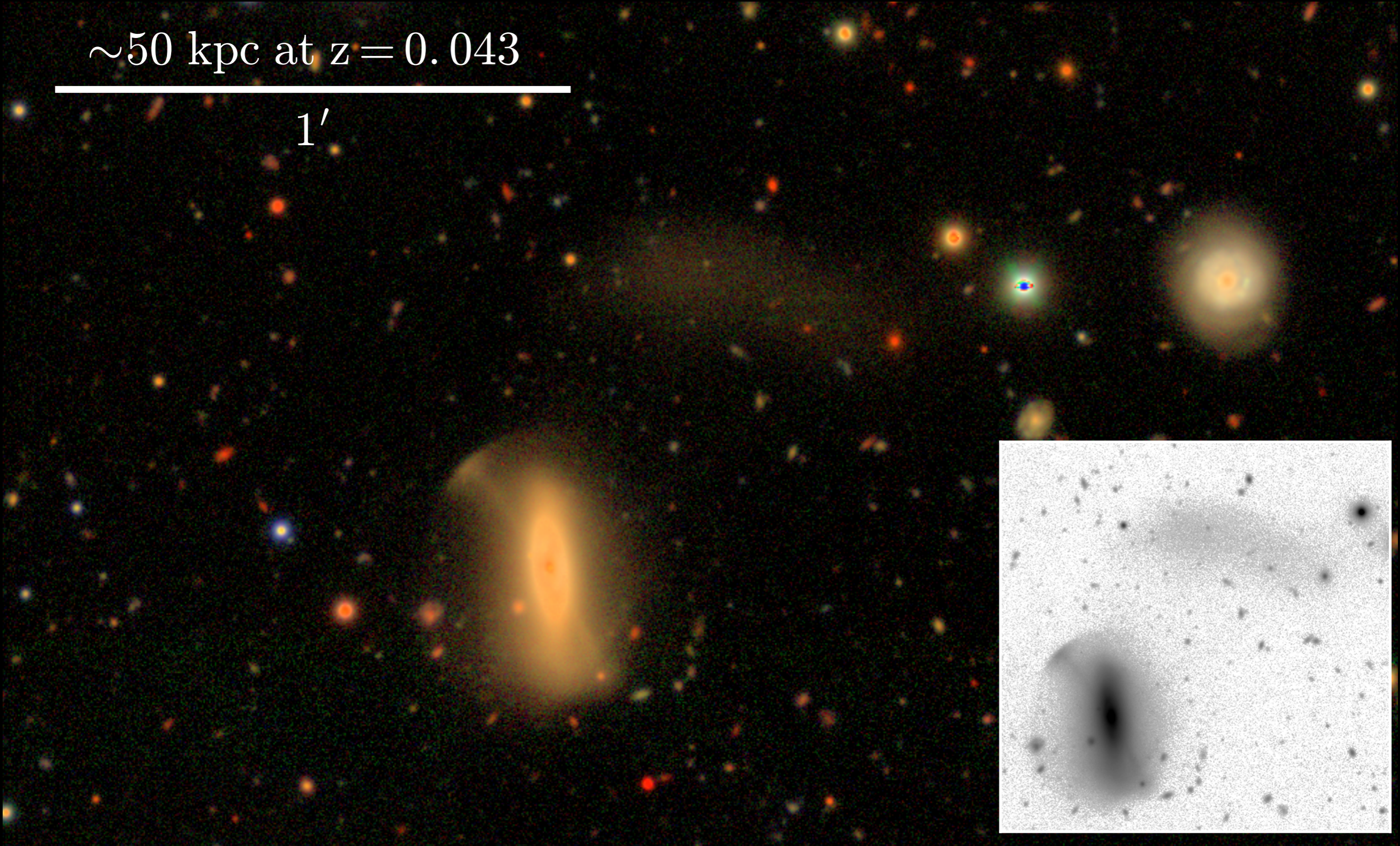
LSB False Positives: **Tidal** Debris?



LSB False Positives: **Tidal** Debris?

~ 50 kpc at $z = 0.043$

1'

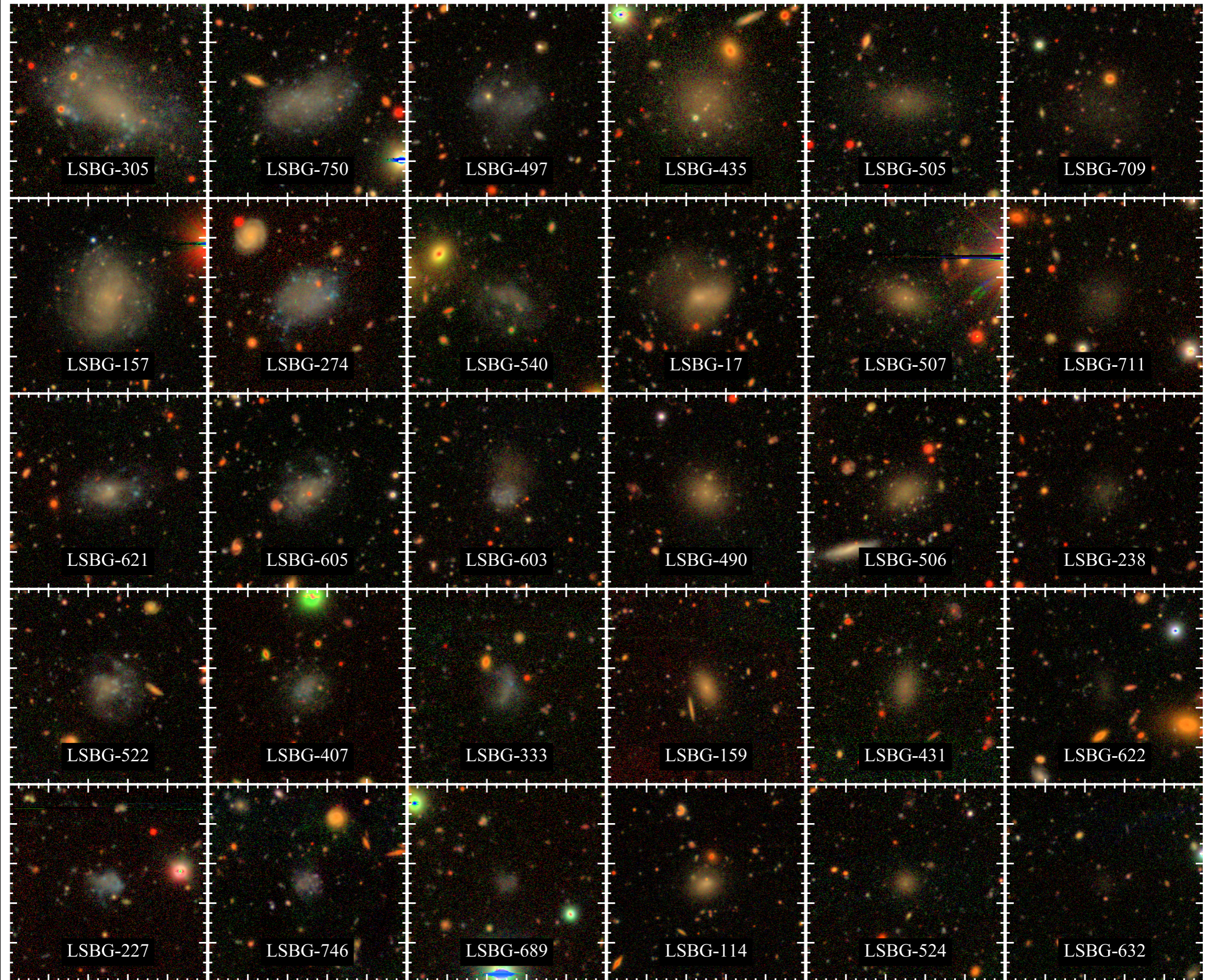


LSBG Sample

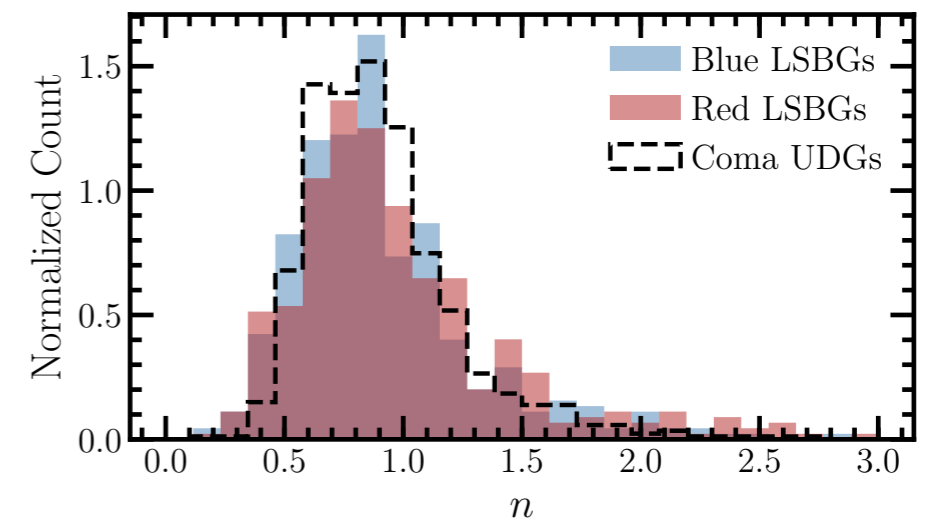
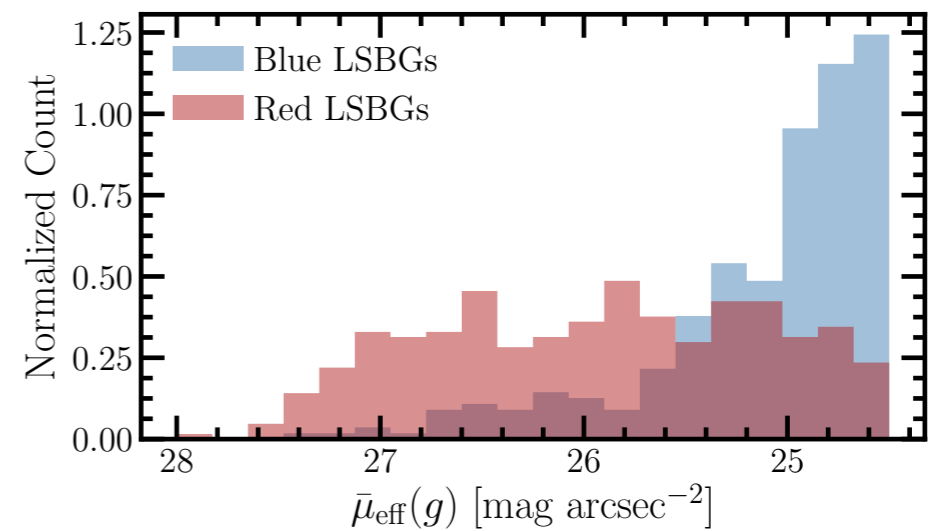
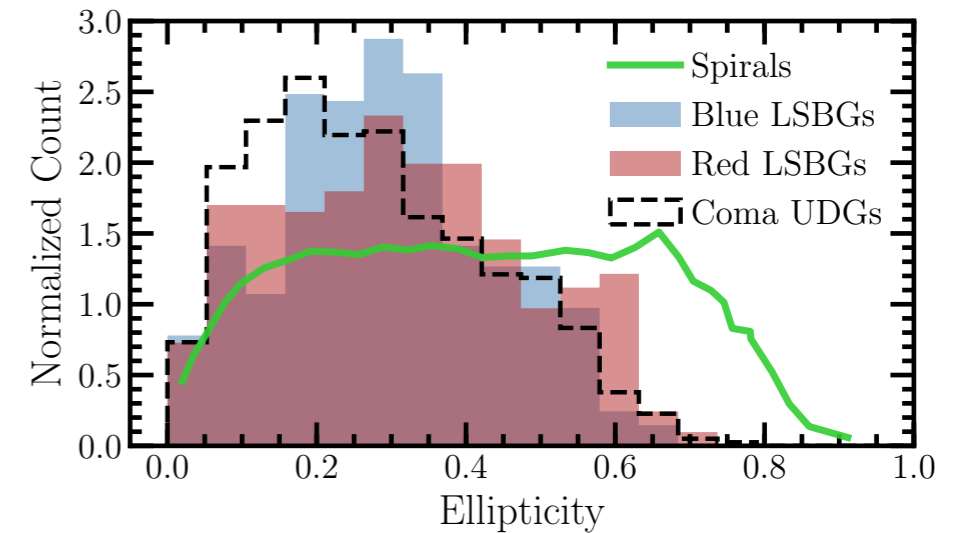
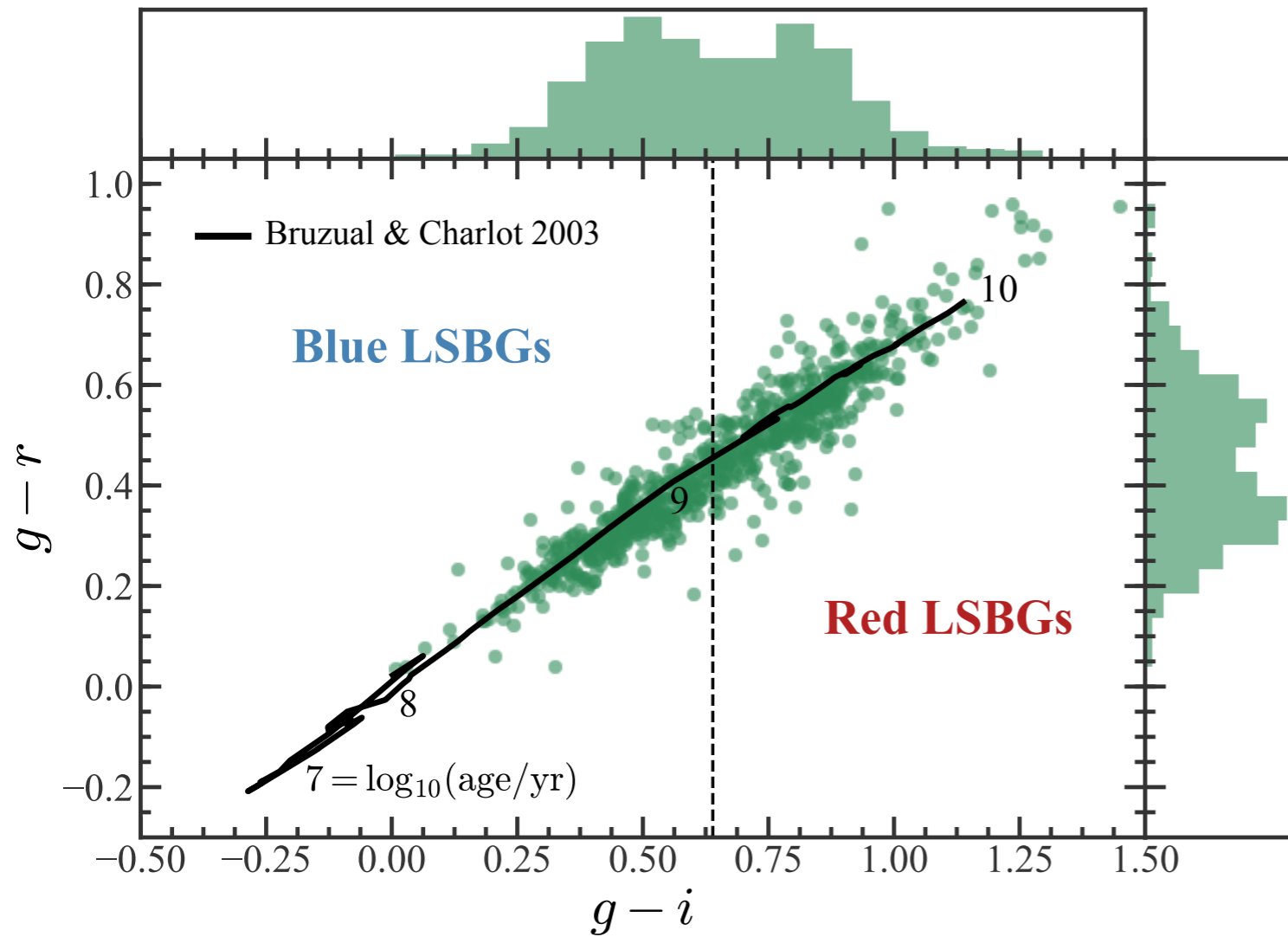
Blue LSBGs

Red LSBGs

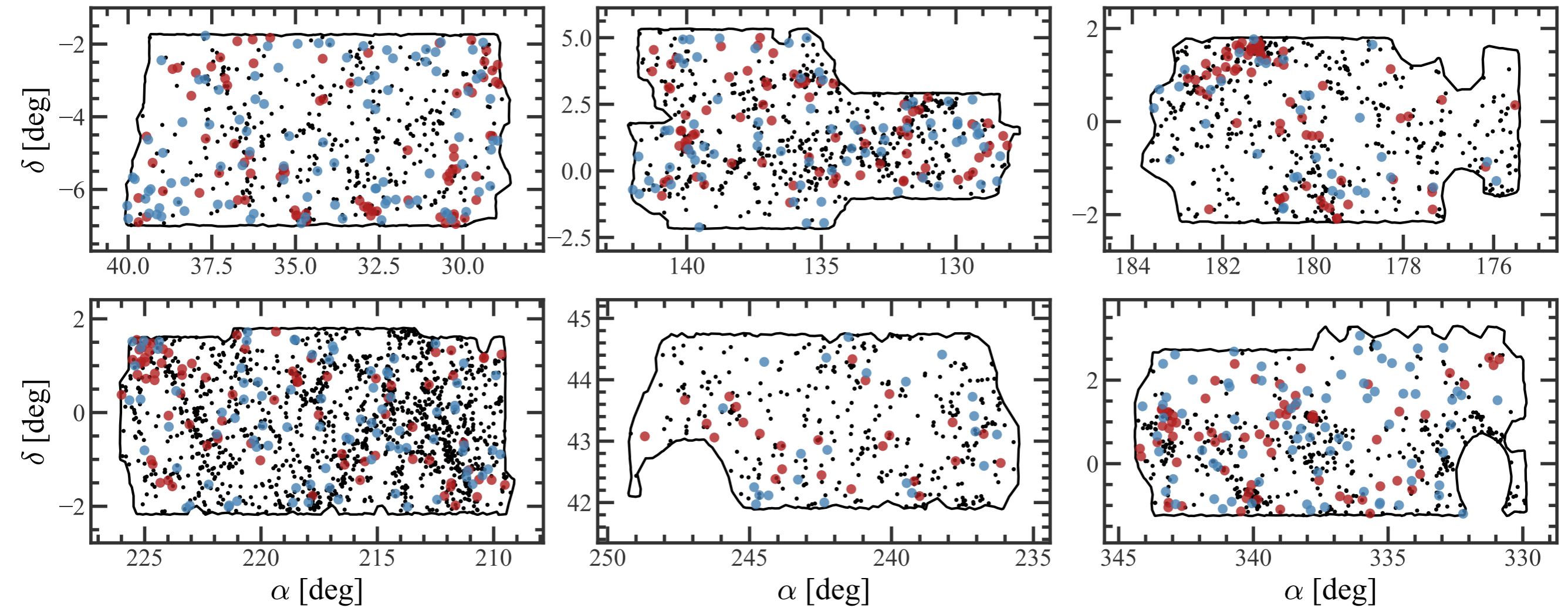
55''



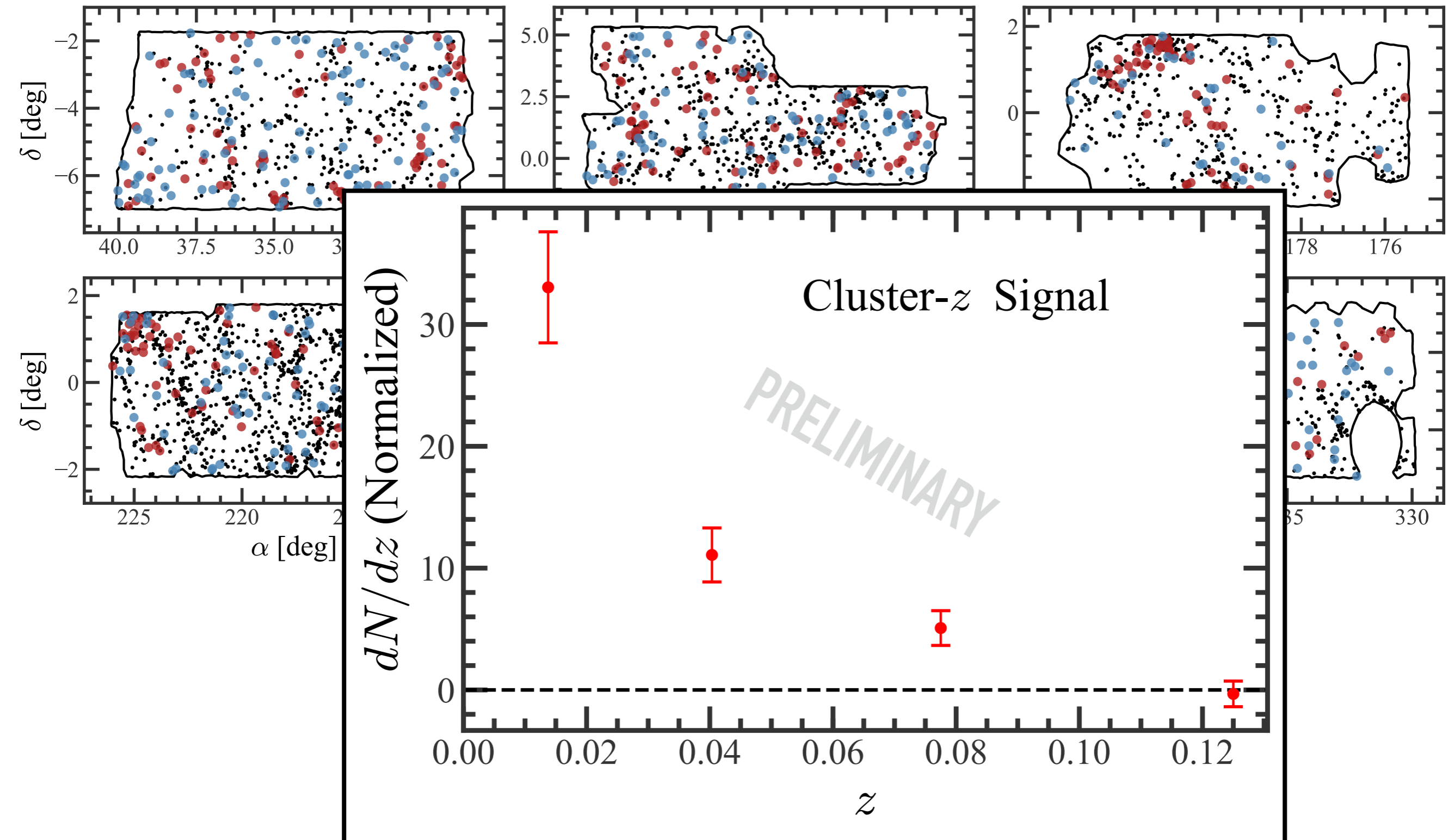
LSBG Sample: Parameter Distributions



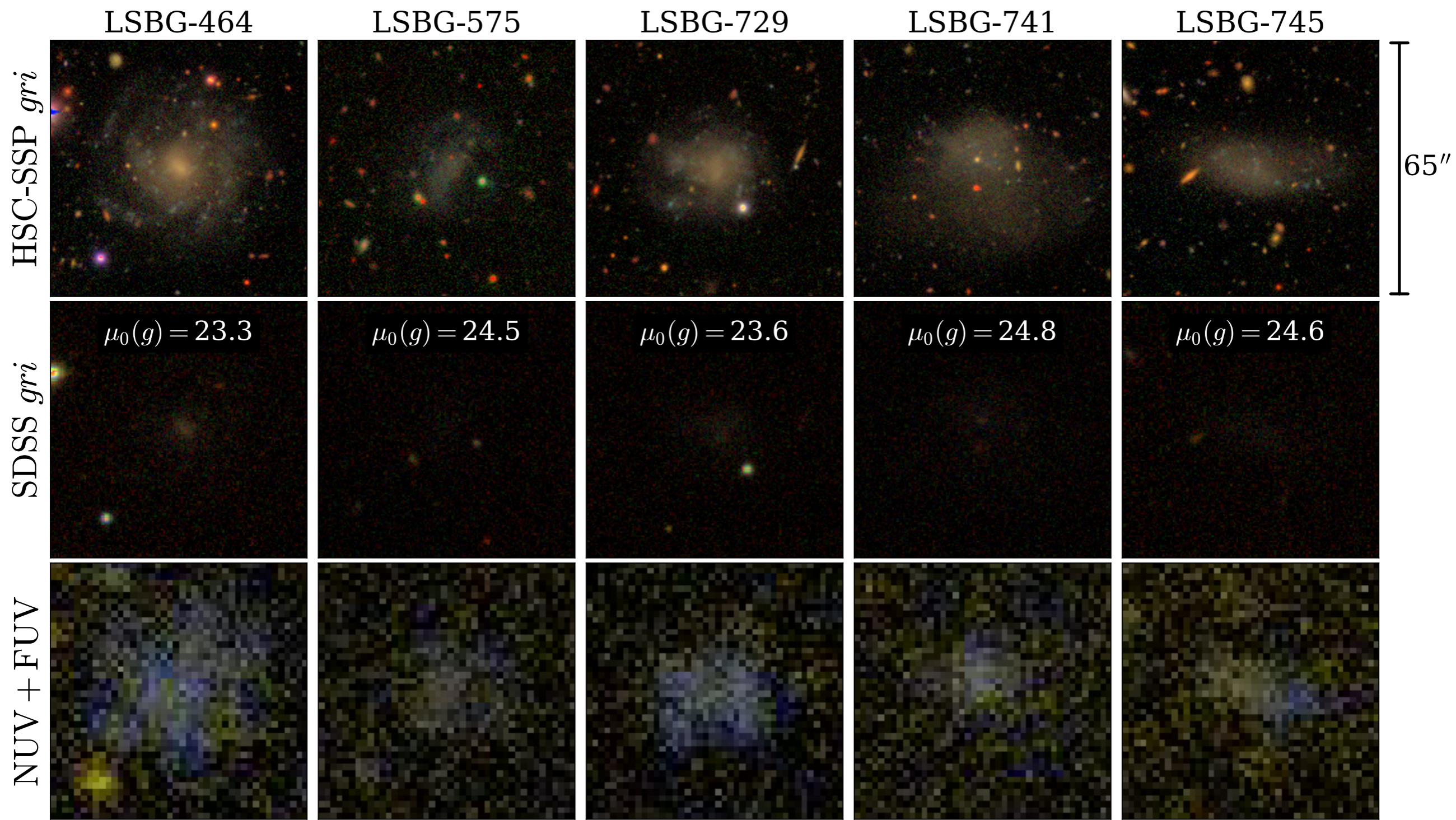
LSBG Sample: Spatial Distribution



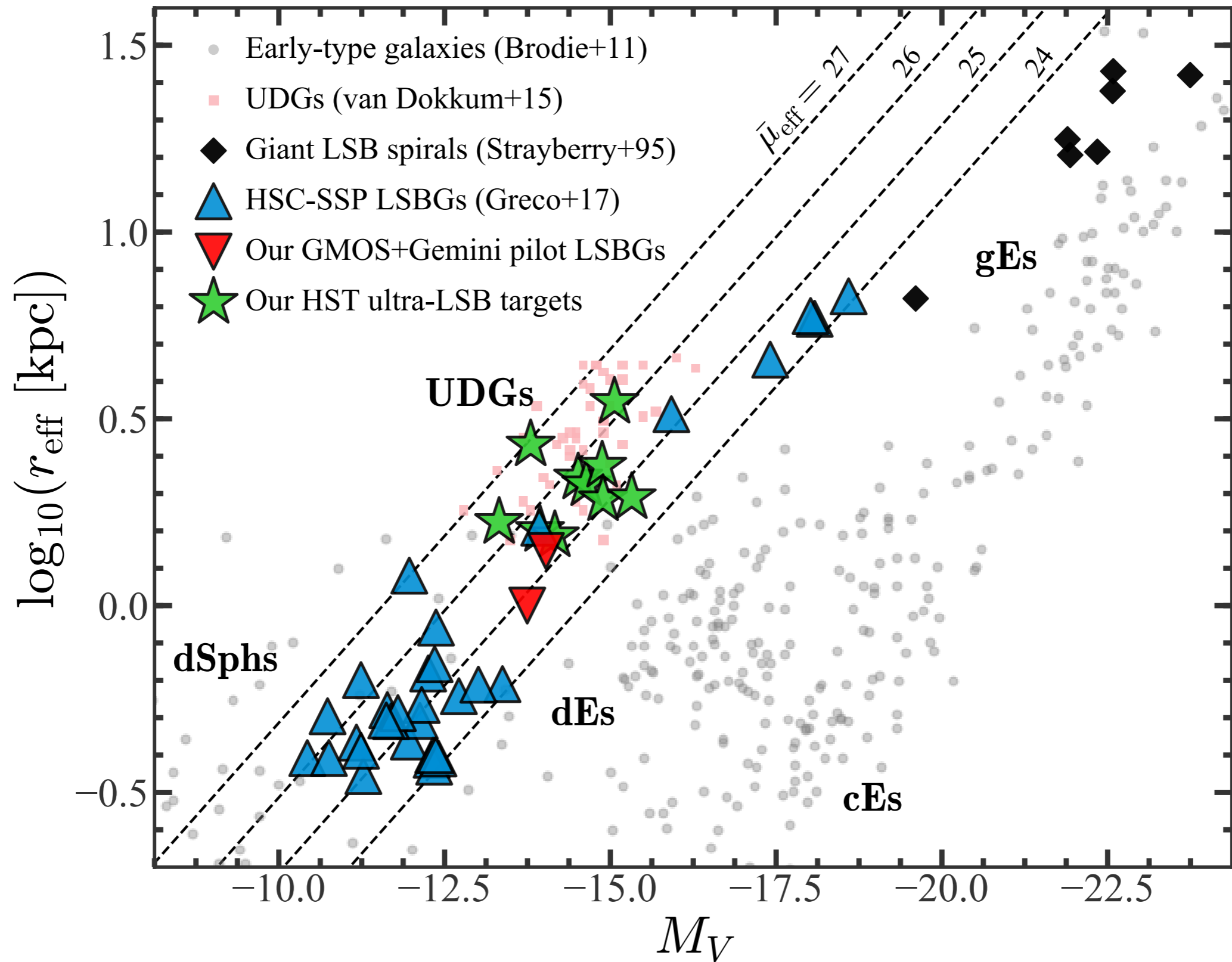
LSBG Sample: Spatial Distribution



LSBG Sample: Catalog Crossmatching

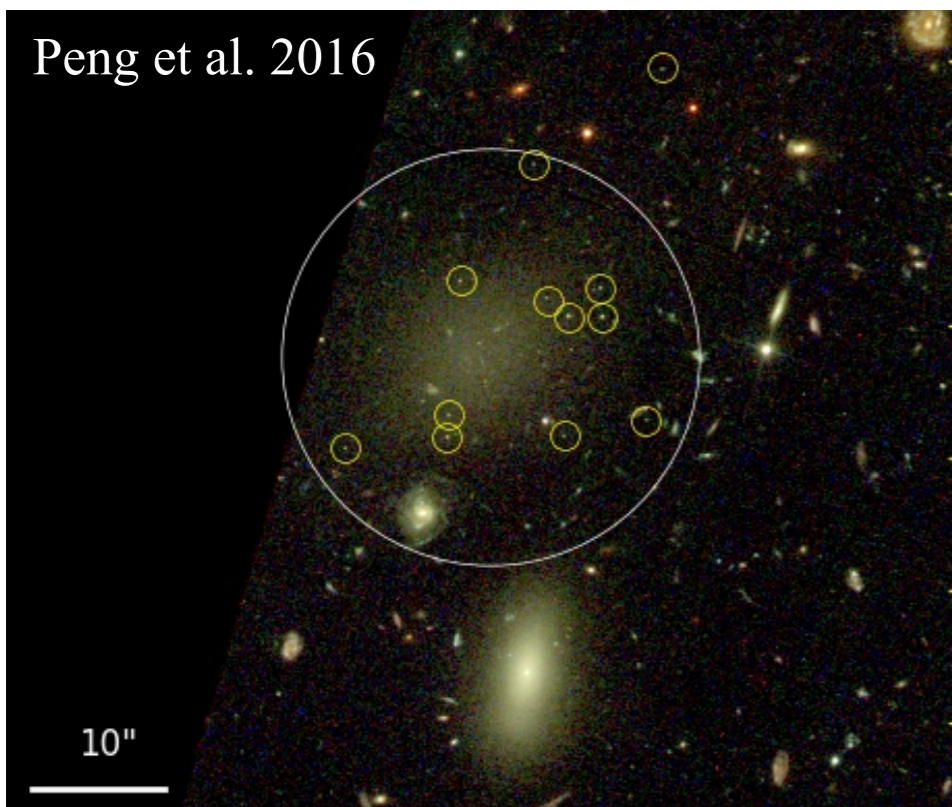
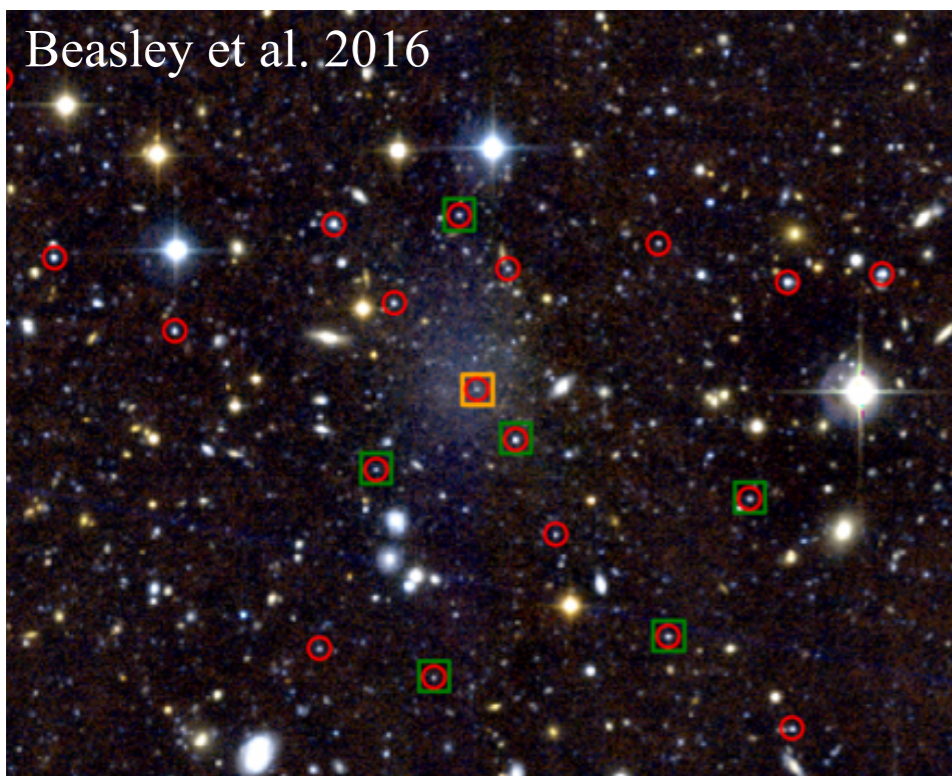


Size-Luminosity Relation

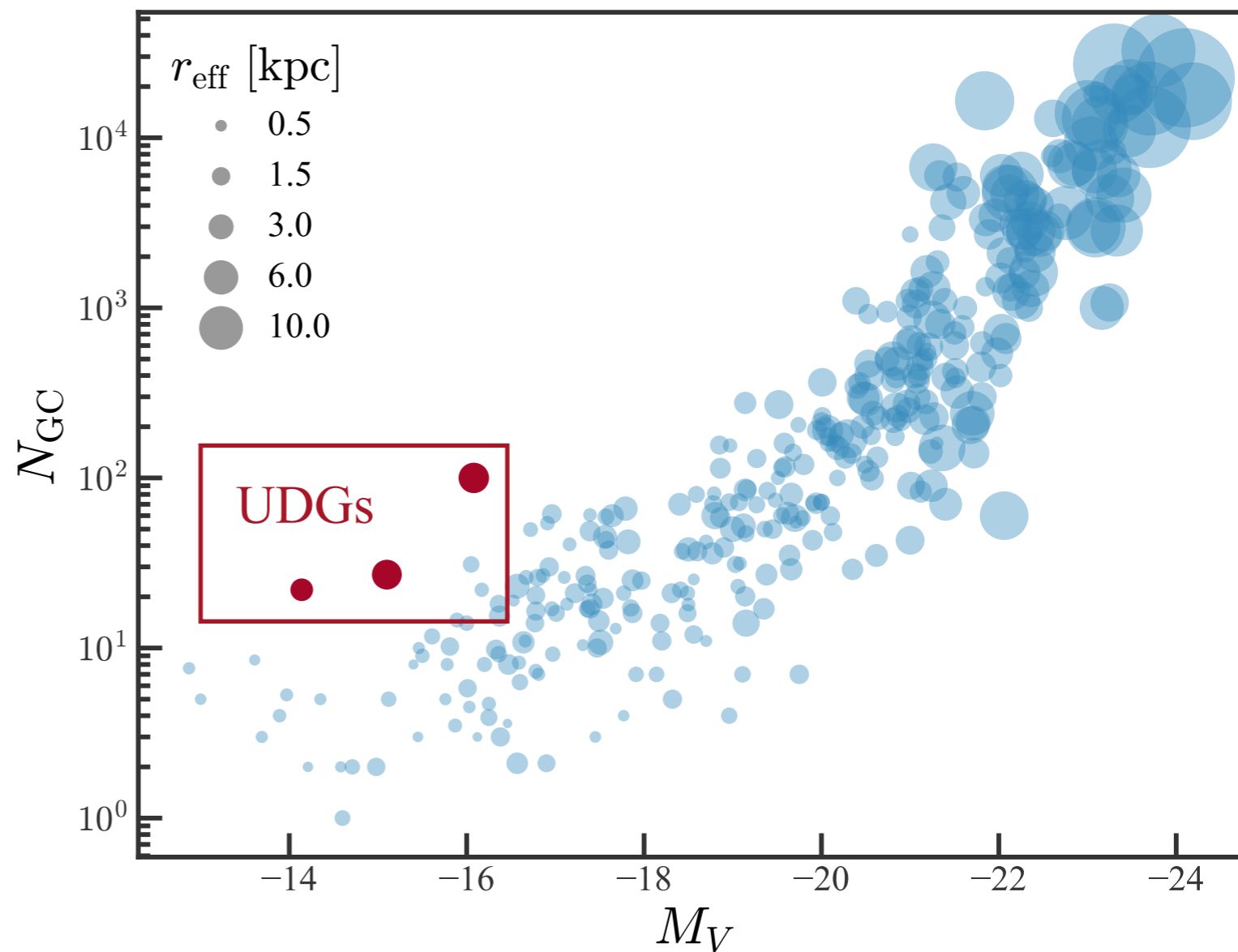


What's **Next**?

What's Next?

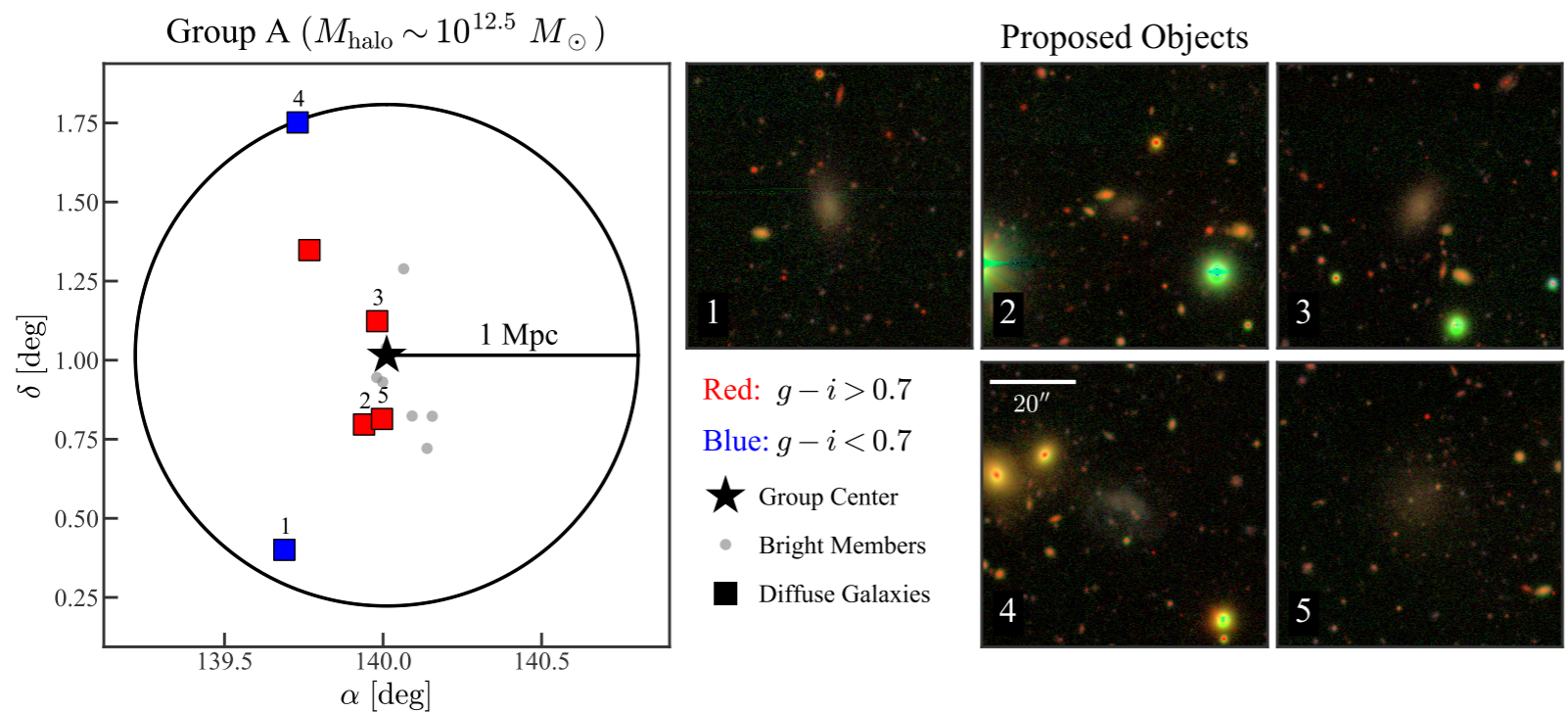
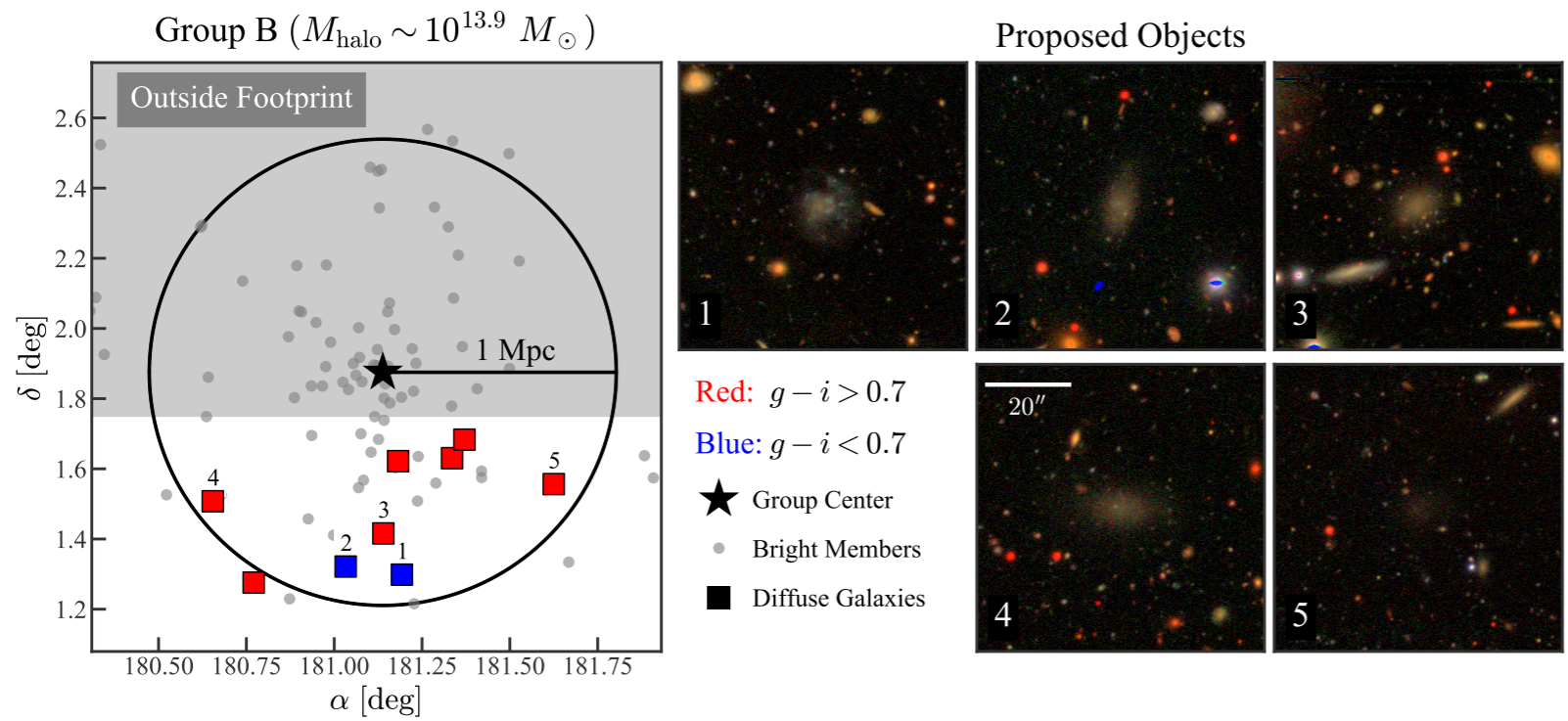


UDGs & rich globular cluster systems



What's Next?

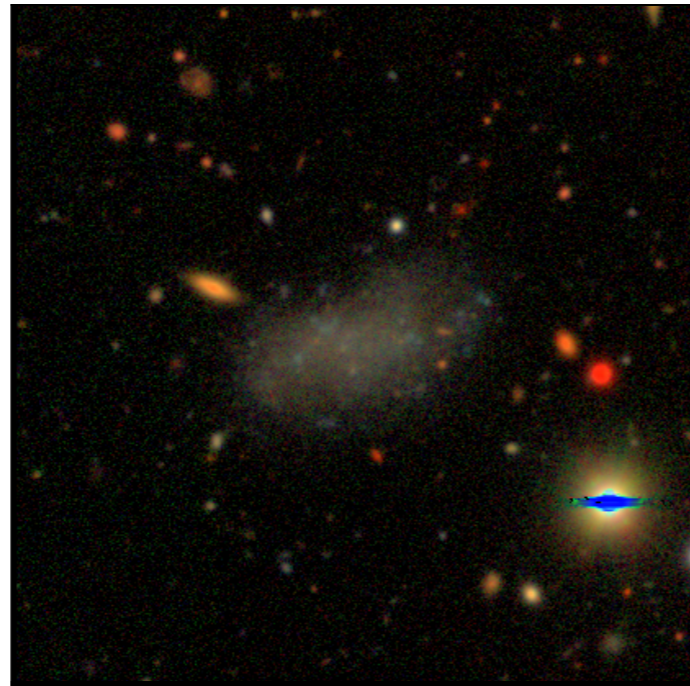
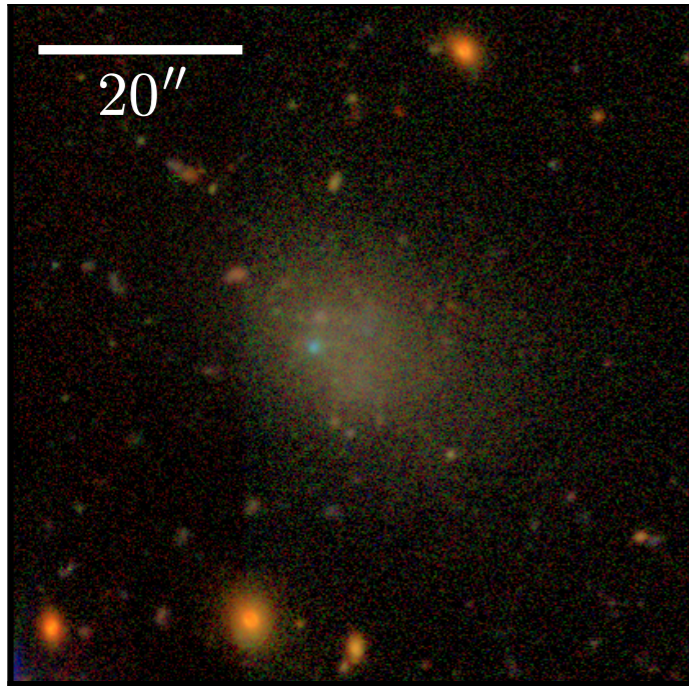
Counting globular clusters with HST



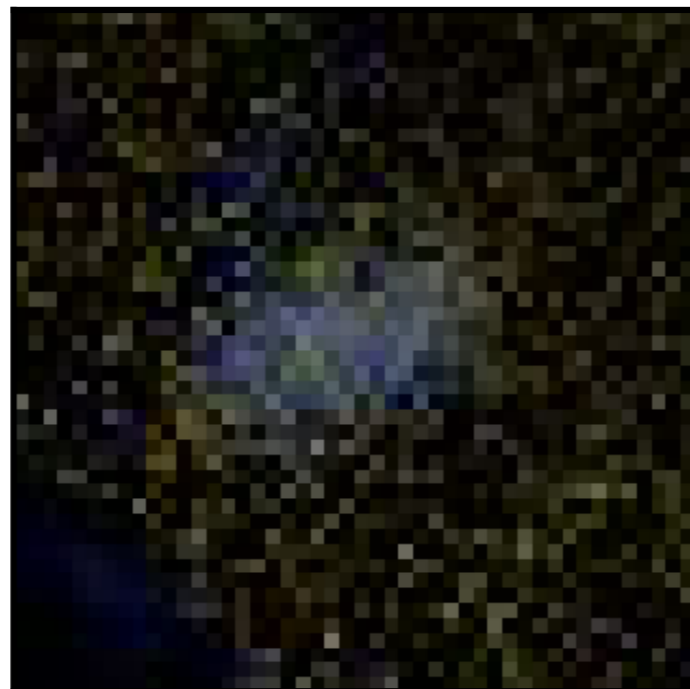
What's **Next?**

Redshifts with GMOS on Gemini

HSC-SPP *gri*

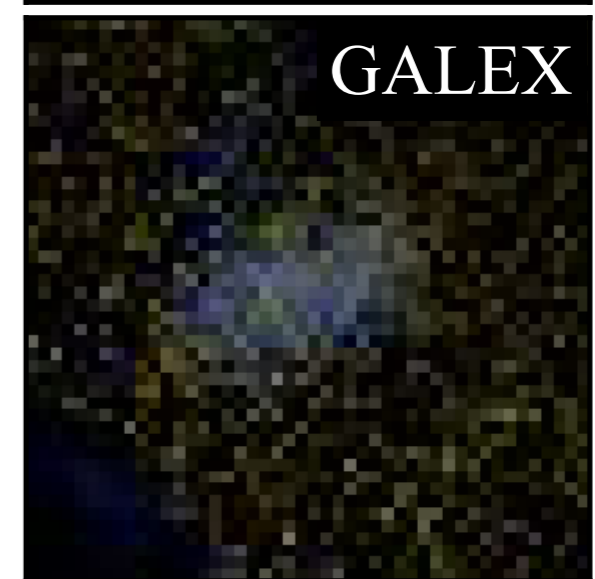
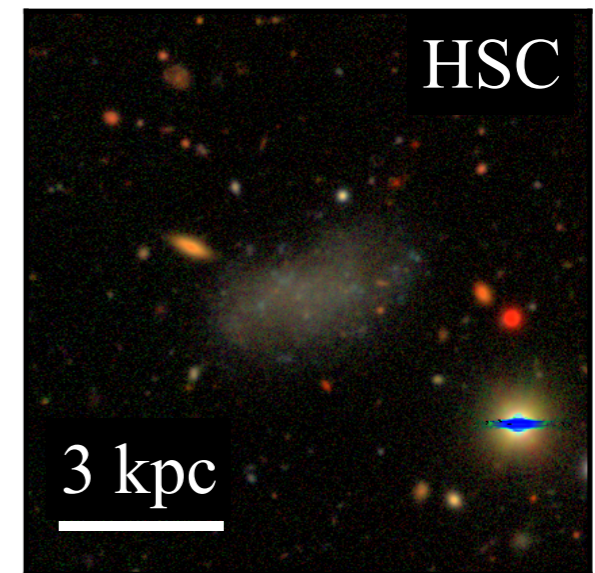
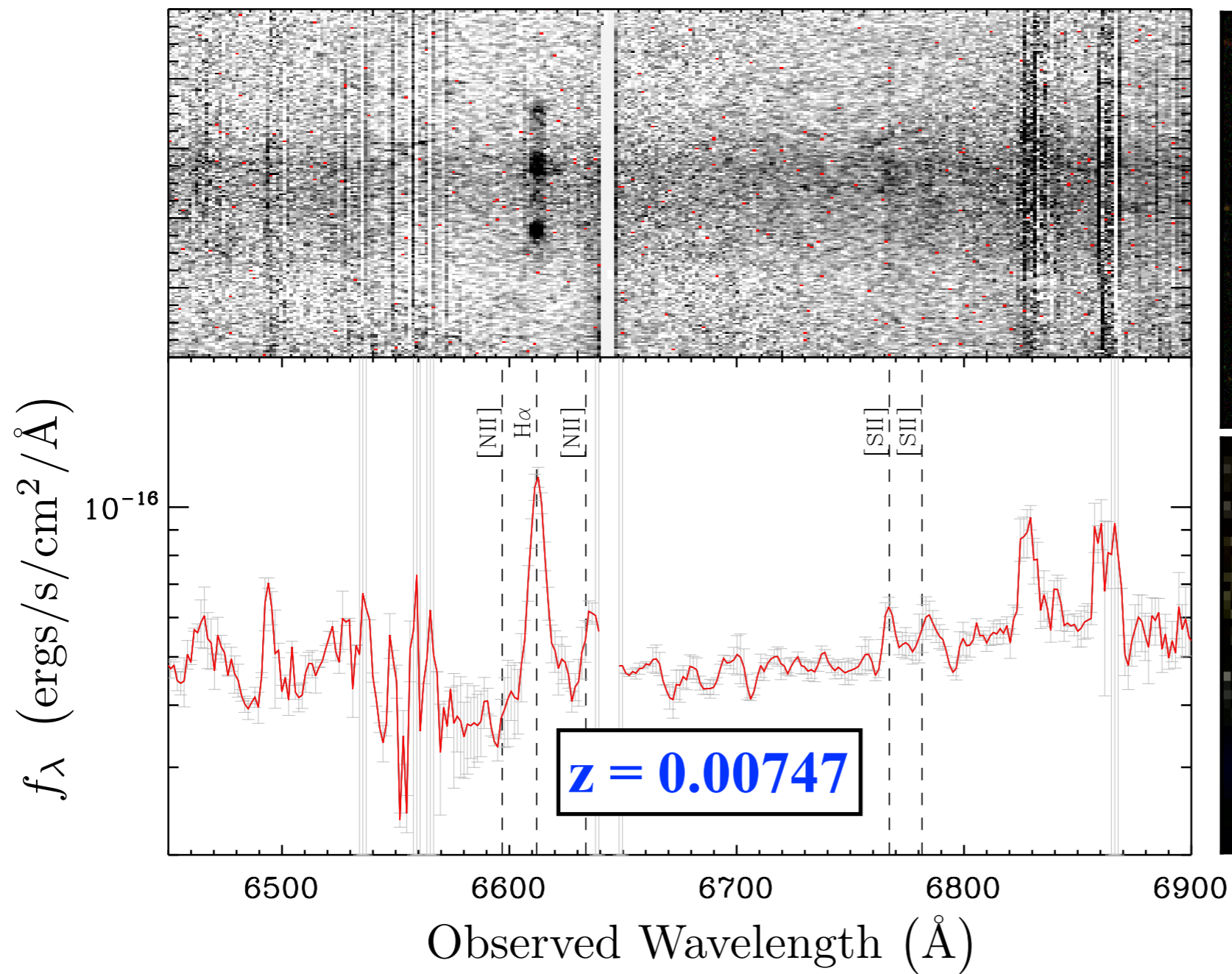


Galex NUV+FUUV

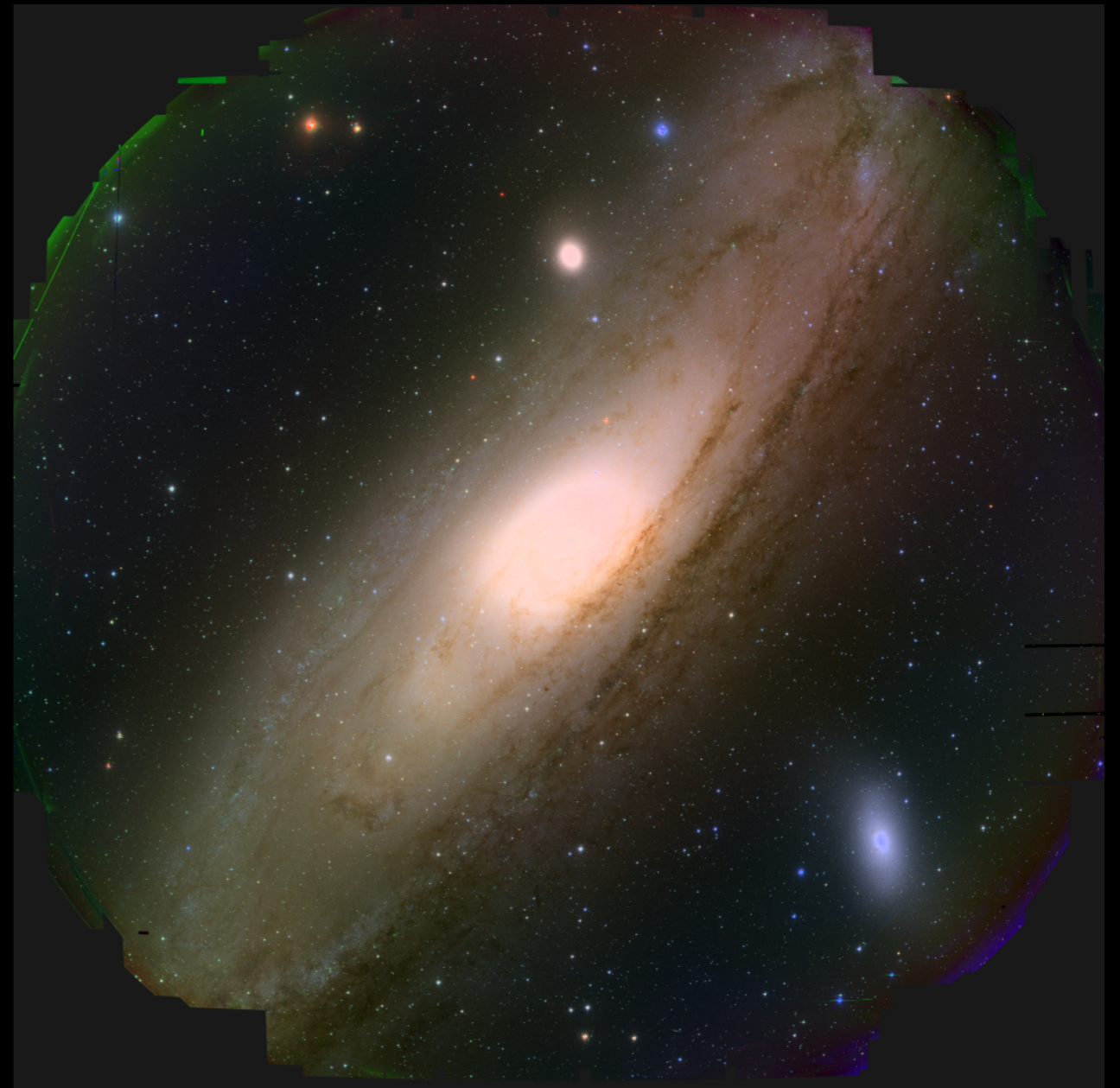
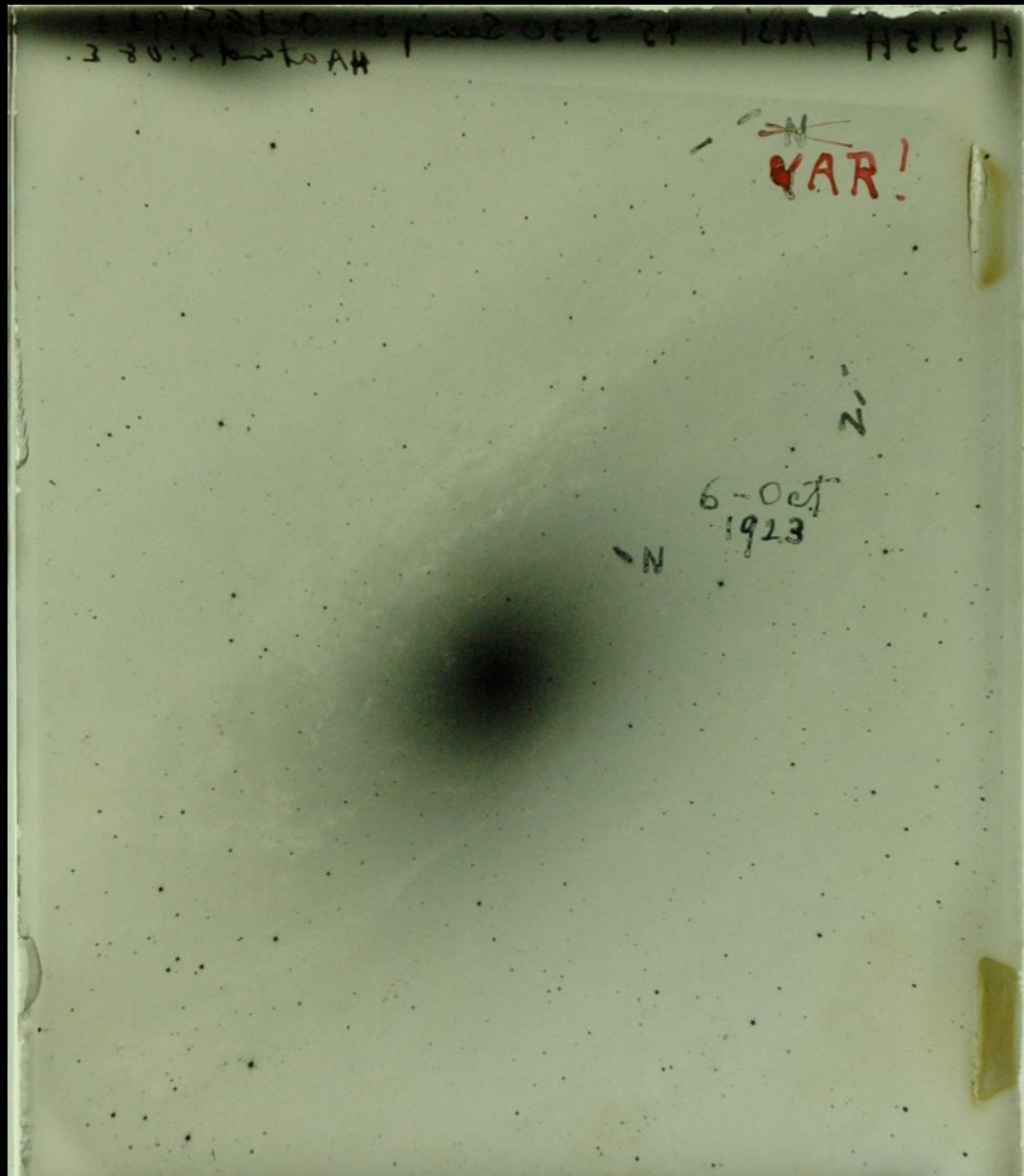


What's Next?

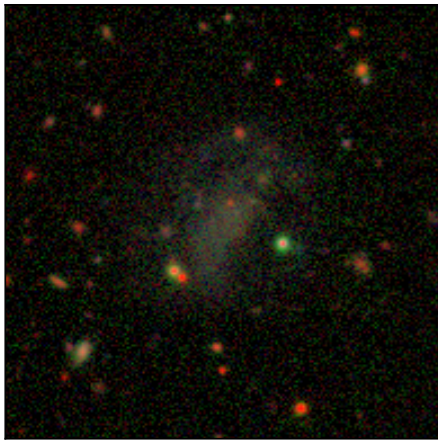
Redshifts with GMOS on Gemini



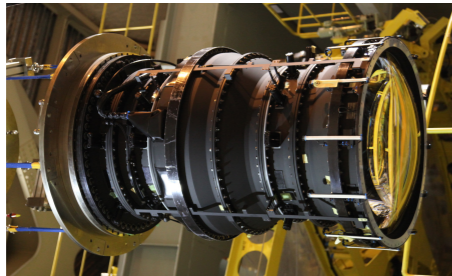
Progress



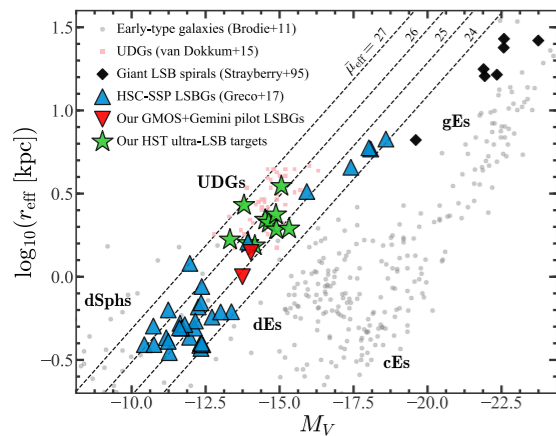
Summary



Low-surface-brightness galaxies (LSBGs) are a significant yet poorly understood component of the galaxy population



Exquisite imaging afforded by HSC promises to provide an unprecedented view of LSBGs over 1400 deg²



We have a diverse sample of ~ 800 LSBGs, which we are actively following up from the ground and space

Special thanks to my collaborators!

Jenny Greene, Michael Strauss, David Spergel, **Andy Goulding**, Lauren MacArthur, Robert Lupton, Alexie Leauthaud, Song Huang, and the HSC collaboration