

# A Unified Scenario for the Origin of Spiral and Elliptic Galaxy Structural Scaling Laws

arXiv:2009.03916

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ISMAEL FERRERO

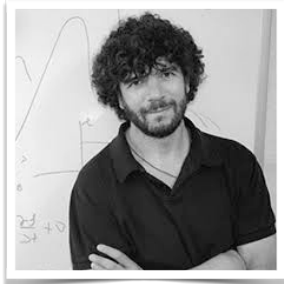


UiO : **Institute of Theoretical Astrophysics**  
University of Oslo



LineA Webinar - Nov 19, 2020

# Collaborators



Julio Navarro

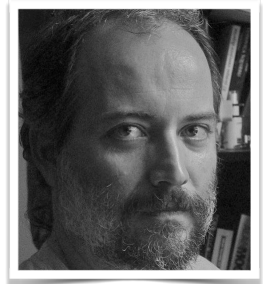
Victoria, Canada



Mario Abadi



José Benavides



Damián Mast

Córdoba, Argentina

# Hubble Morphological Classification

Hubble 1926, 1936

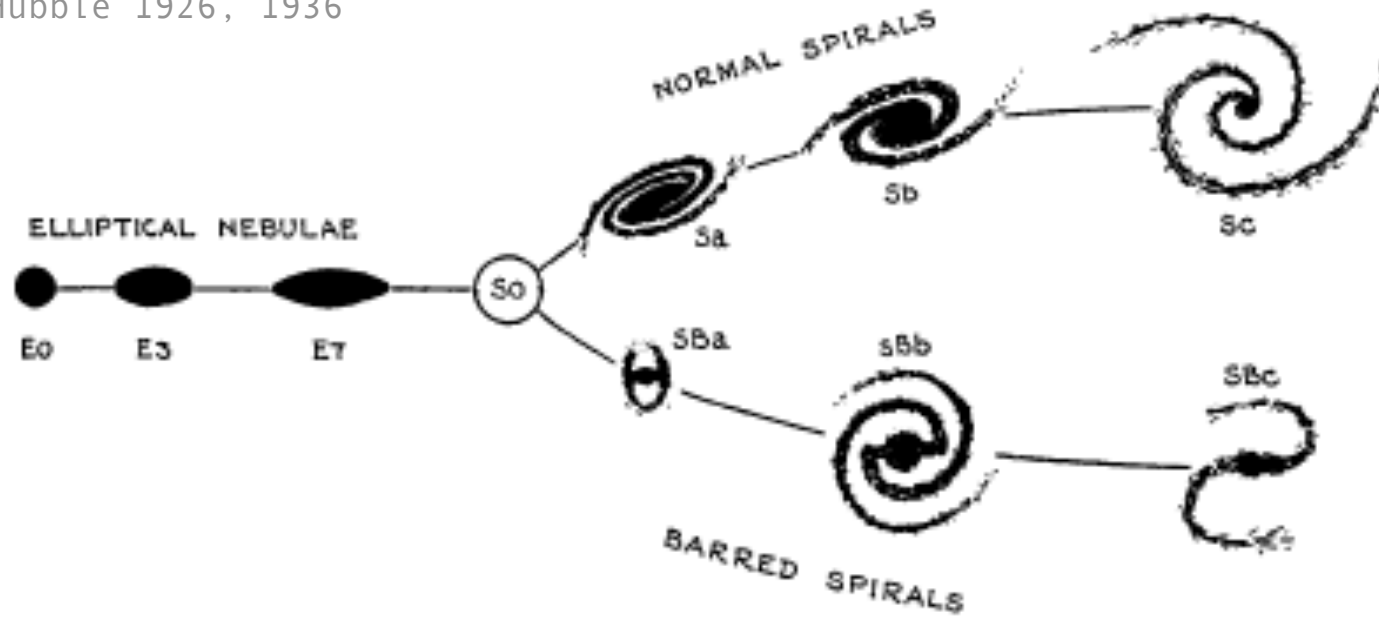
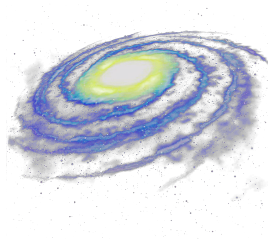
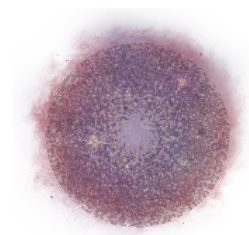


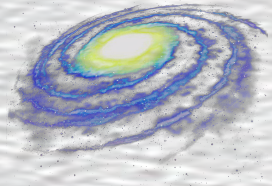
FIG. 1. *The Sequence of Nebular Types.*



# Spirals vs Ellipticals

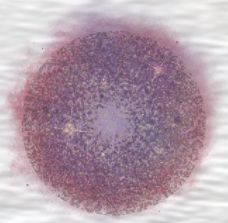




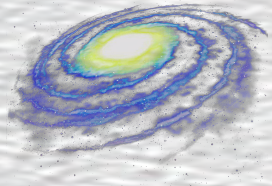


Discoidal

# Spirals vs Ellipticals

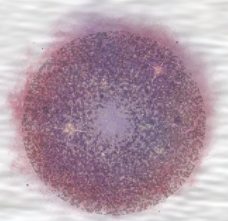


Spheroidal

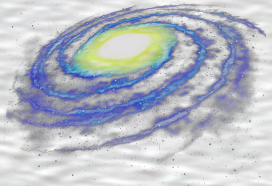


Discoidal  
Blue

# Spirals vs Ellipticals

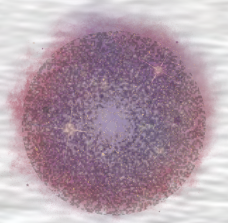


Spheroidal  
Red

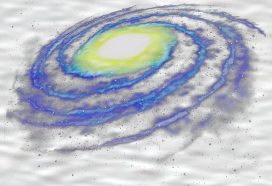


Discoidal  
Blue  
Structured

# Spirals vs Ellipticals

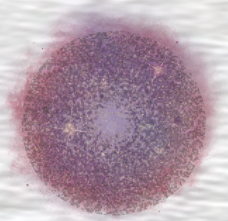


Spheroidal  
Red  
Smooth



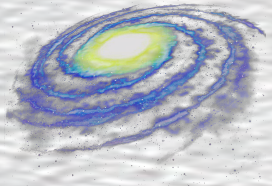
Discoidal  
Blue  
Structured  
Star Forming

# Spirals vs Ellipticals



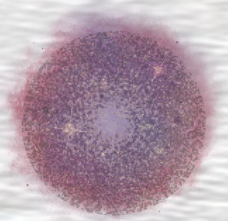
Spheroidal  
Red  
Smooth  
Quenched



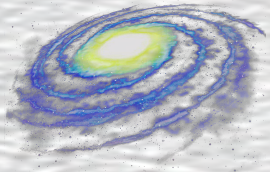


Discoidal  
Blue  
Structured  
Star Forming  
Field

# Spirals vs Ellipticals

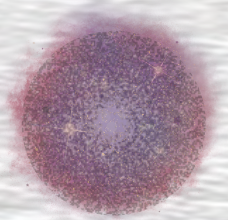


Spheroidal  
Red  
Smooth  
Quenched  
Cluster

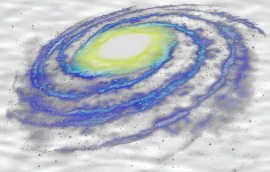


Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich

# Spirals vs Ellipticals

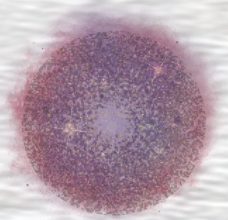


Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor



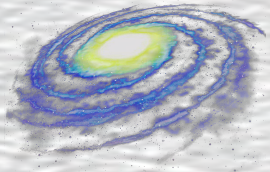
Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor

# Spirals vs Ellipticals



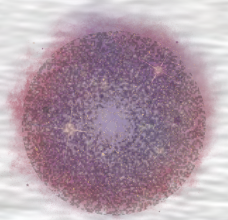
Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich





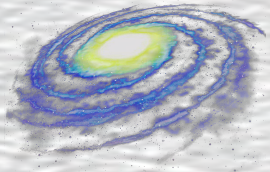
Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young

# Spirals vs Ellipticals



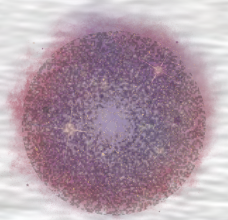
Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old



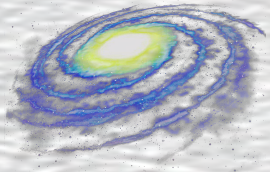


Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse

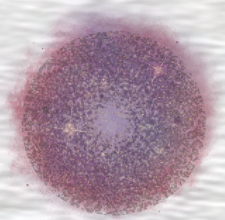
# Spirals vs Ellipticals



Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge



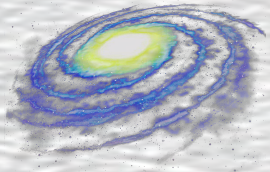
# Spirals vs Ellipticals



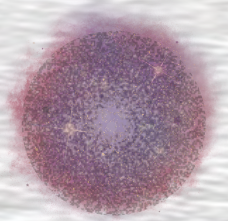
Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse  
Exponential

Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge  
Power law



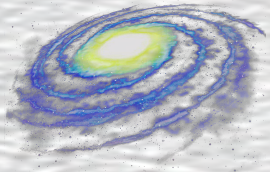


# Spirals vs Ellipticals

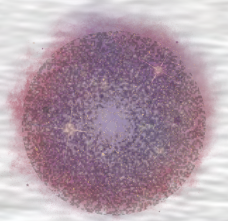


Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse  
Exponential  
Rotational

Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge  
Power law  
Dispersion



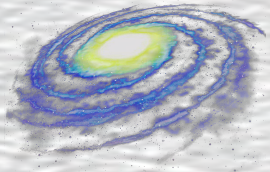
# Spirals vs Ellipticals



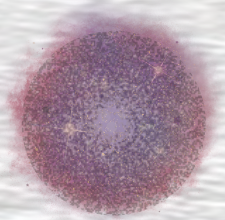
Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse  
Exponential  
Rotational  
Low mass end

Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge  
Power law  
Dispersion  
High mass end





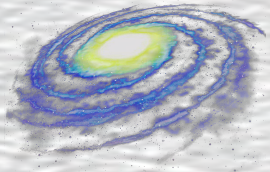
# Spirals vs Ellipticals



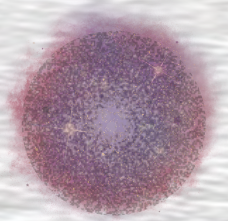
Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse  
Exponential  
Rotational  
Low mass end  
Extended

Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge  
Power law  
Dispersion  
High mass end  
Compact



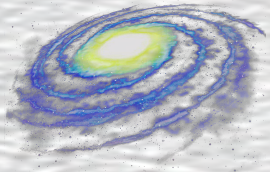


# Spirals vs Ellipticals

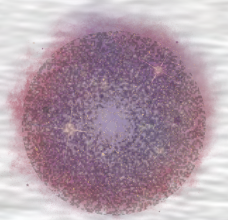


Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse  
Exponential  
Rotational  
Low mass end  
Extended  
Tully-Fisher

Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge  
Power law  
Dispersion  
High mass end  
Compact  
Faber-Jackson



# Spirals vs Ellipticals



Discoidal  
Blue  
Structured  
Star Forming  
Field  
Gas Rich  
Metal Poor  
Young  
Collapse  
Exponential

Spheroidal  
Red  
Smooth  
Quenched  
Cluster  
Gas Poor  
Metal Rich  
Old  
Merge  
Power law

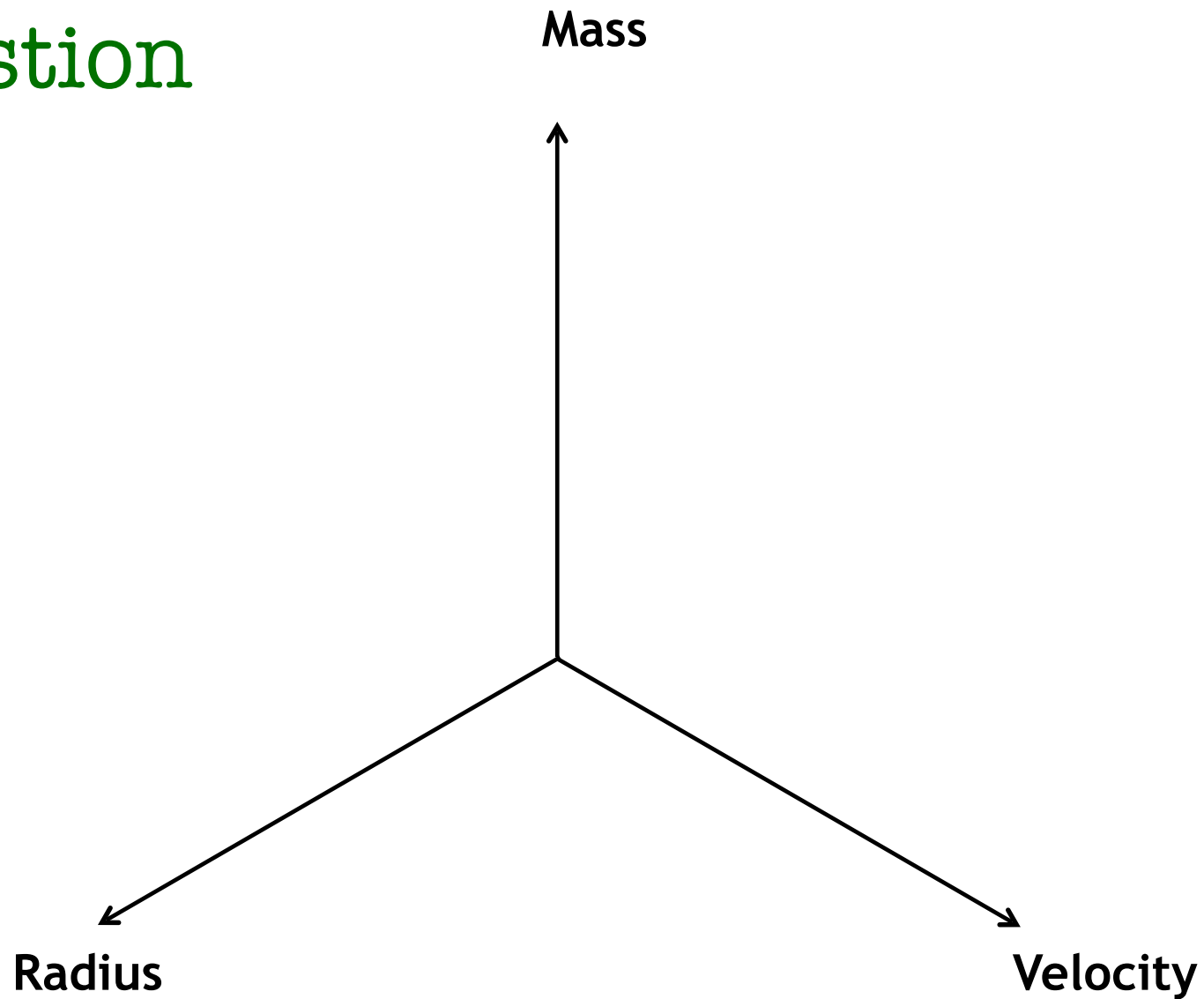
Rotational  
Low mass end  
Extended  
Tully-Fisher

Dispersion  
High mass end  
Compact  
Faber-Jackson



# Key Question

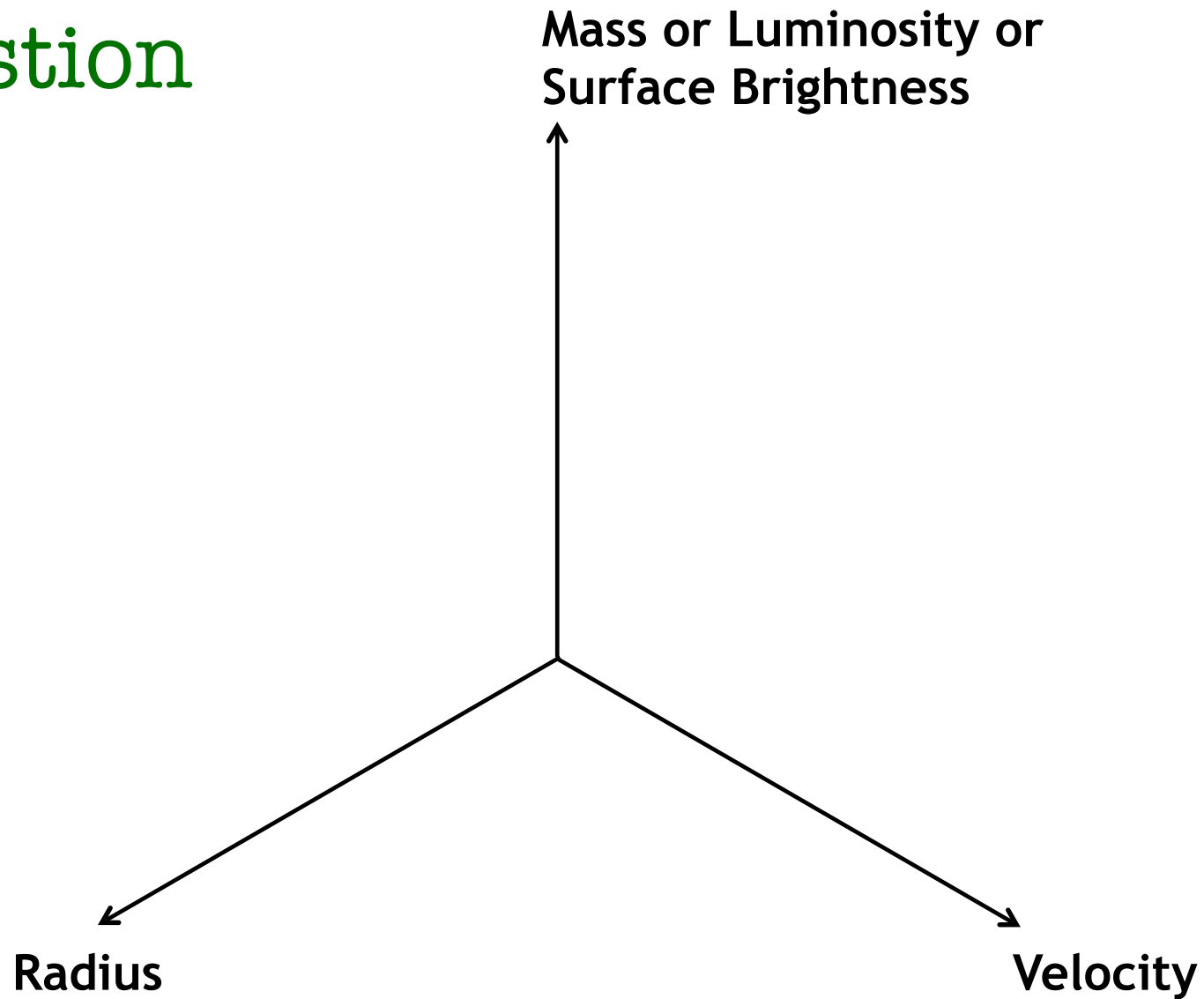
Why spiral and ellipticals obey **different** scaling laws?





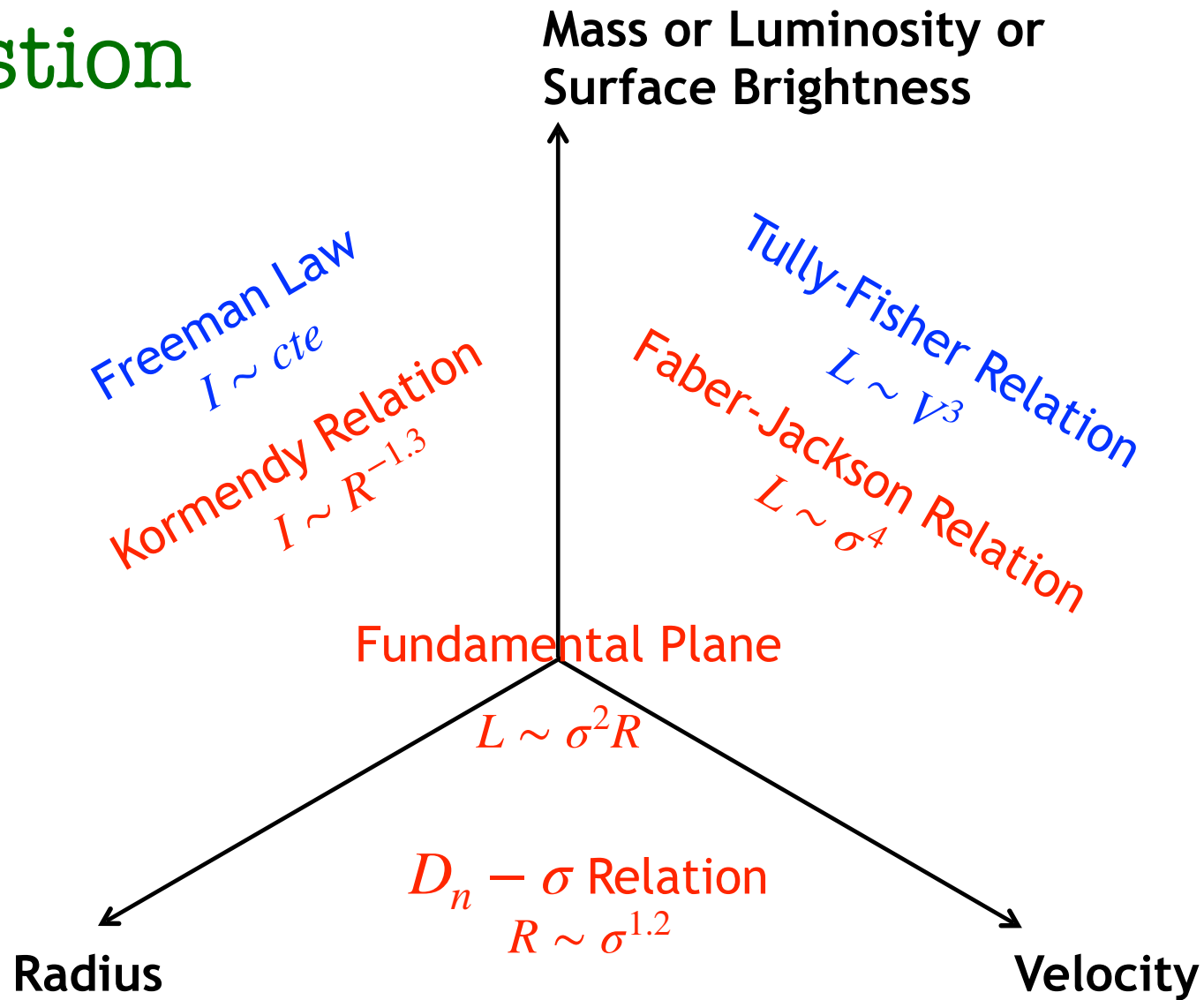
# Key Question

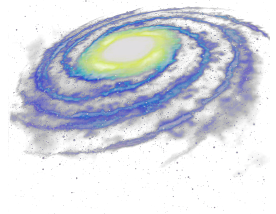
Why spiral and ellipticals obey **different** scaling laws?



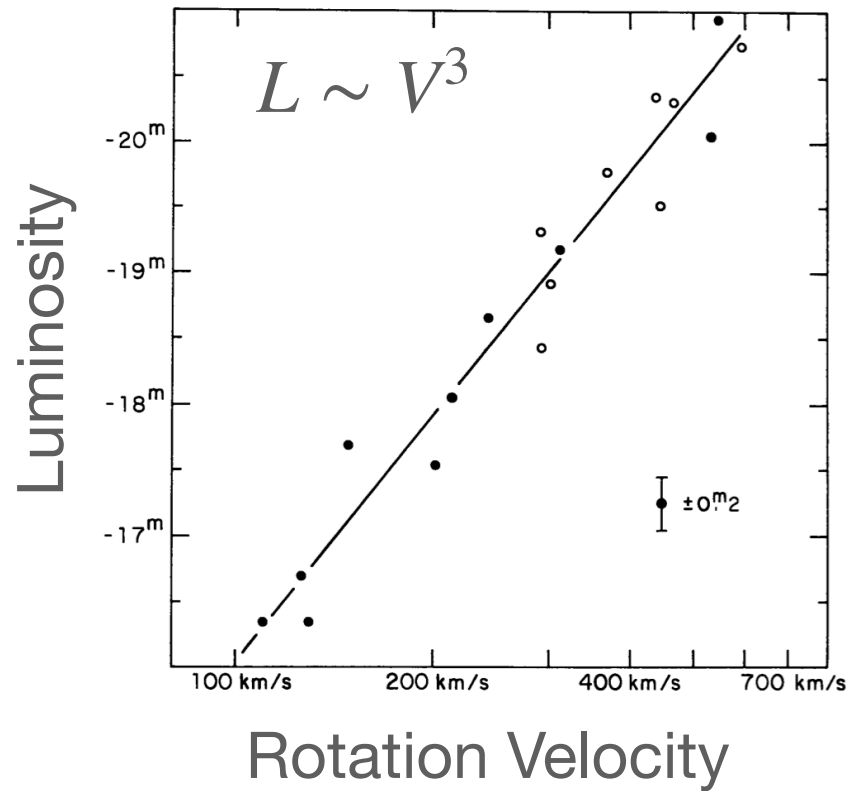
# Key Question

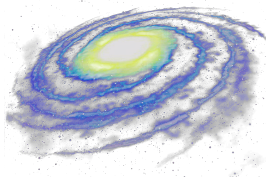
Why **spiral** and **ellipticals** obey **different** scaling laws?





# Tully-Fisher 1977

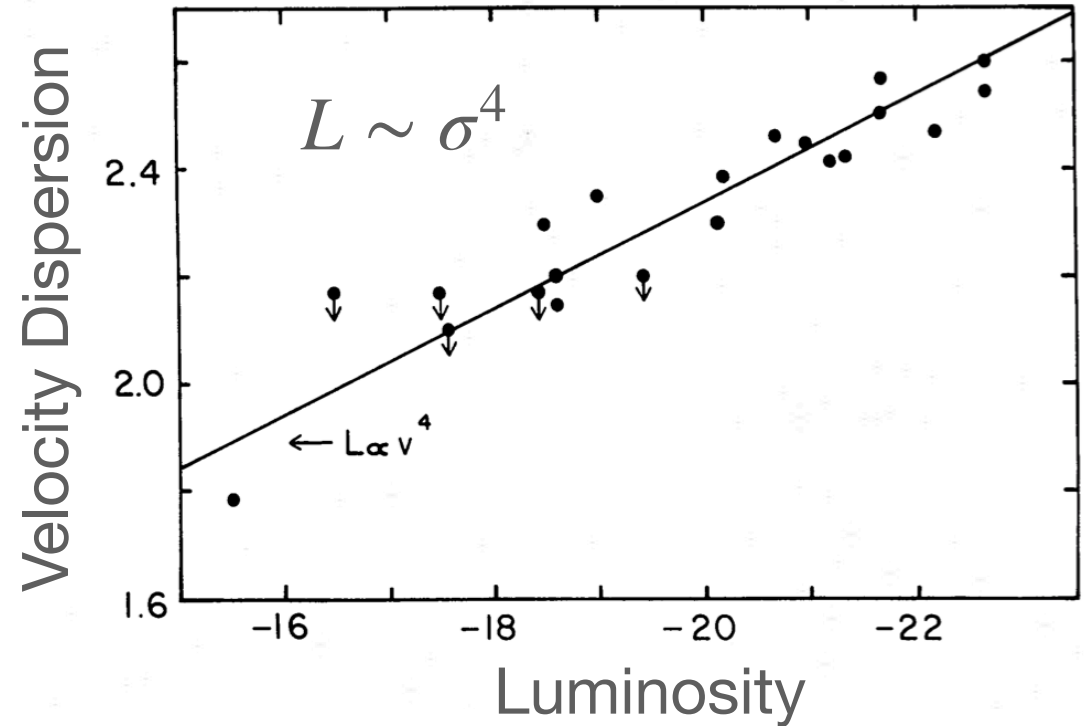
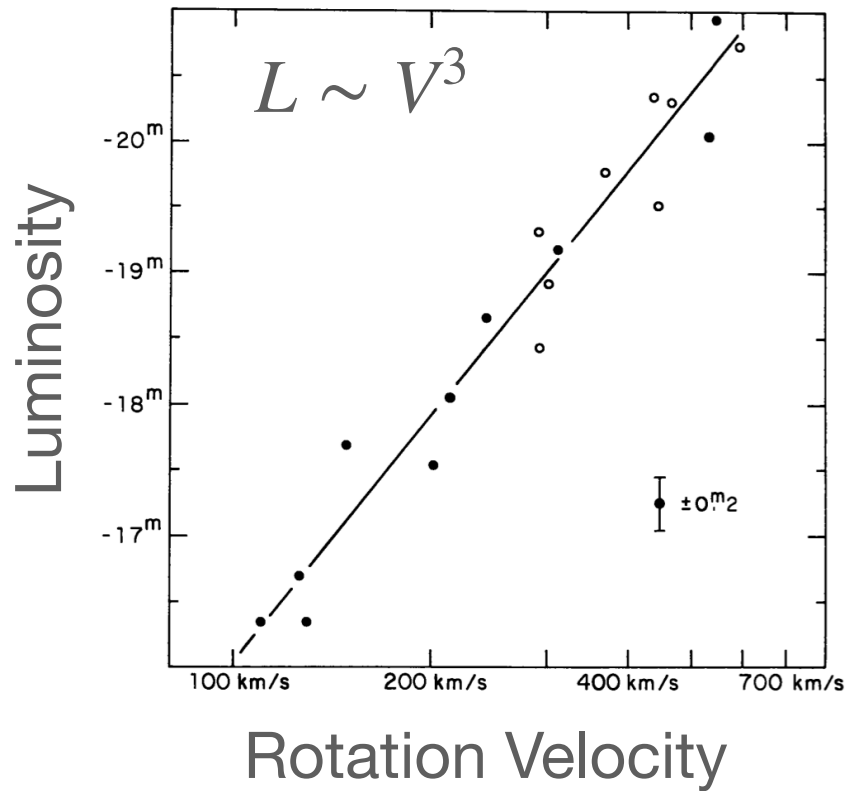
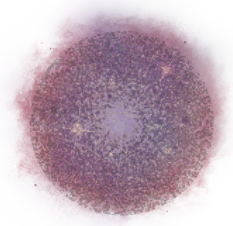


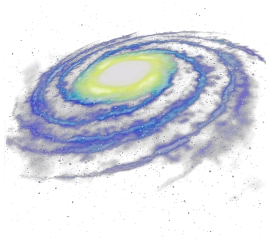


# Tully-Fisher 1977

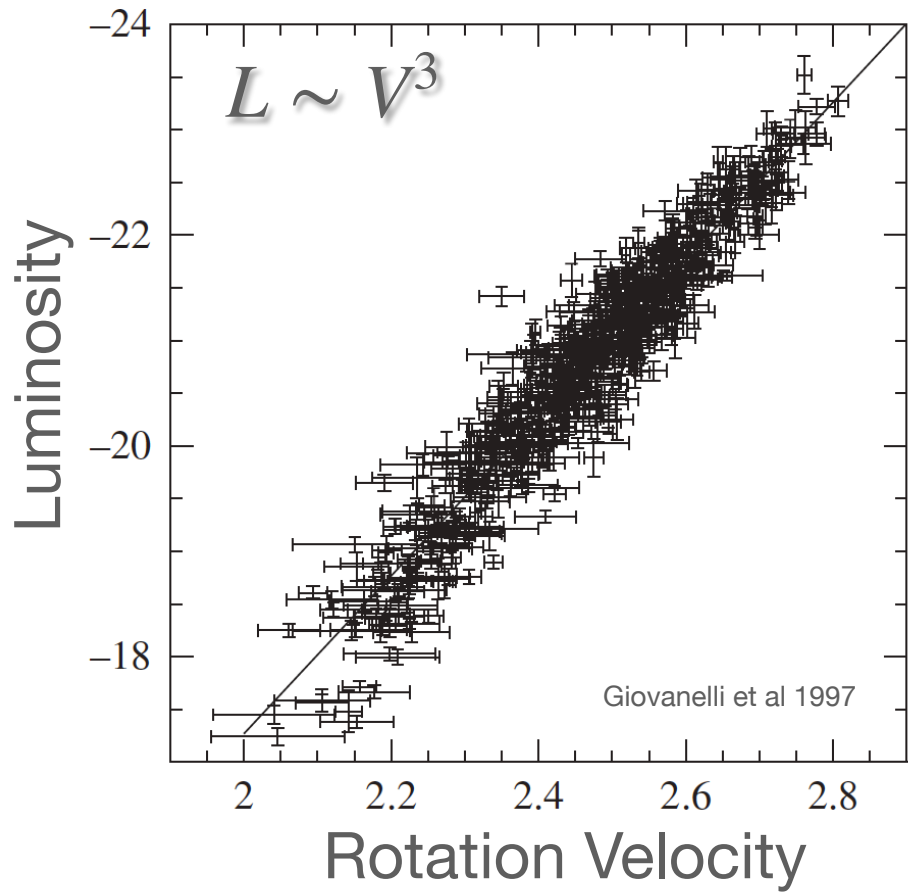
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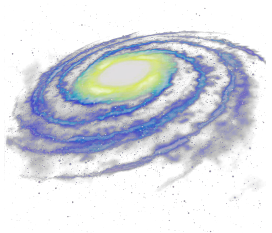
# Faber-Jackson 1976



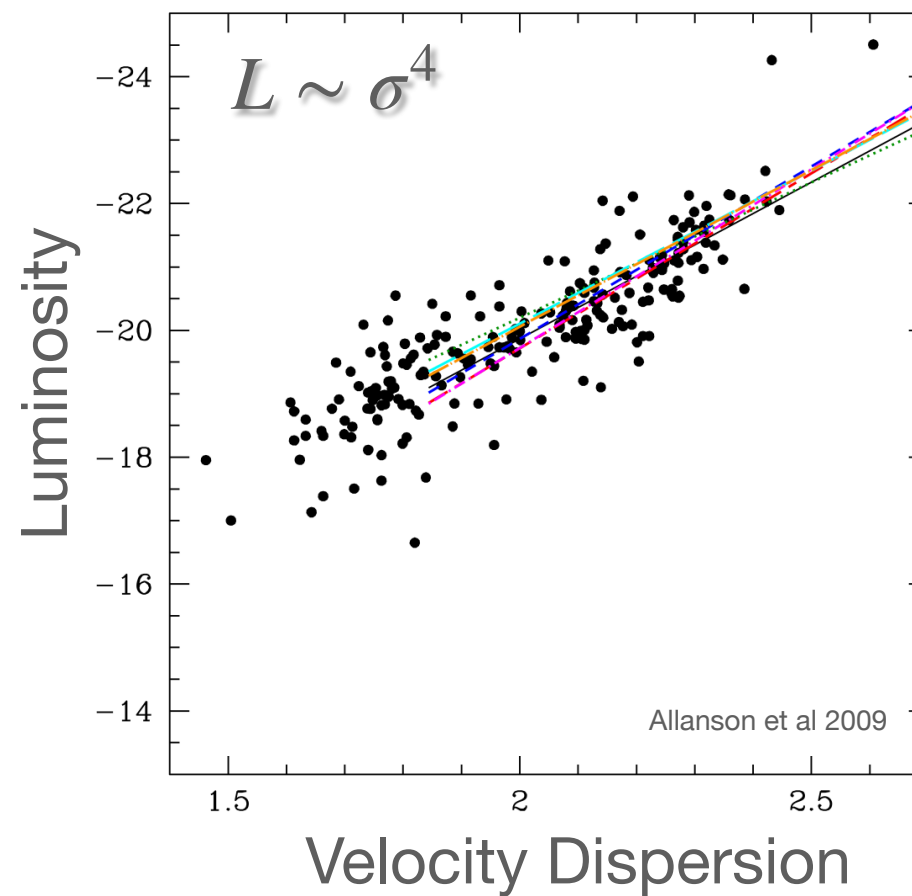
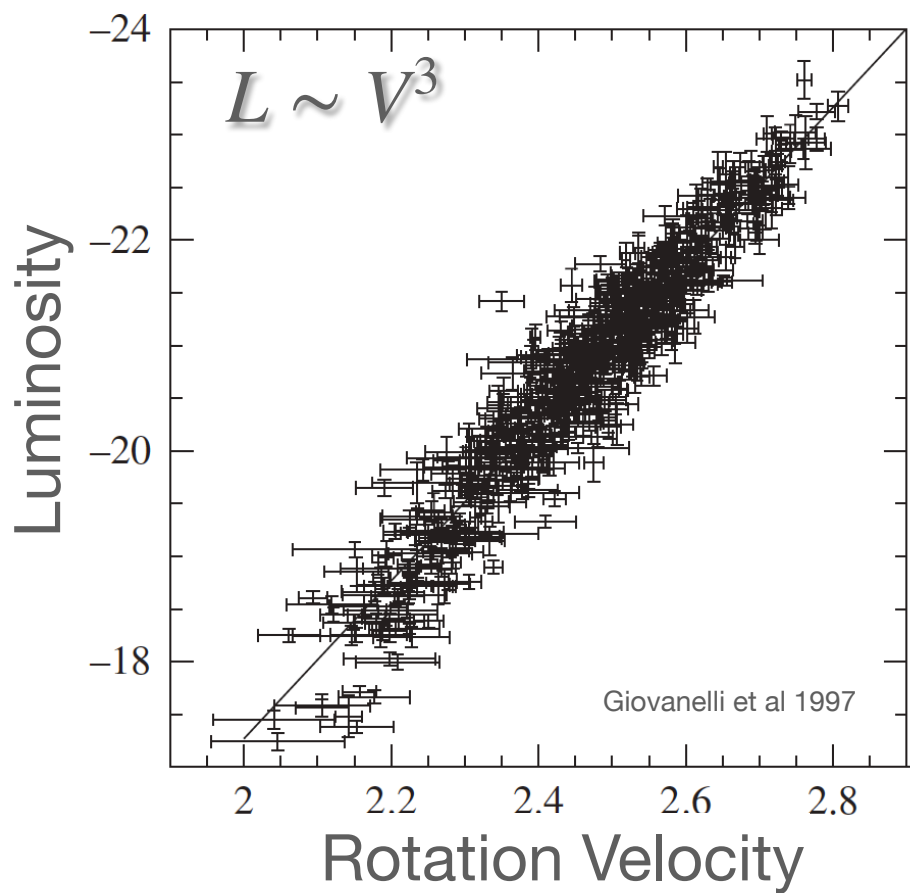
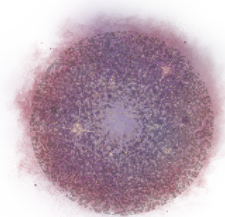


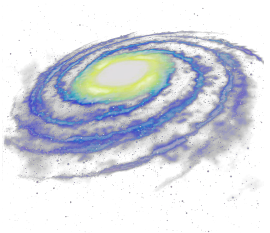
# Spiral & Elliptical Scaling Relations



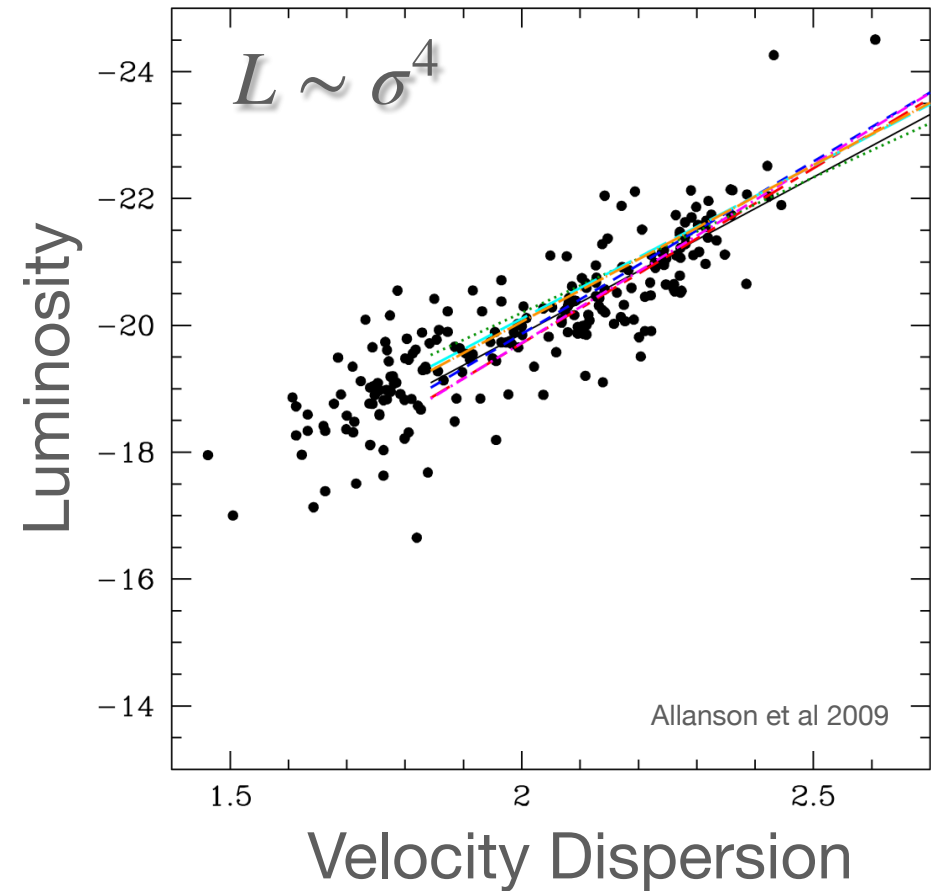
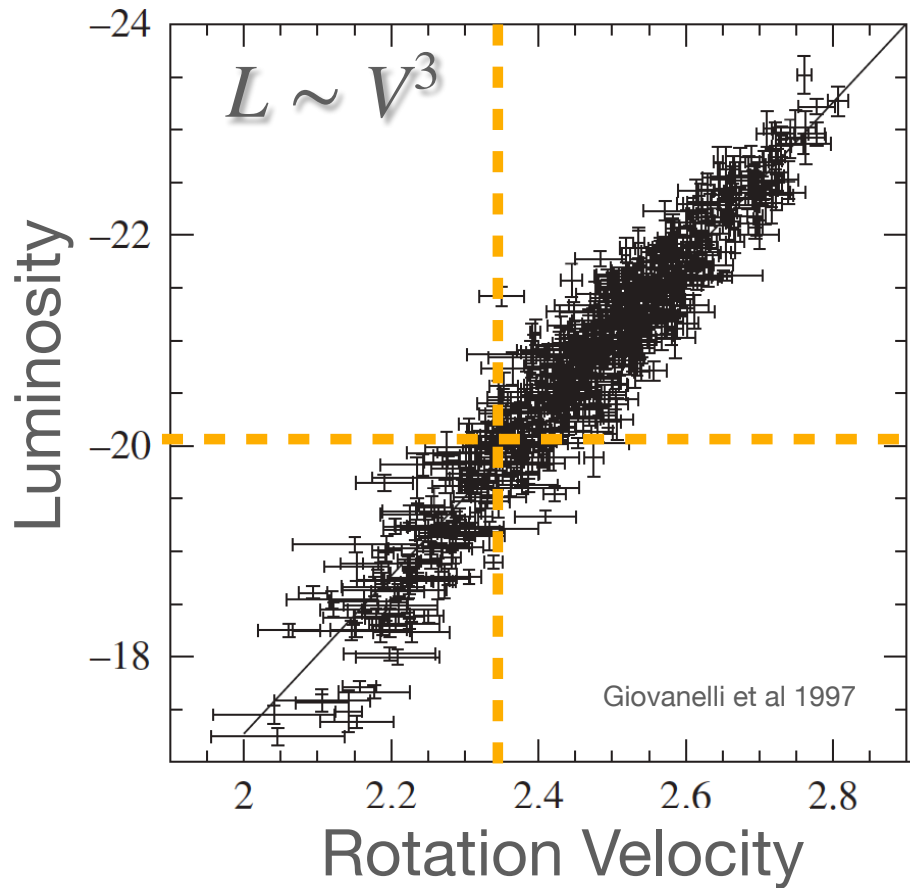
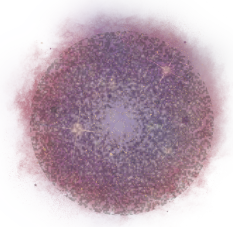


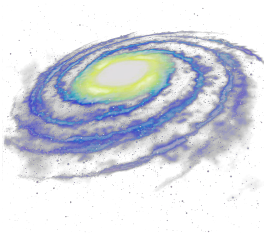
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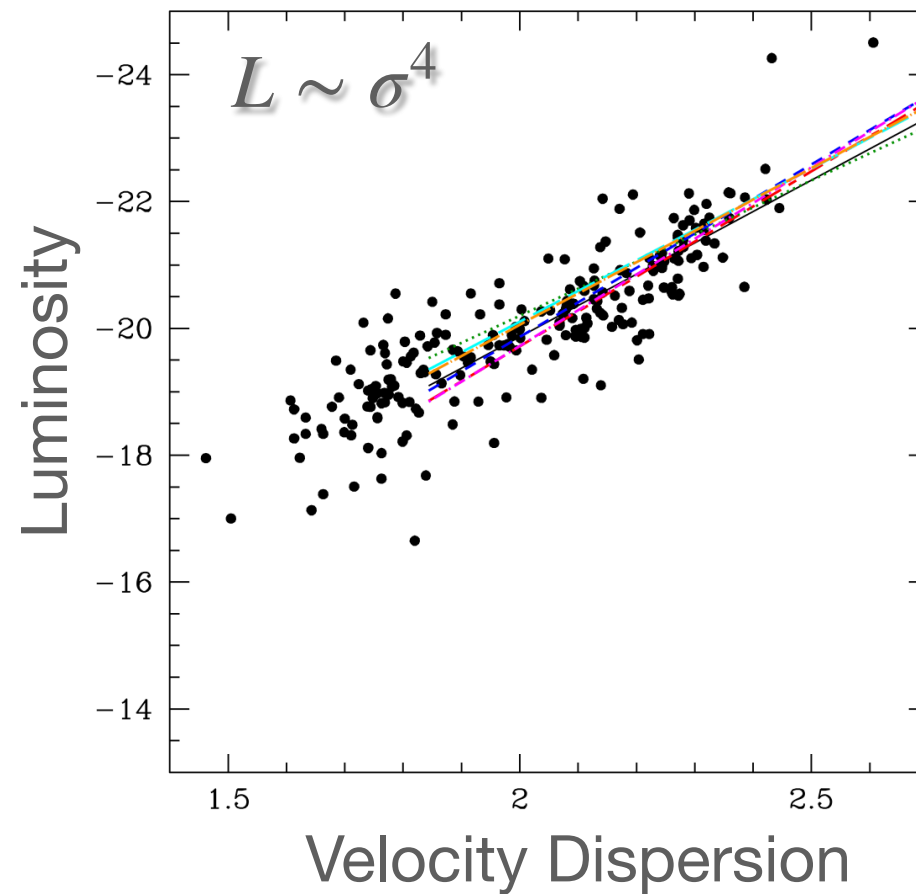
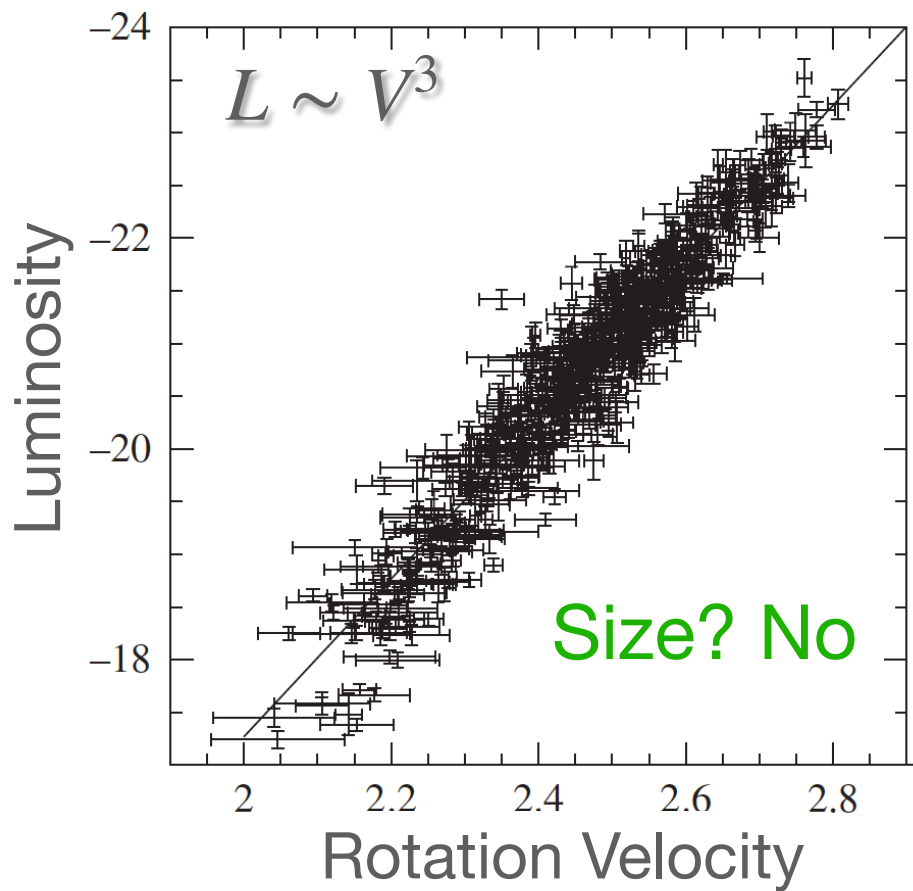
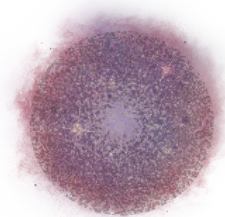


# Spiral & Elliptical Scaling Relations

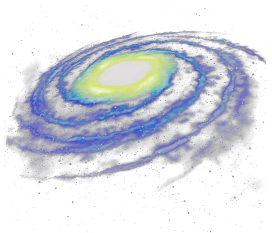




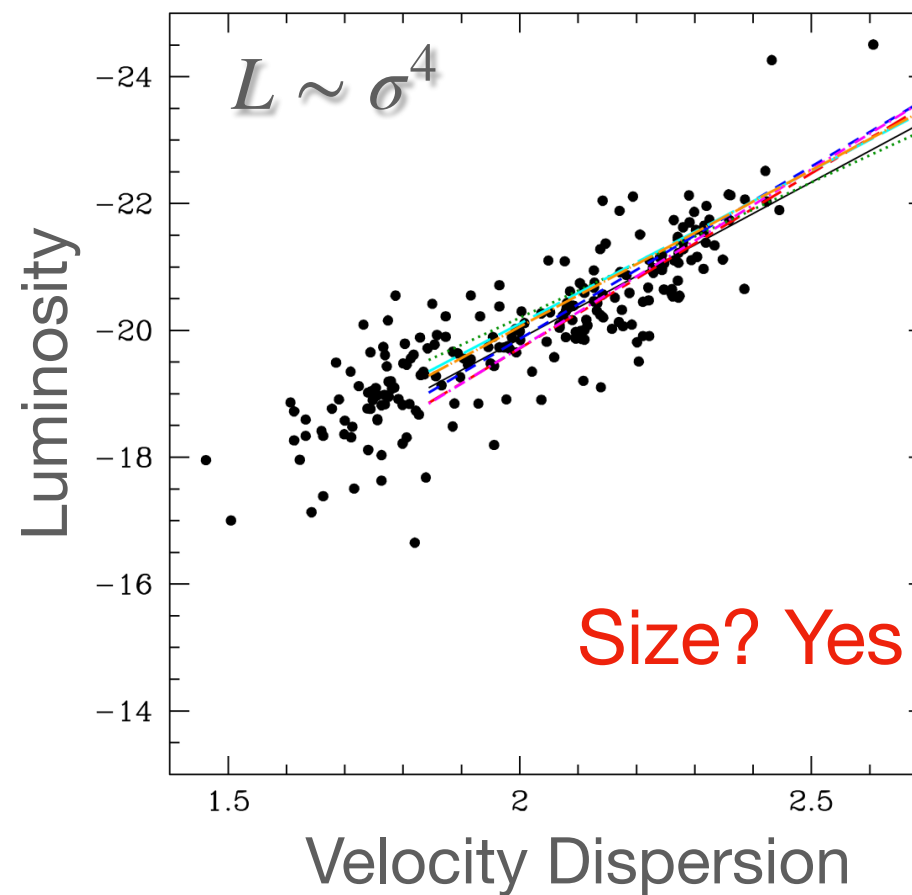
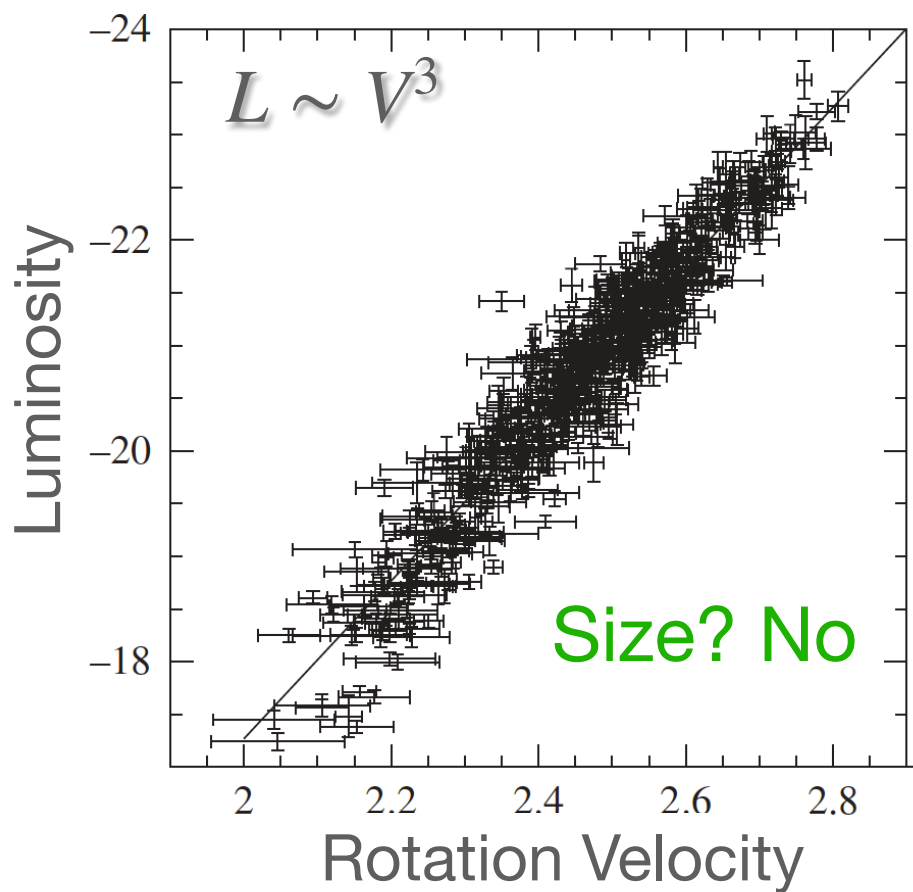
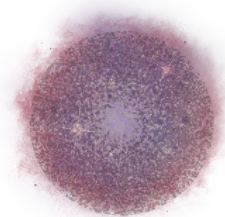
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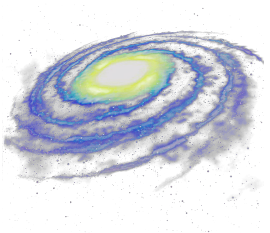




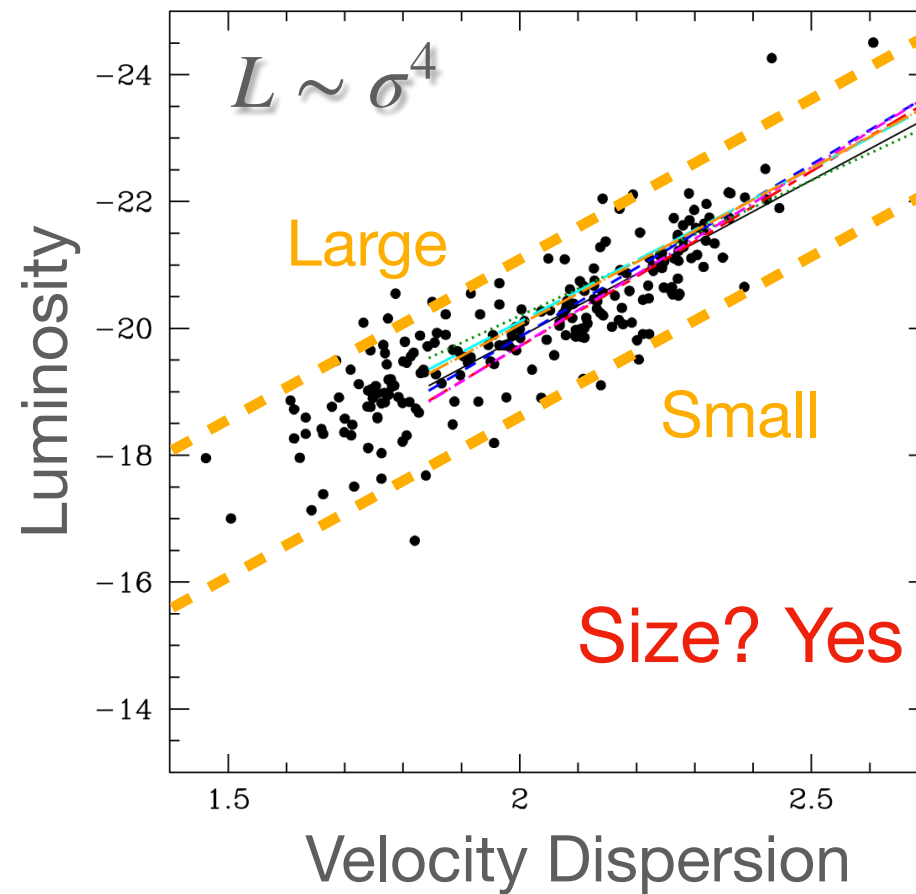
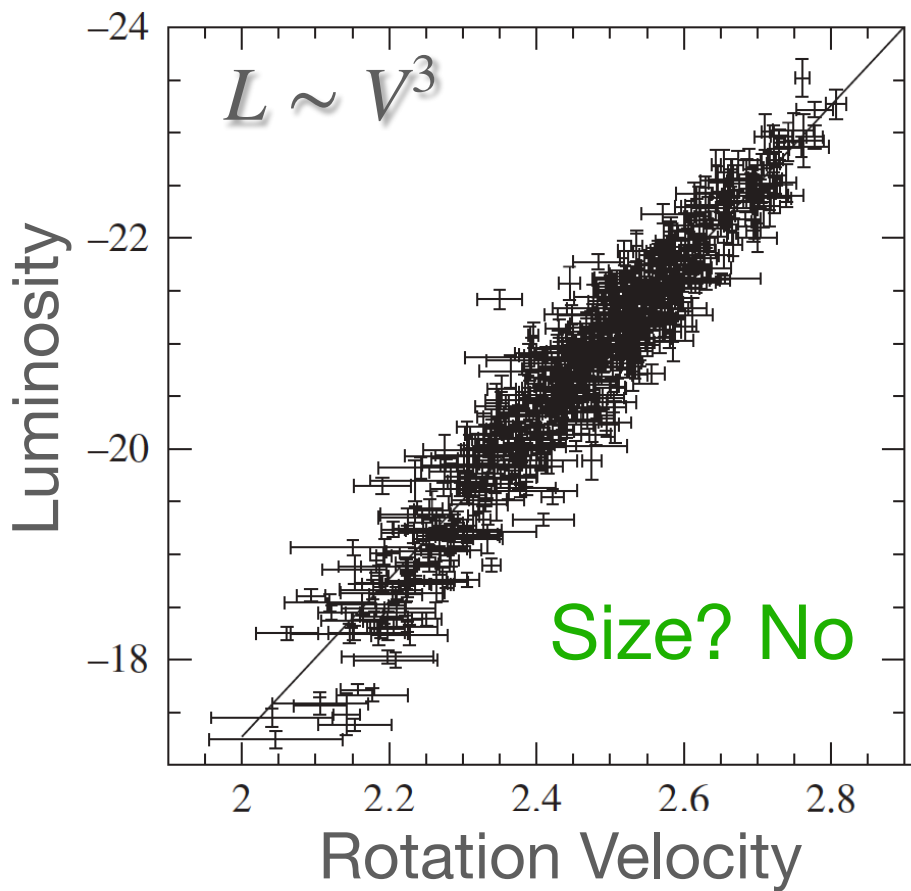
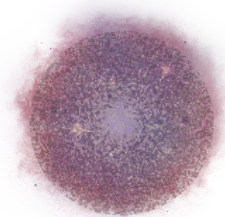


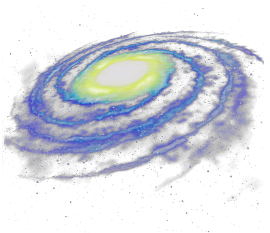
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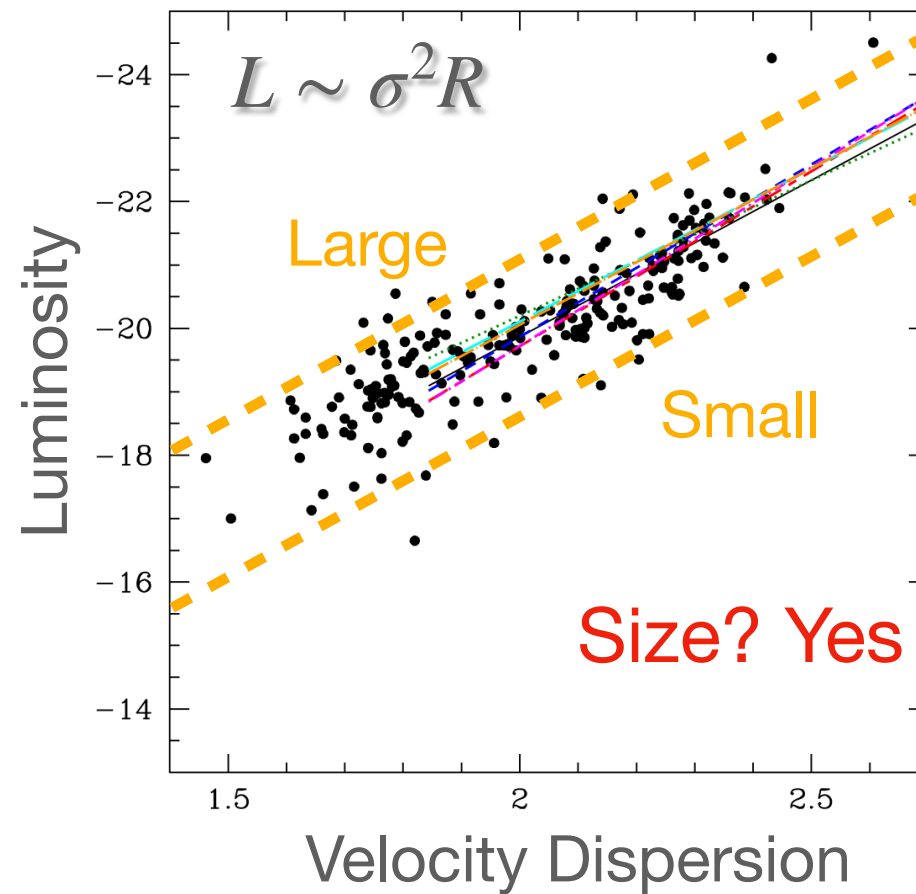
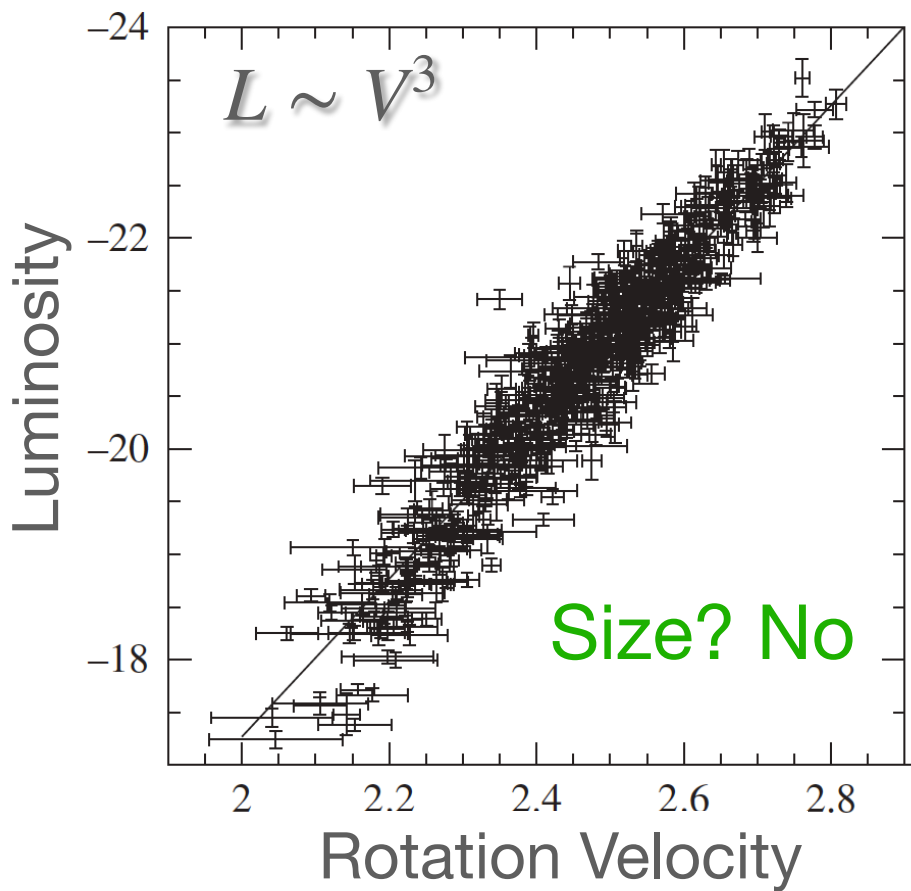
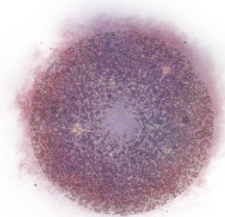


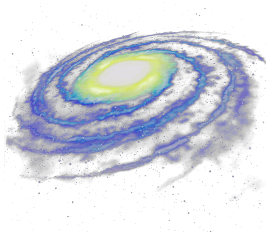
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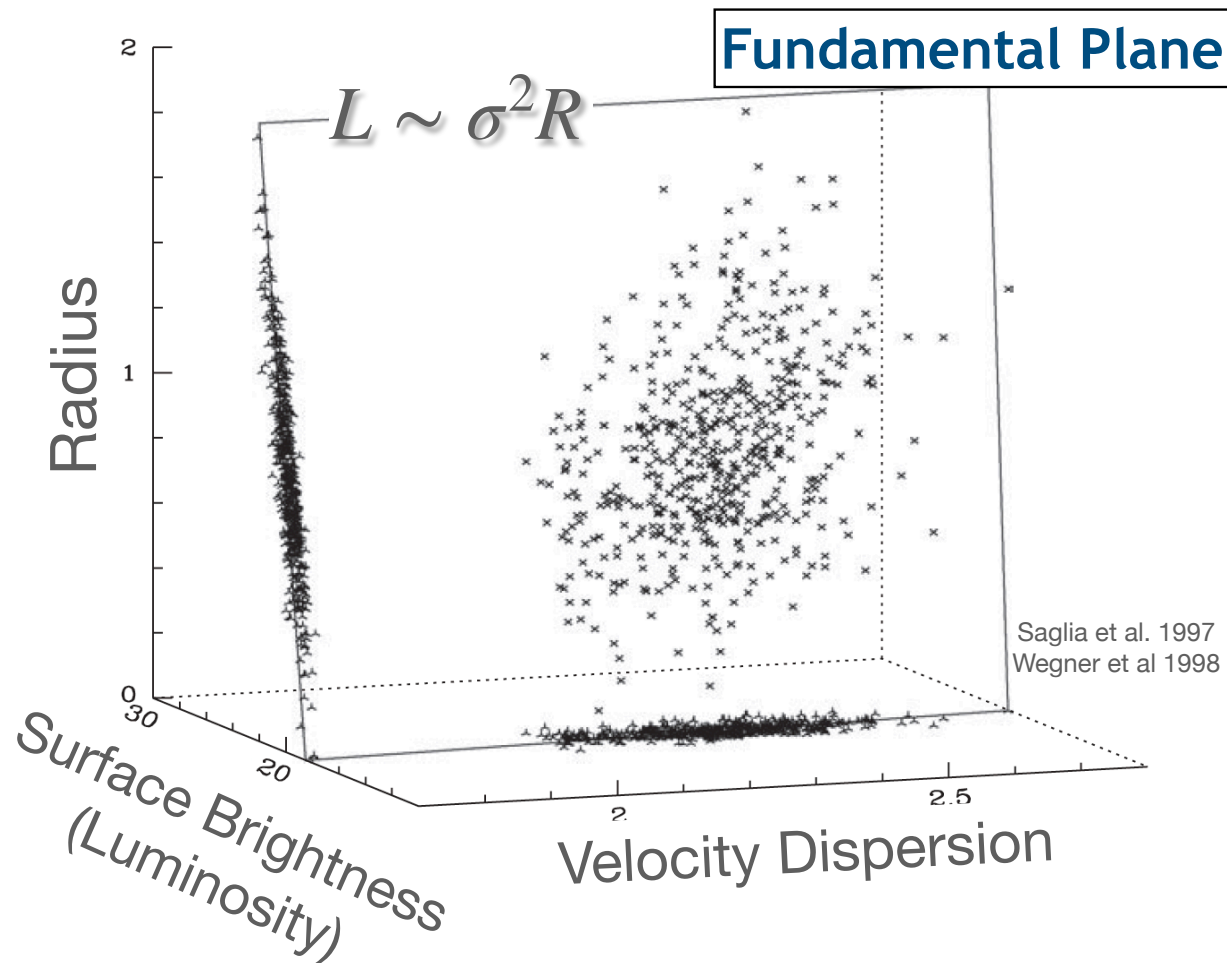
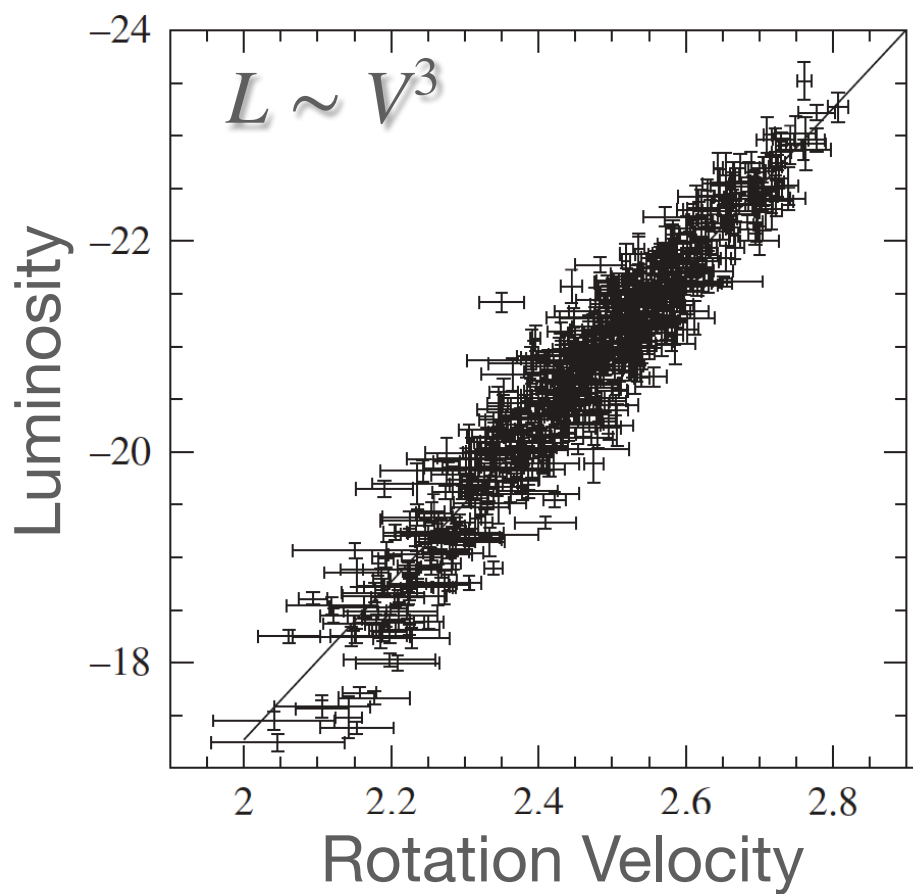
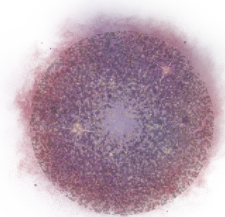


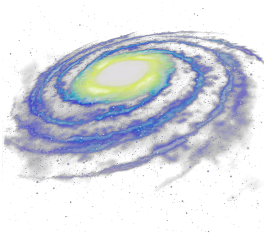
# Spiral & Elliptical Scaling Relations



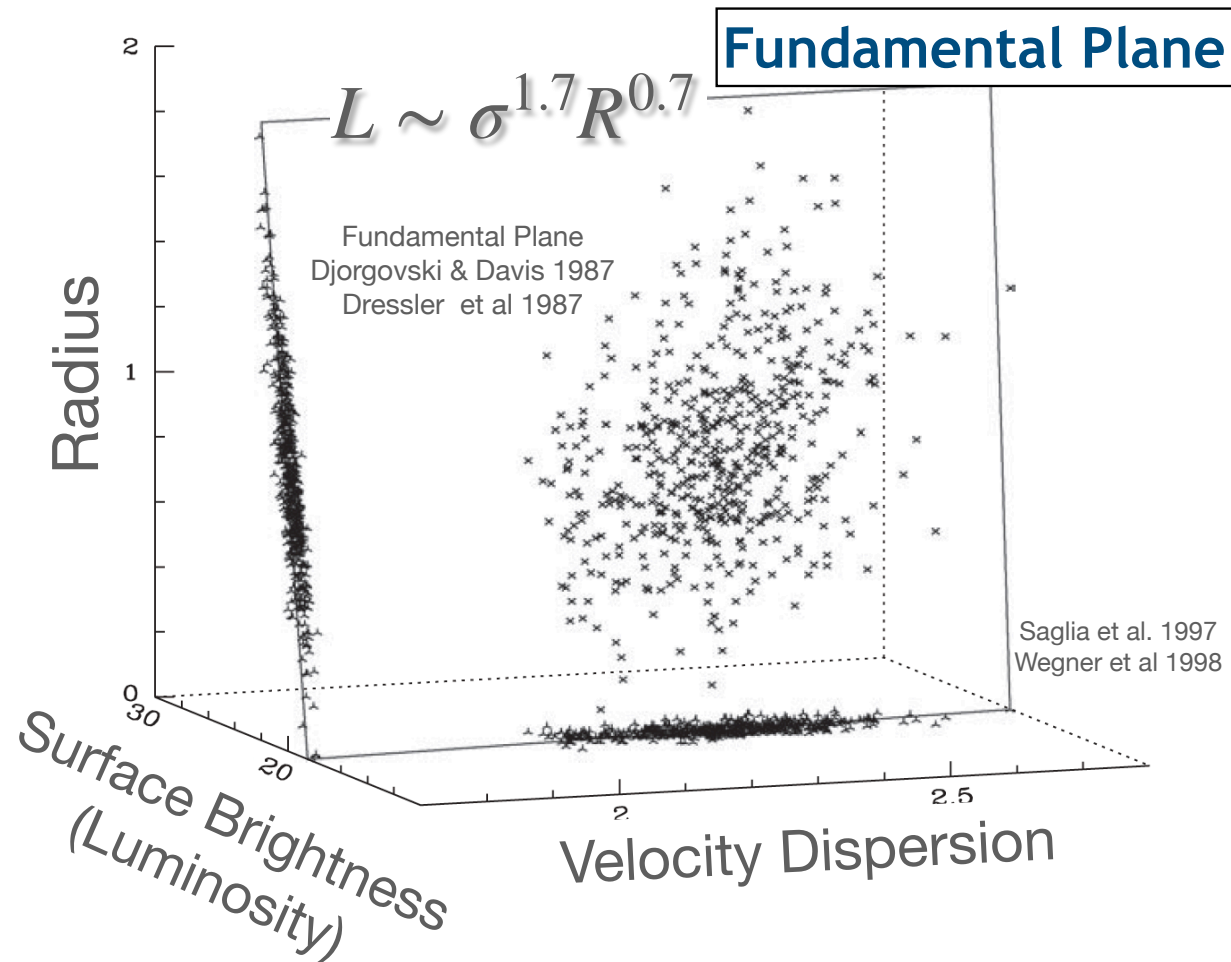
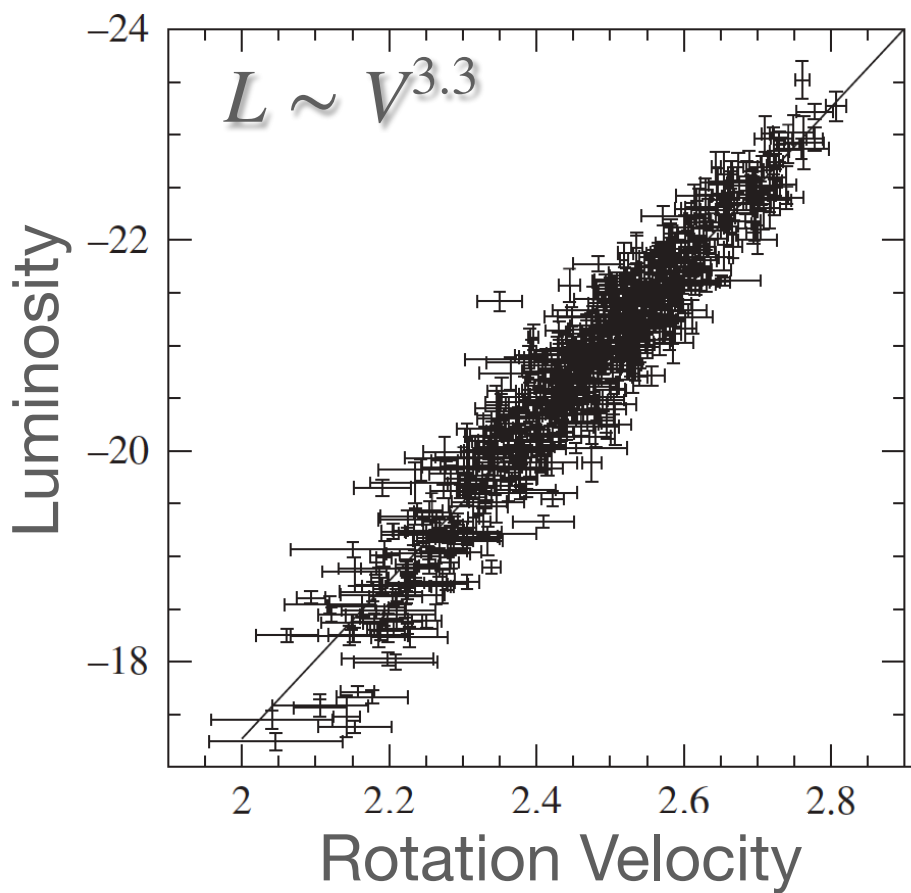
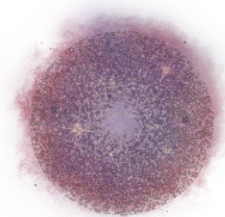


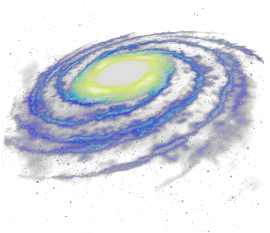
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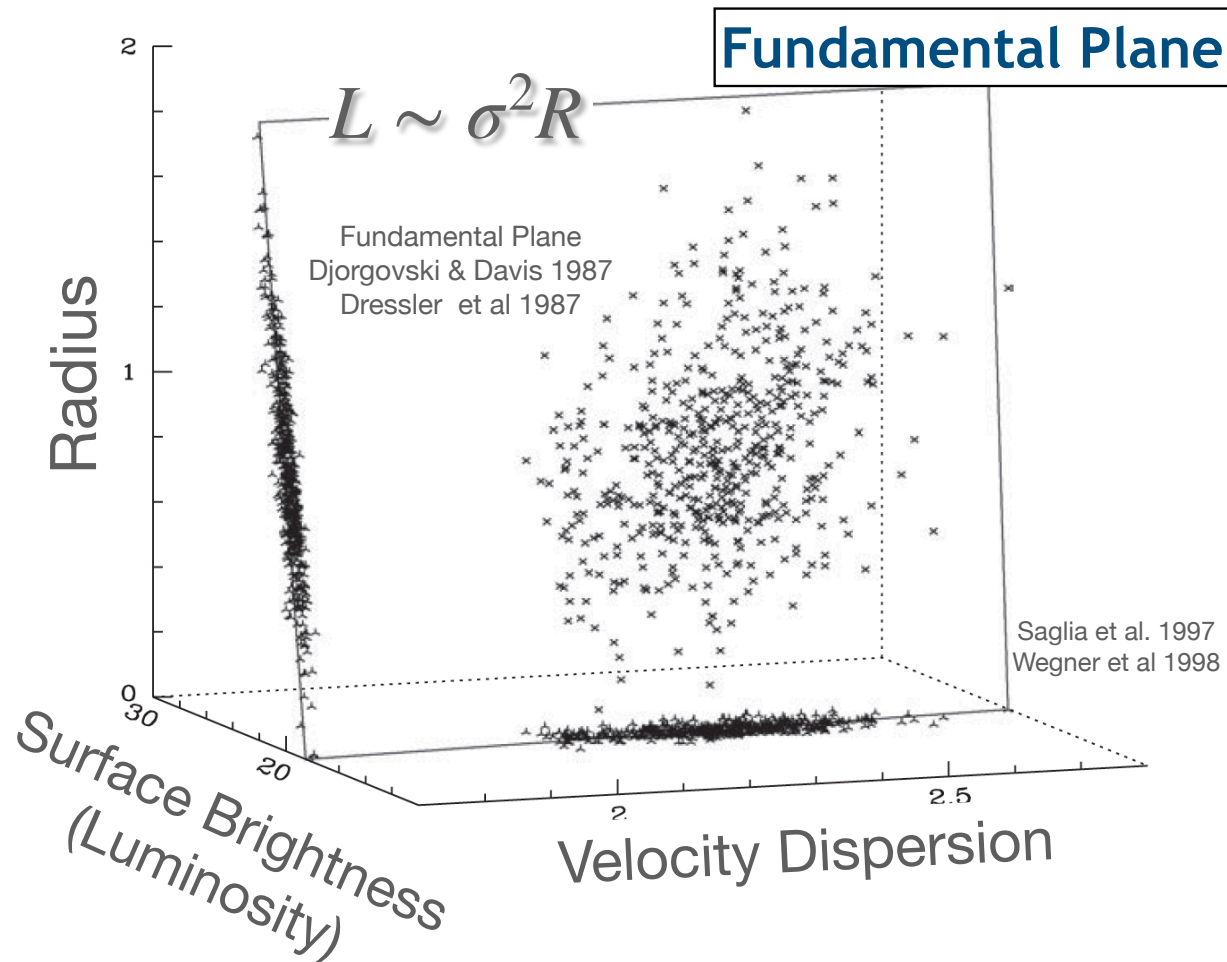
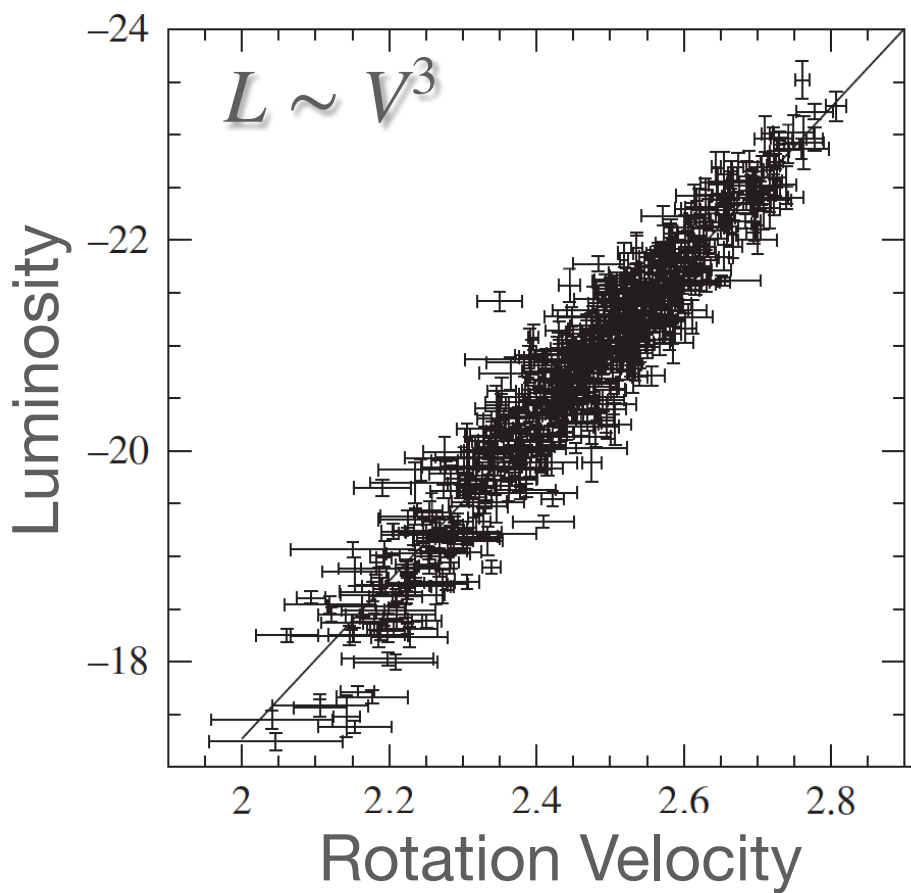
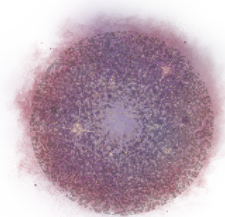


# Spiral & Elliptical Scaling Relations

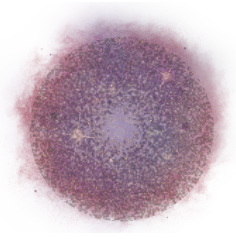




# Spiral & Elliptical Scaling Relations

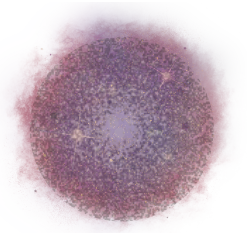


# System in Equilibrium

$$2T+W=0$$


Virial theorem  $\sigma^2 = \frac{GM}{R}$

# System in Equilibrium

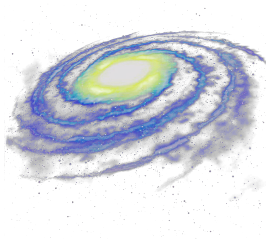
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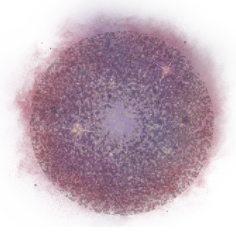
=

Fundamental Plane  $M = \frac{\sigma^2 R}{G}$





# System in Equilibrium

$$2T+W=0$$


Virial theorem

$$V^2 = \frac{GM}{R}$$

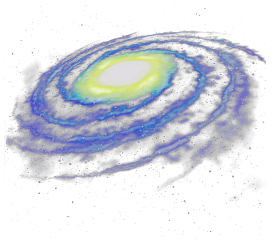
Virial theorem

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=

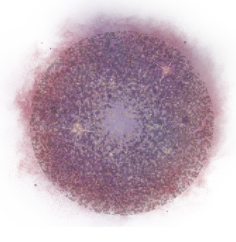
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# System in Equilibrium

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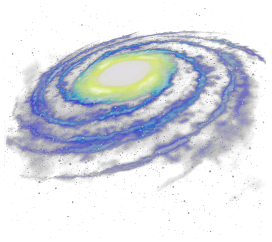
+

Integrated density  
constant

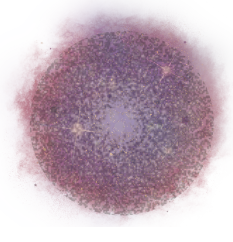
$$\left\{ \begin{array}{l} \rho = \Delta\rho_c \\ \frac{3M}{4\pi R^3} = \Delta \frac{3H^2}{8\pi G} \end{array} \right.$$

=

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=

Tully Fisher

$$M = \sqrt{\frac{2}{\Delta}} \frac{V^3}{GH}$$

=

Fundamental Plane

$$M = \frac{\sigma^2 R}{G}$$

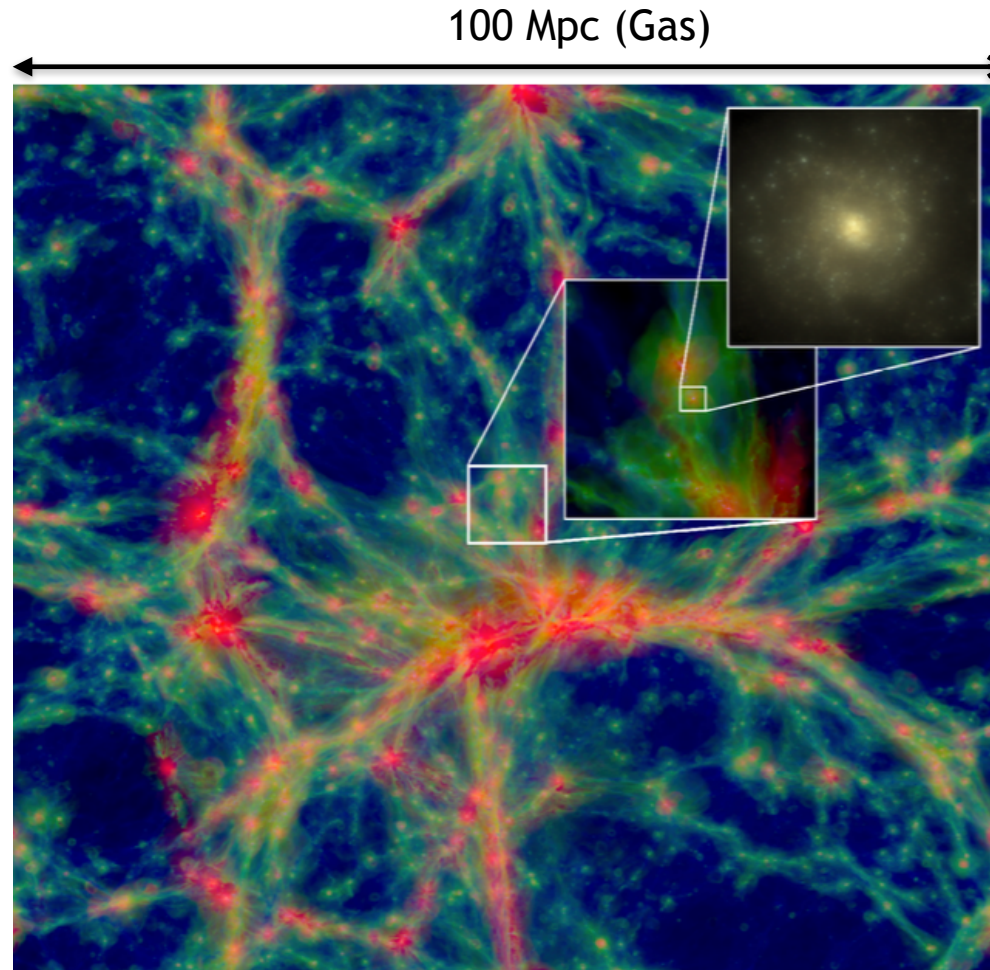
# Cosmological Numerical Simulations



# EAGLE Cosmological Simulations

## Physical Process

- Gravity
- Hydrodynamics
- Radiative cooling
- Star formation
- Feedback SN
- Feedback AGN
- Metallicity



100 Mpc (Gas)

60 kpc (Stars)

$$M_{stellar} = 3 \times 10^{10} M_{\odot}$$

## Particles:

Gas, stars and dark matter

## Code:

GADGET-3 Springel et al 2005

## Cosmological Parameters:

$\Lambda$ CDM model Plank et al 2014

## Particle Mass:

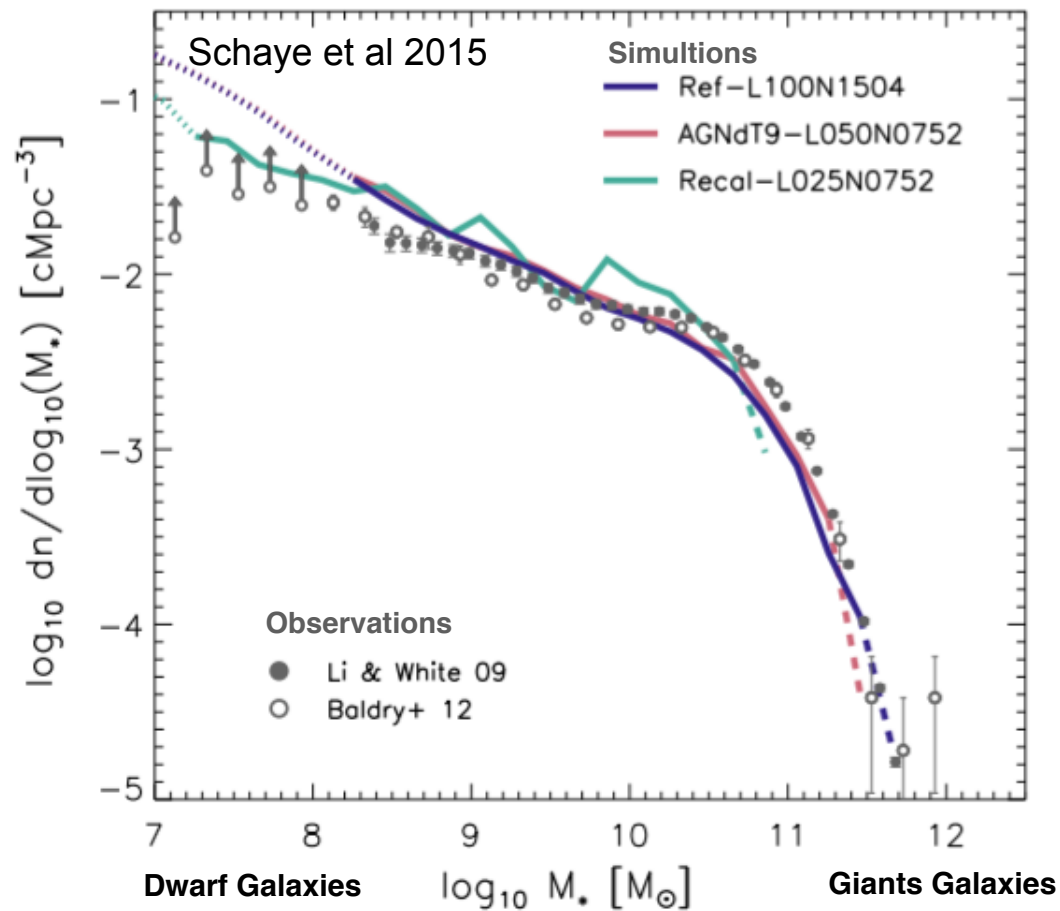
$$M_{gas} = 1.81 \times 10^6 M_{\odot}$$

$$M_{dark} = 9.70 \times 10^6 M_{\odot}$$

Schaye et al 2015



# Stellar Mass Function



The galaxy stellar mass function at  $z = 0.1$  for the EAGLE simulations compared to observations.

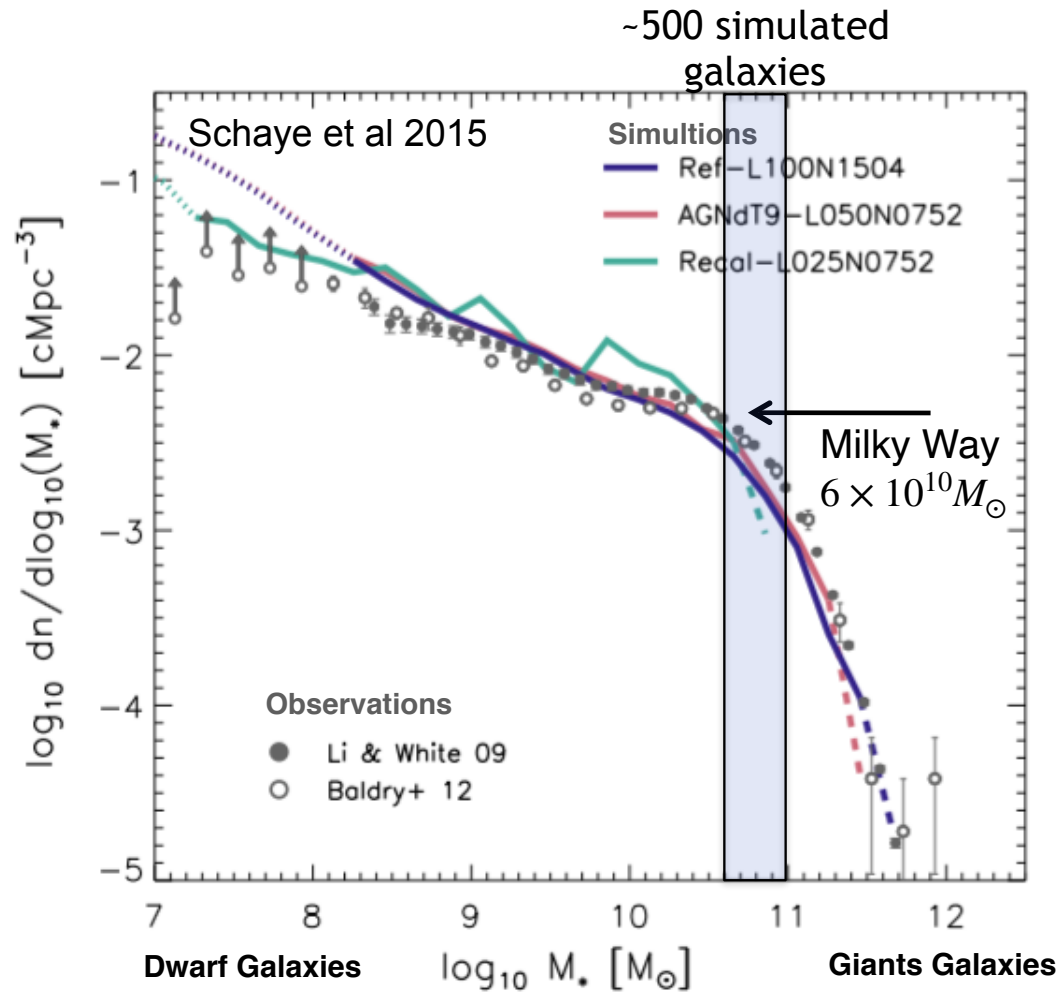
The galaxy number density agrees with the data to  $< \sim 0.2$  dex.

High-mass end fewer than 10 objects per (0.2 dex) stellar mass bin.

Low-mass end stellar mass falls below 100 baryonic particles.

GAMA survey ( $z < 0.06$ ; Baldry et al. 2012)  
SDSS ( $z \sim 0.07$ ; Li & White 2009).

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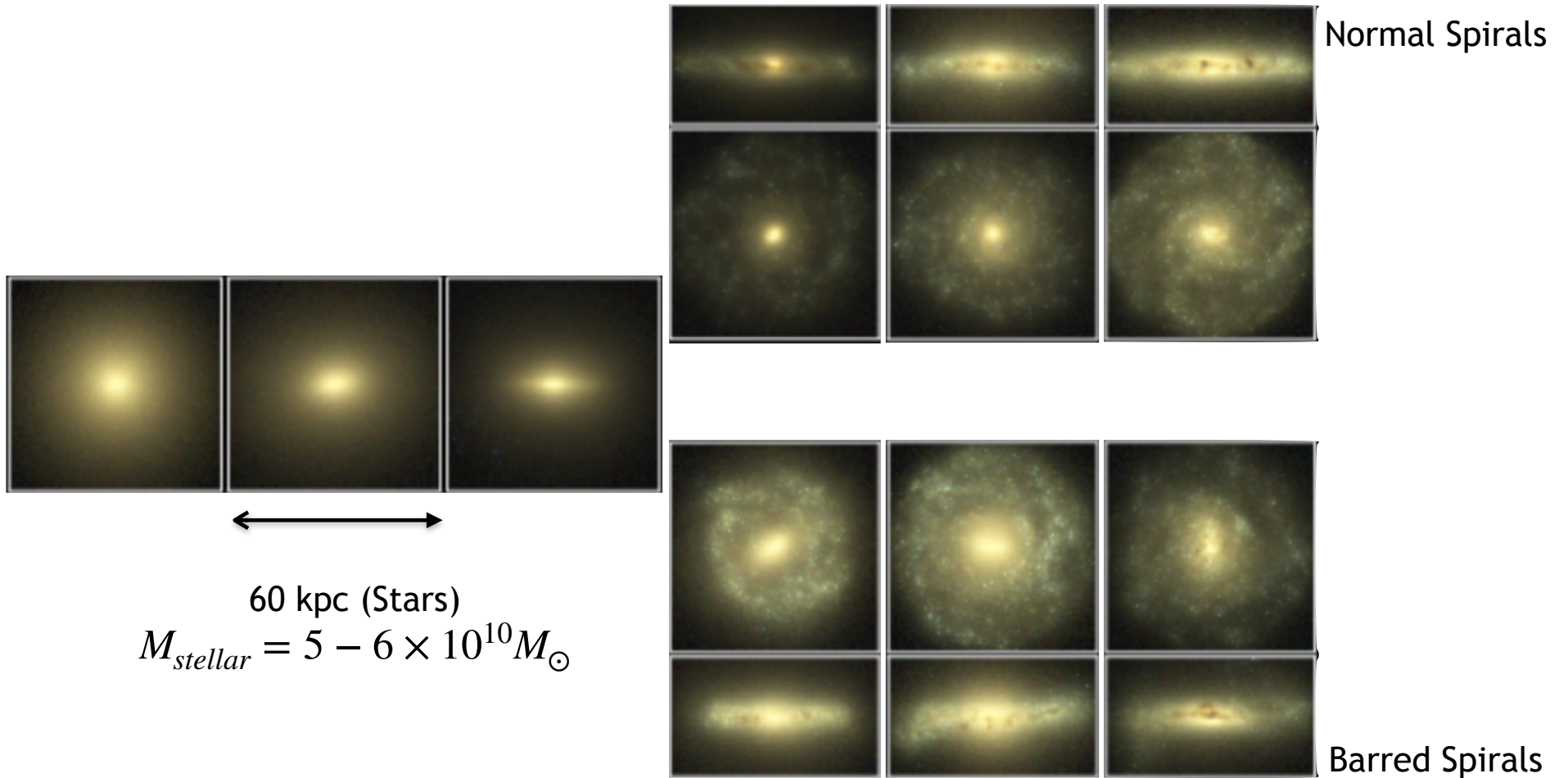
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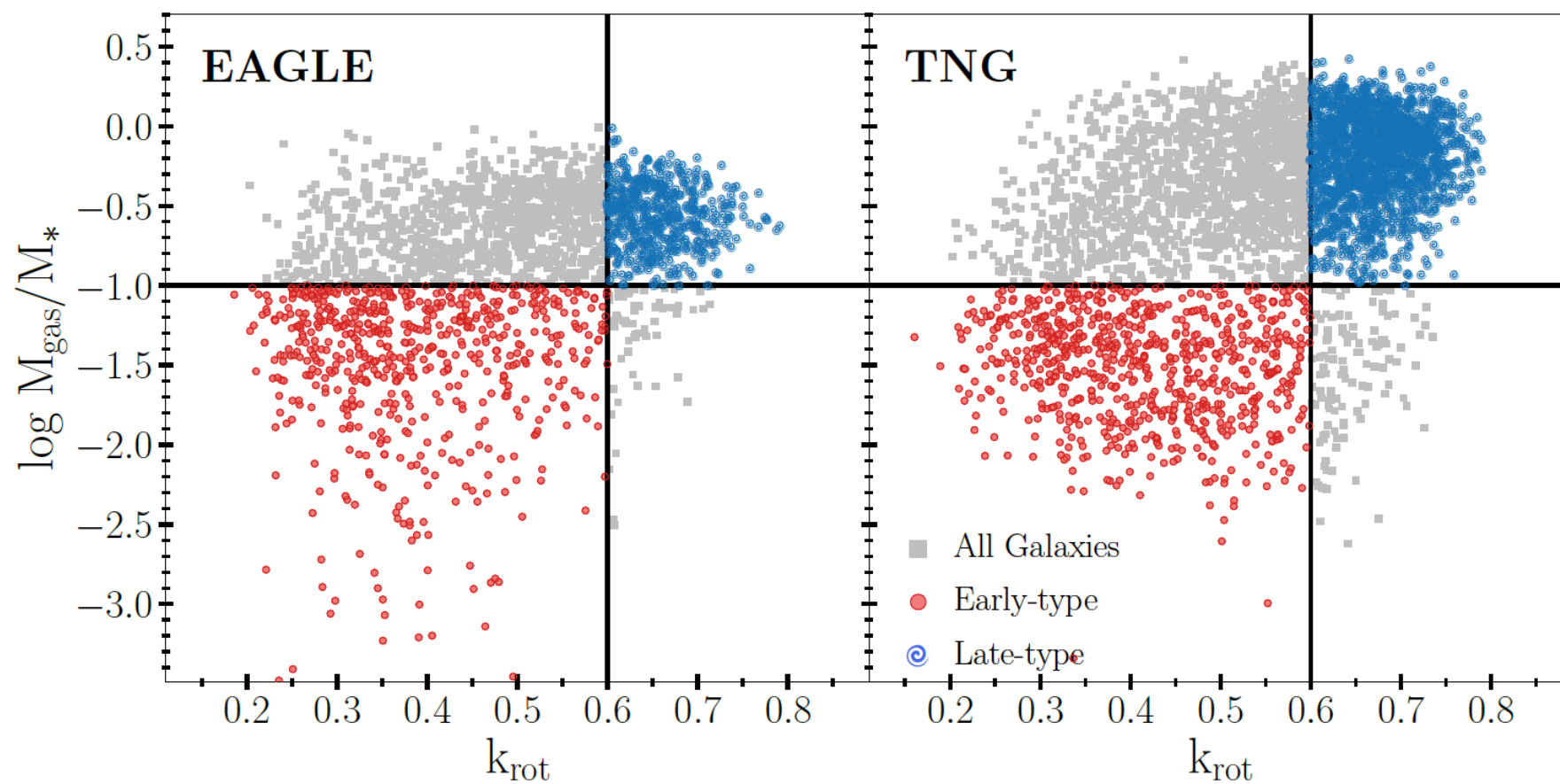
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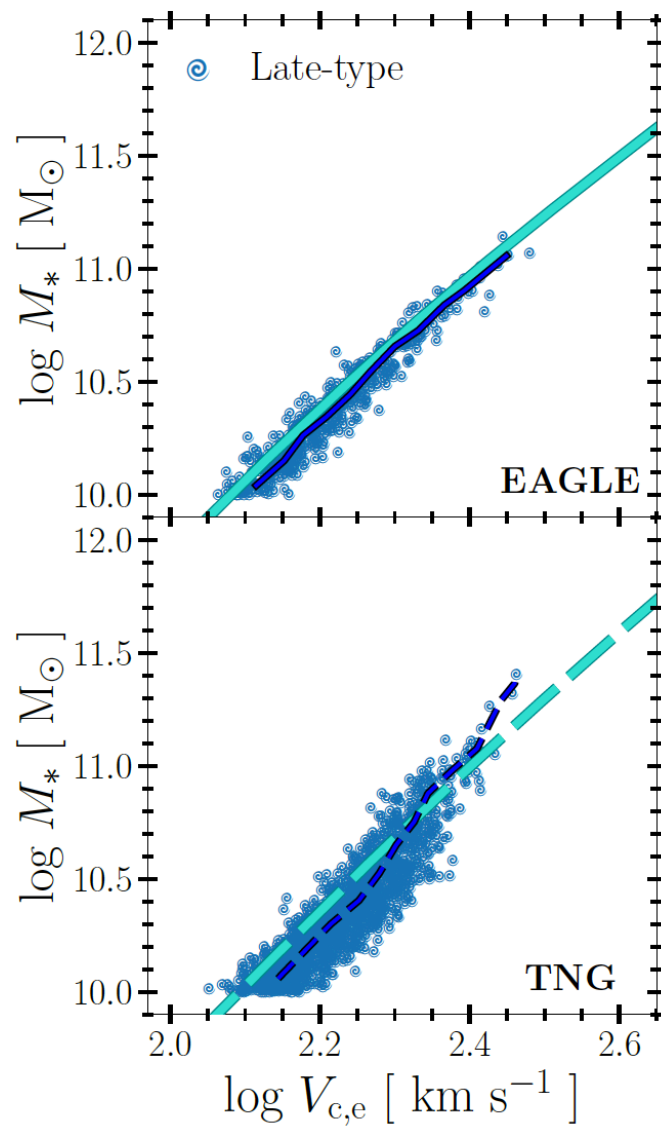
# EAGLE Morphological Classification



# Spiral & Elliptical Classification

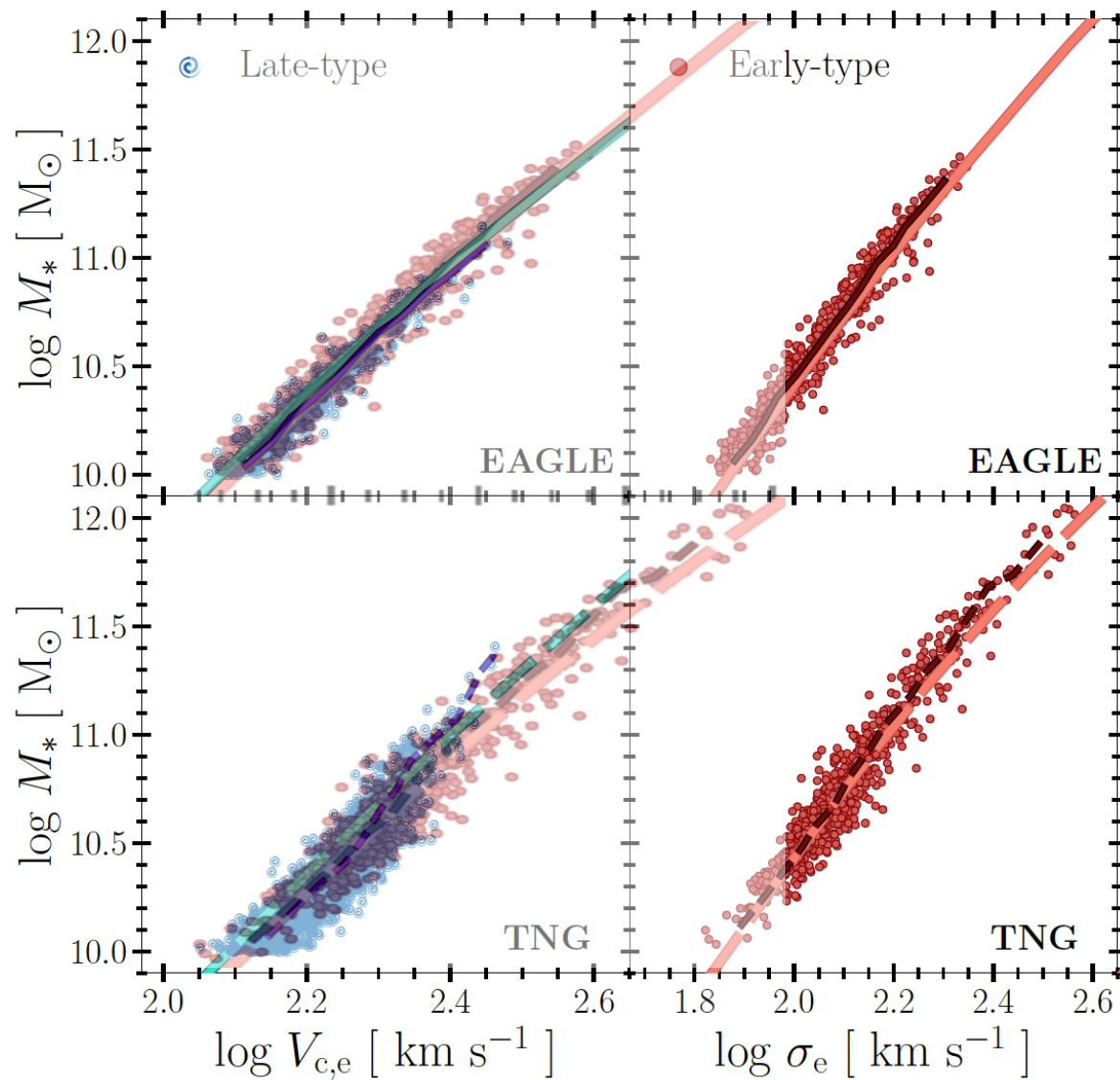


# Simulated Scaling Relations

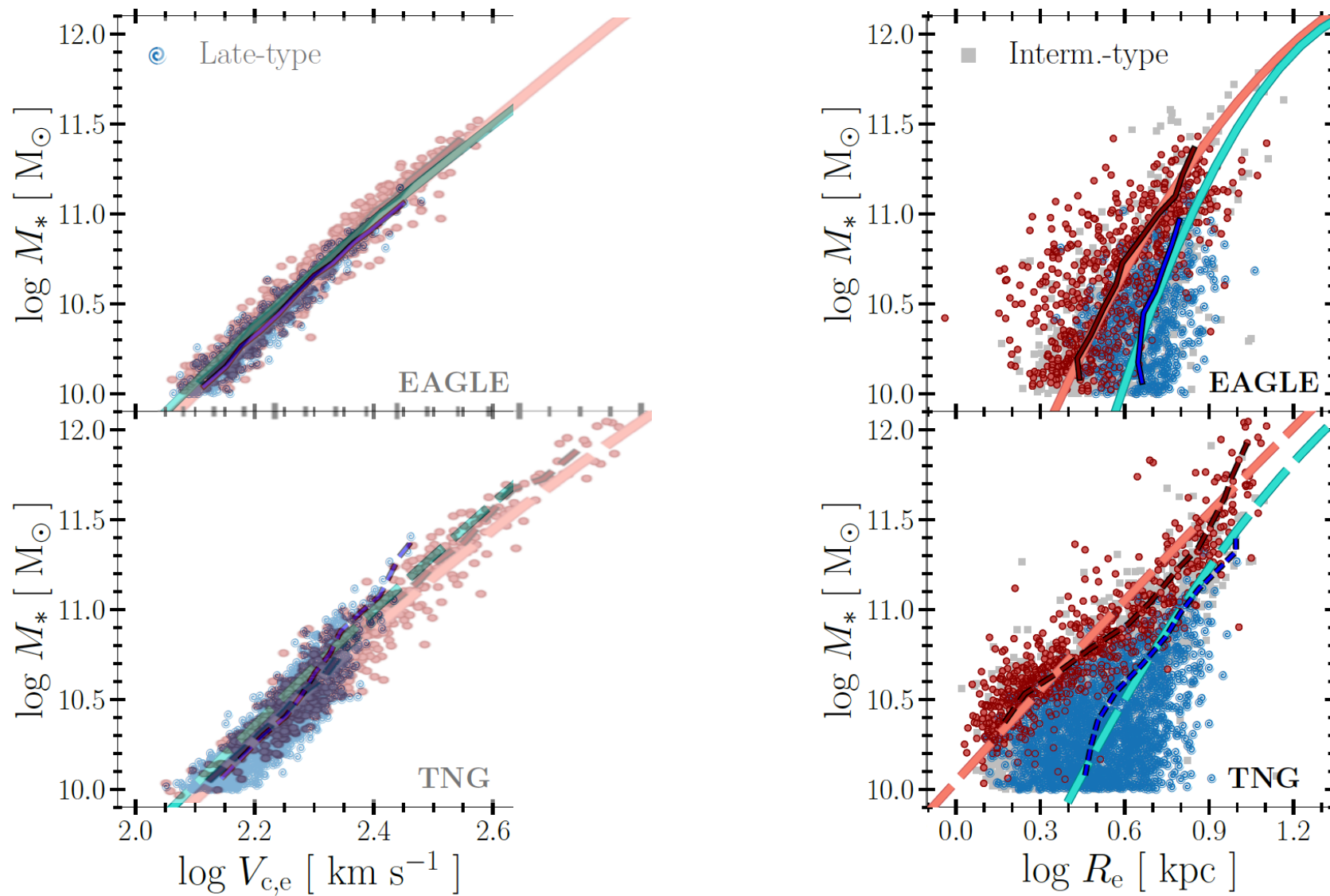




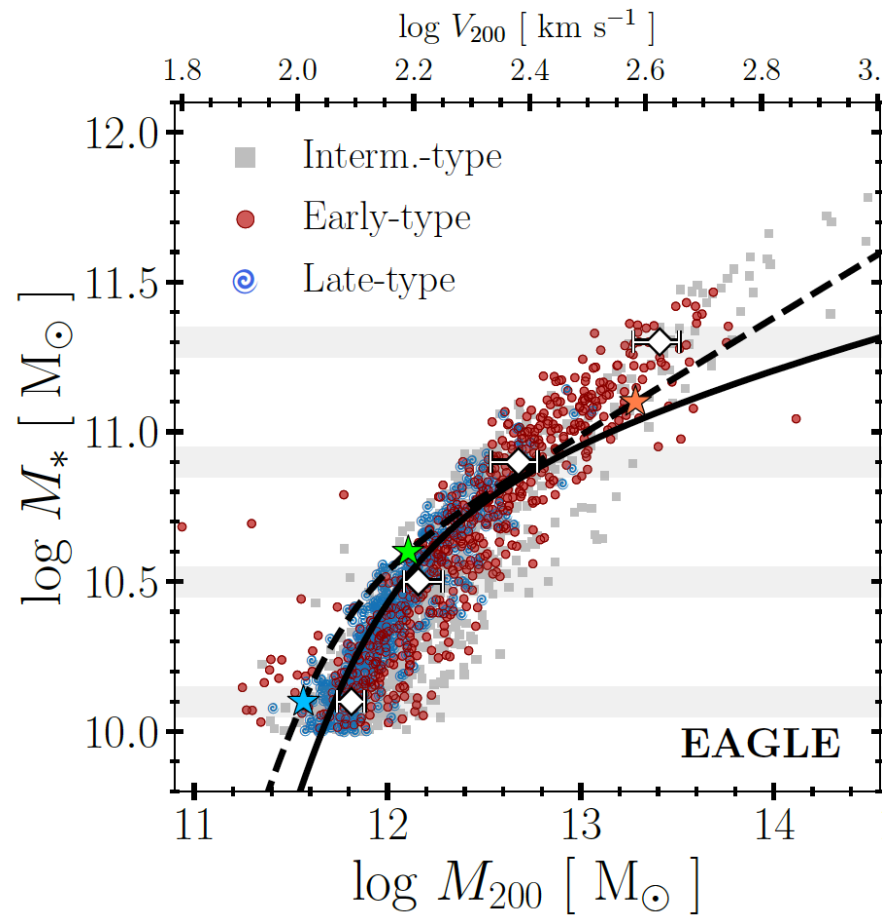
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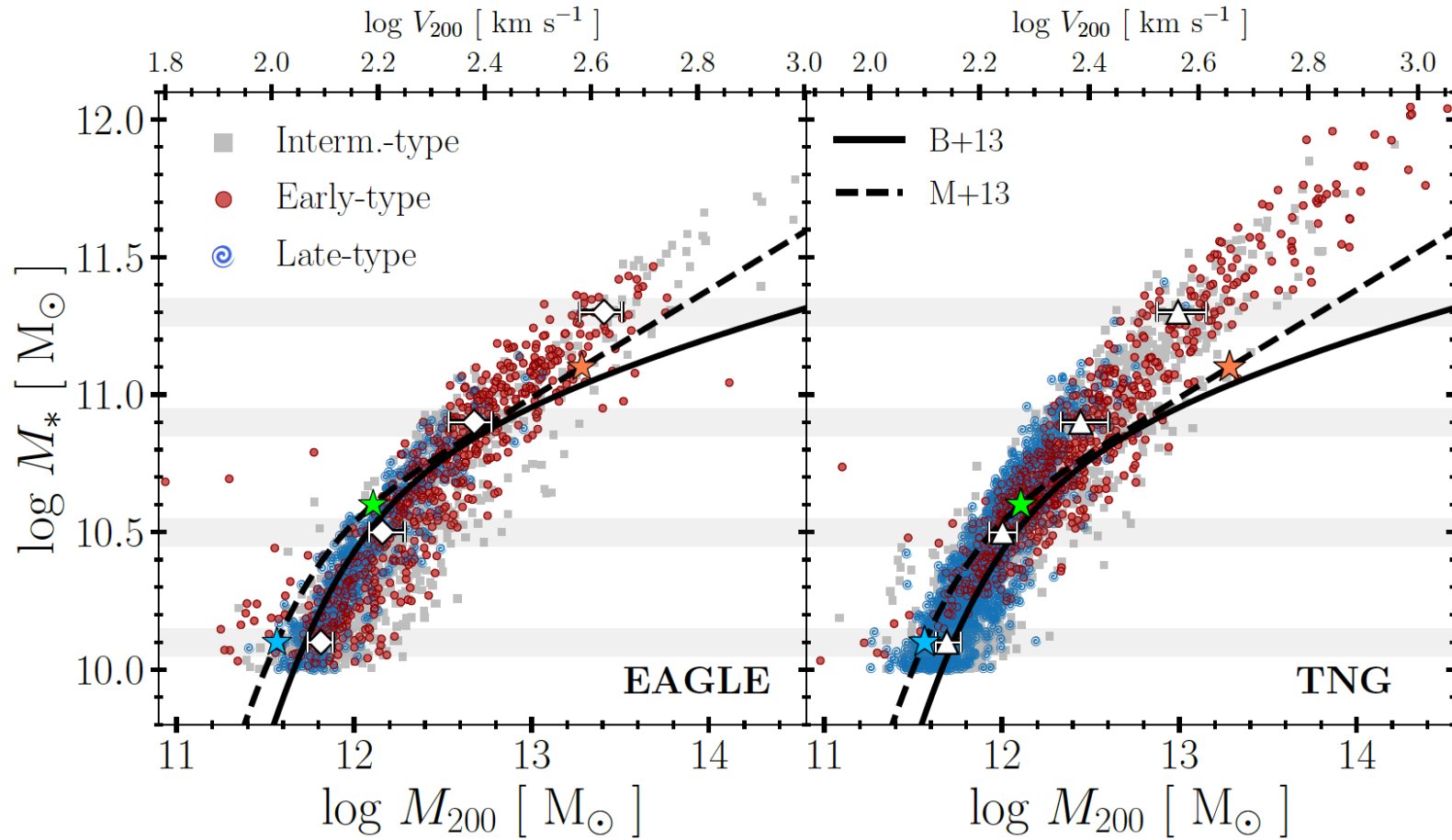
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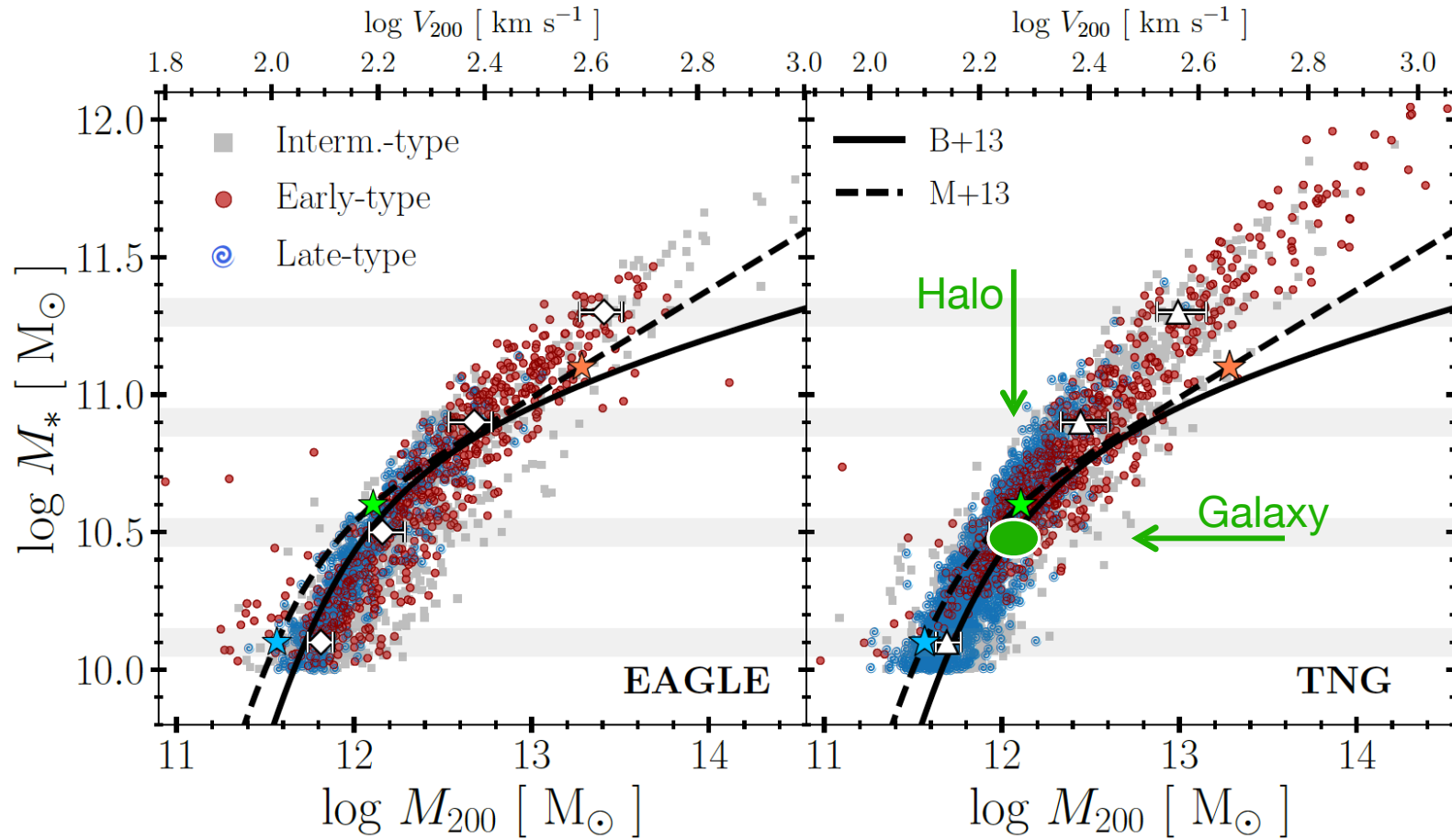
# Stellar Halo Mass Relation



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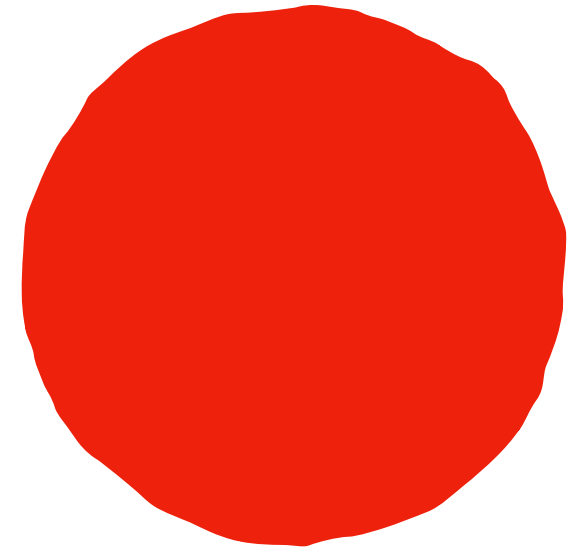


# Stellar Halo Mass Relation



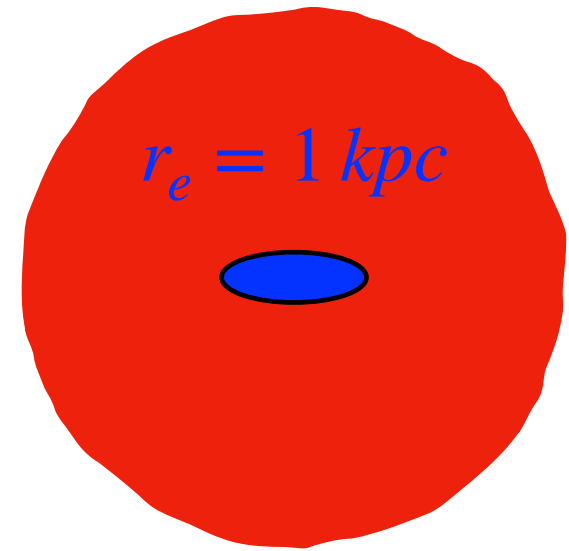
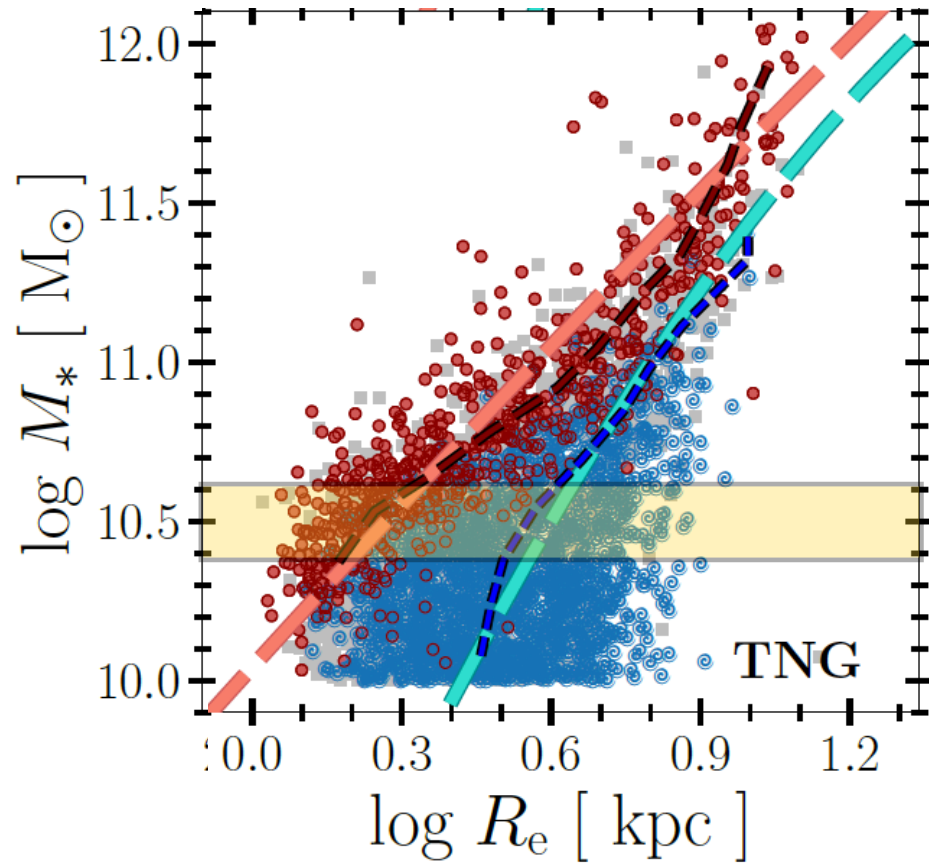


# Theoretical Model



Dark matter halo  
 $10^{12} M_{\odot}$

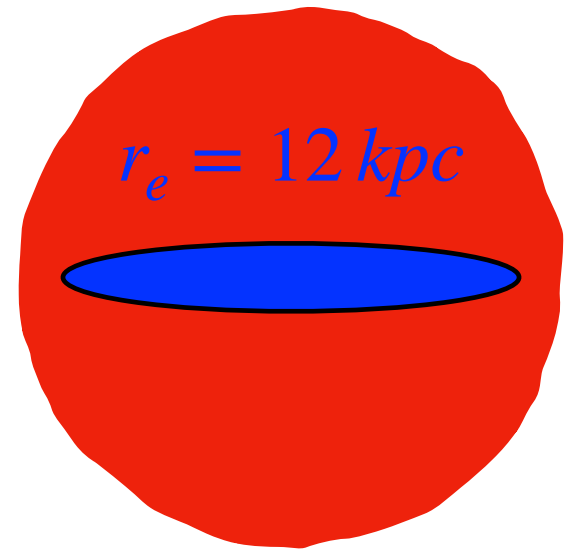
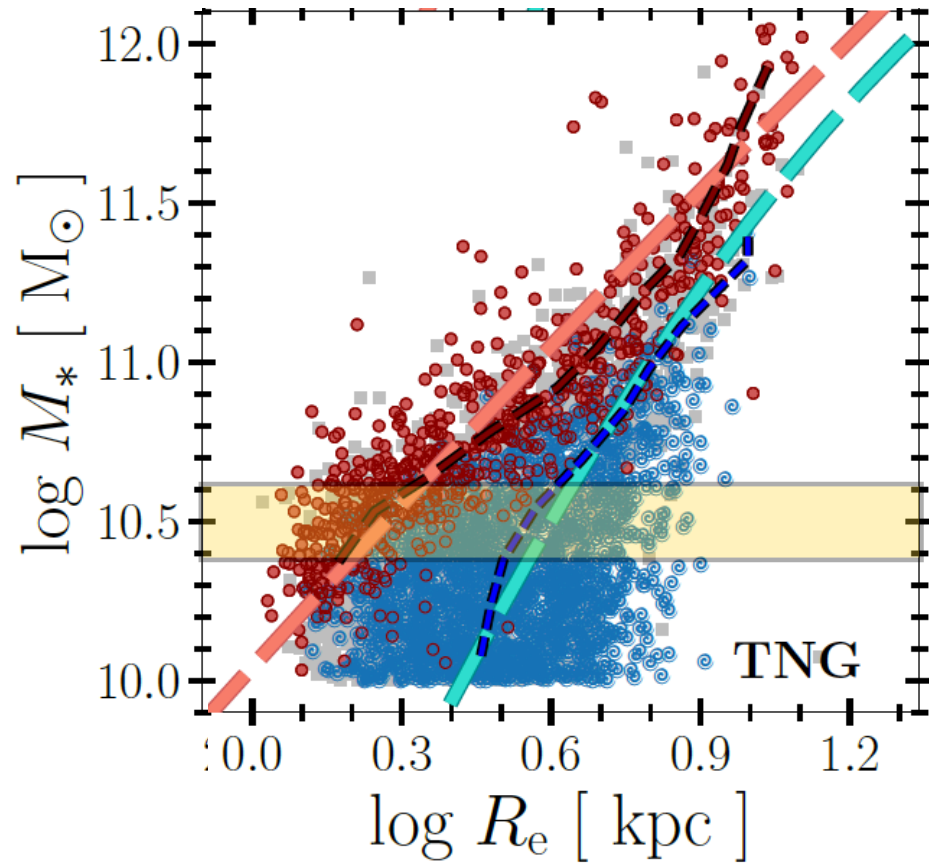
# Theoretical Model



Dark matter halo  
 $10^{12} M_\odot$

**COMPACT** Galaxy  
 $10^{10.5} M_\odot$

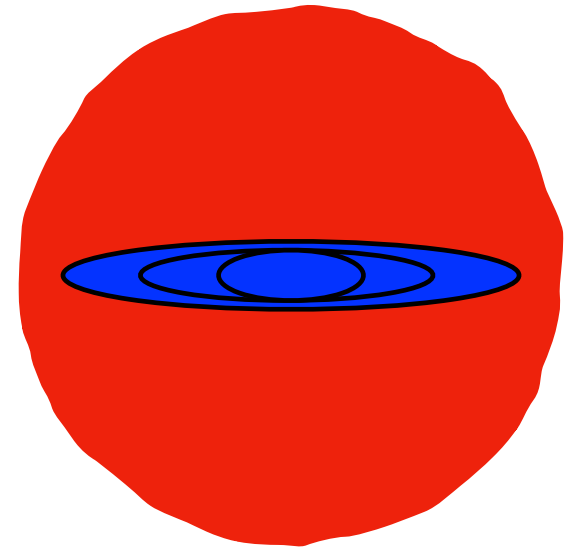
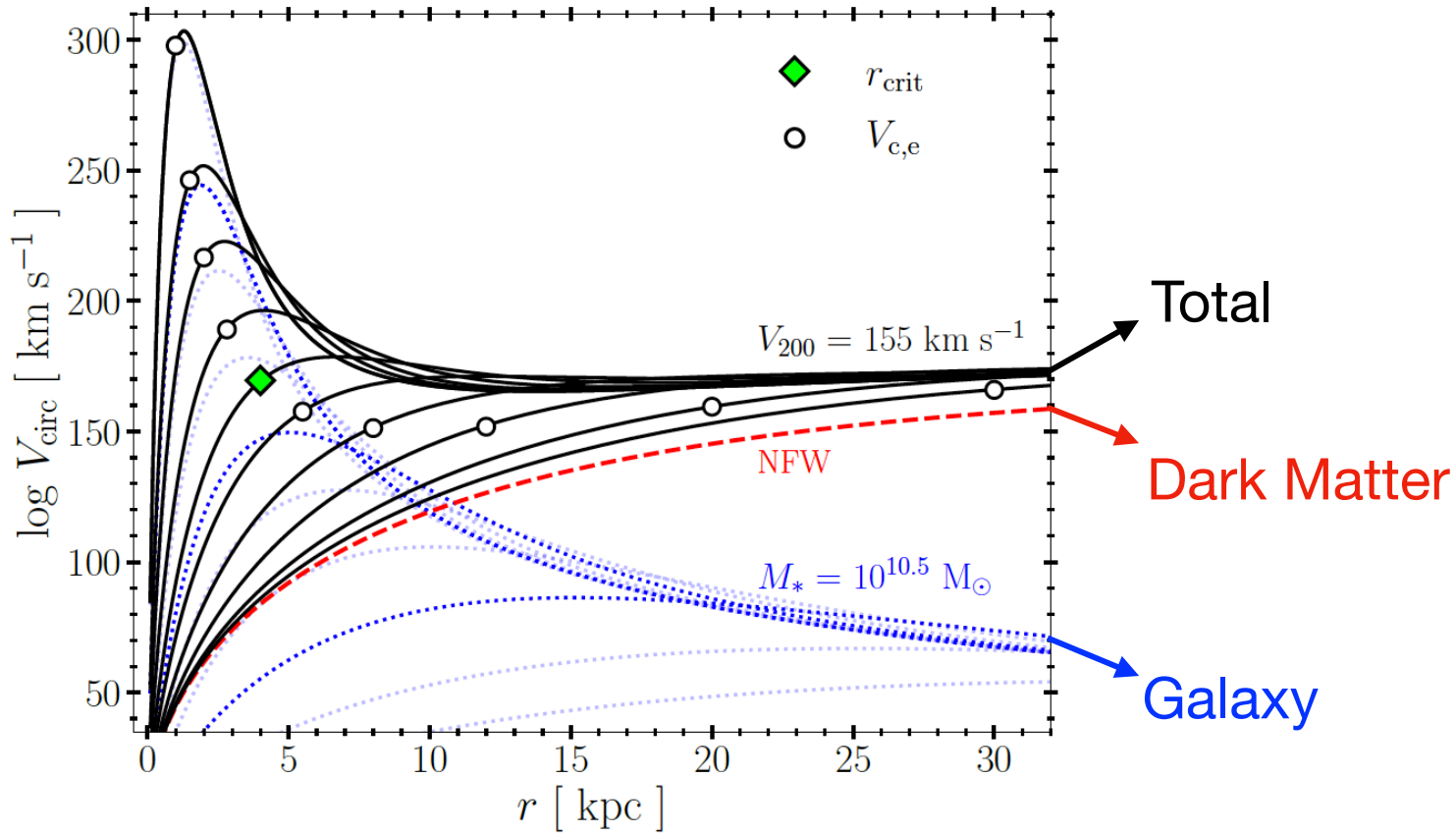
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Dark matter halo  
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**EXTENDED** Galaxy  
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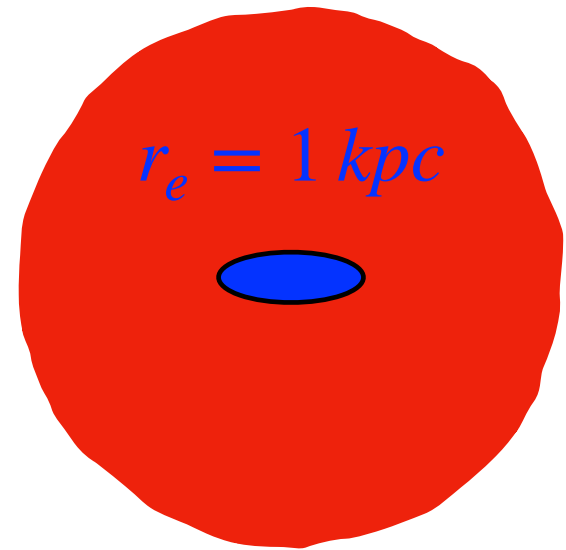
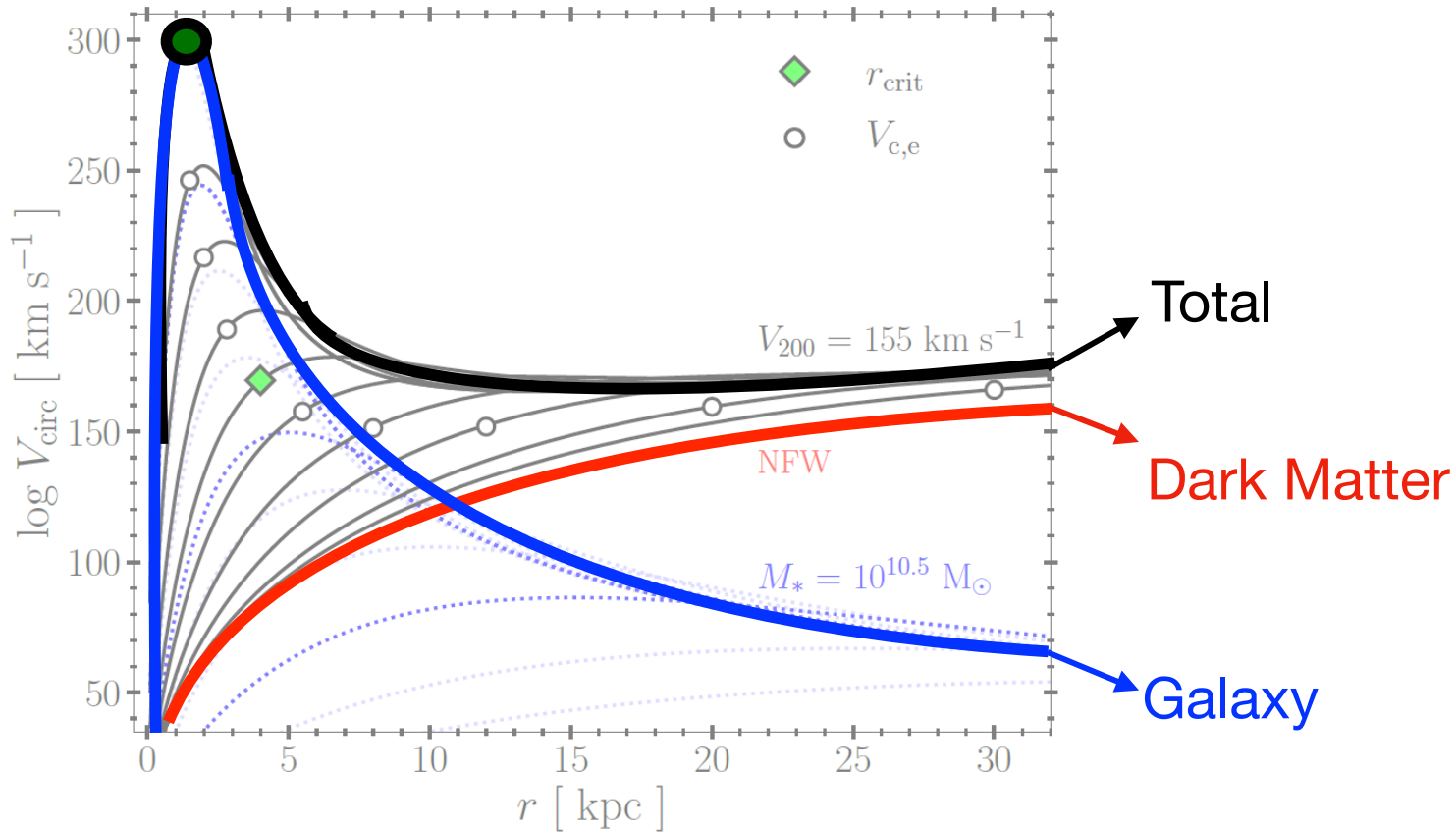
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Dark matter halo  
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Galaxy  
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# Theoretical Model

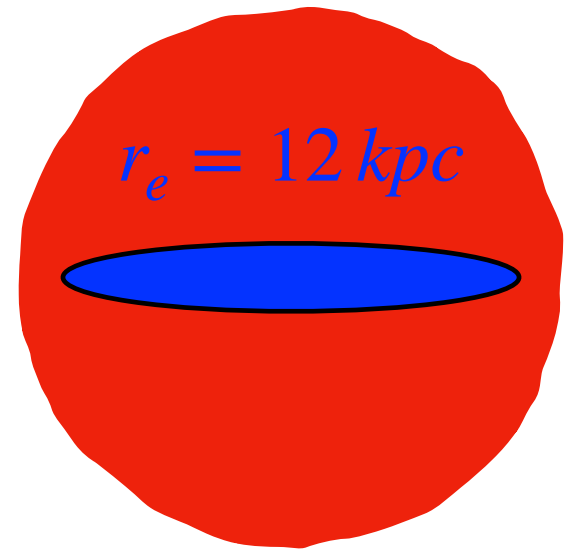
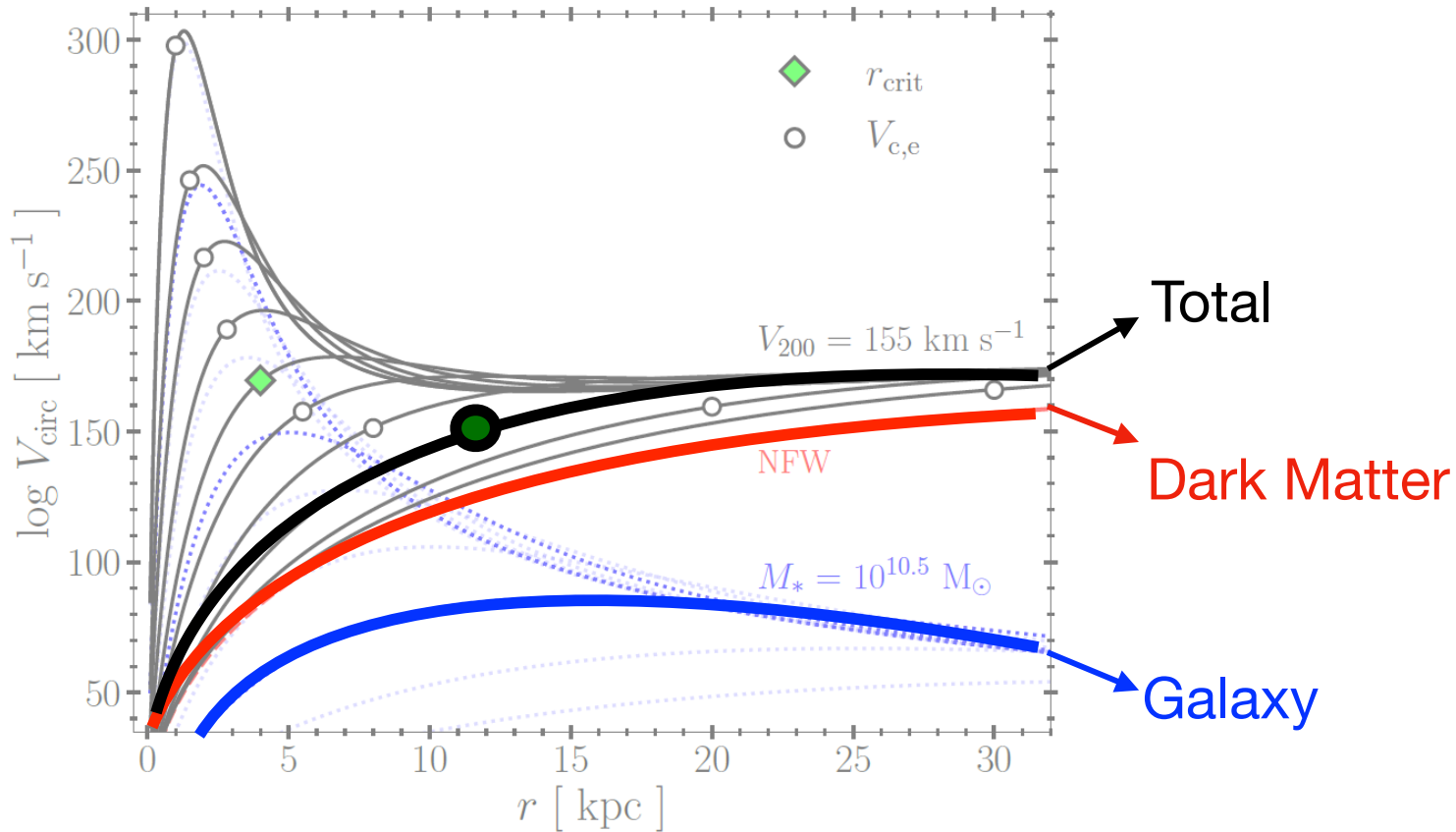


Dark matter halo  
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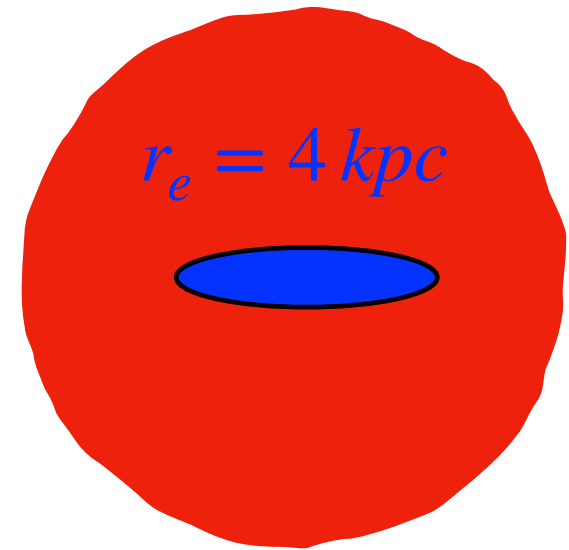
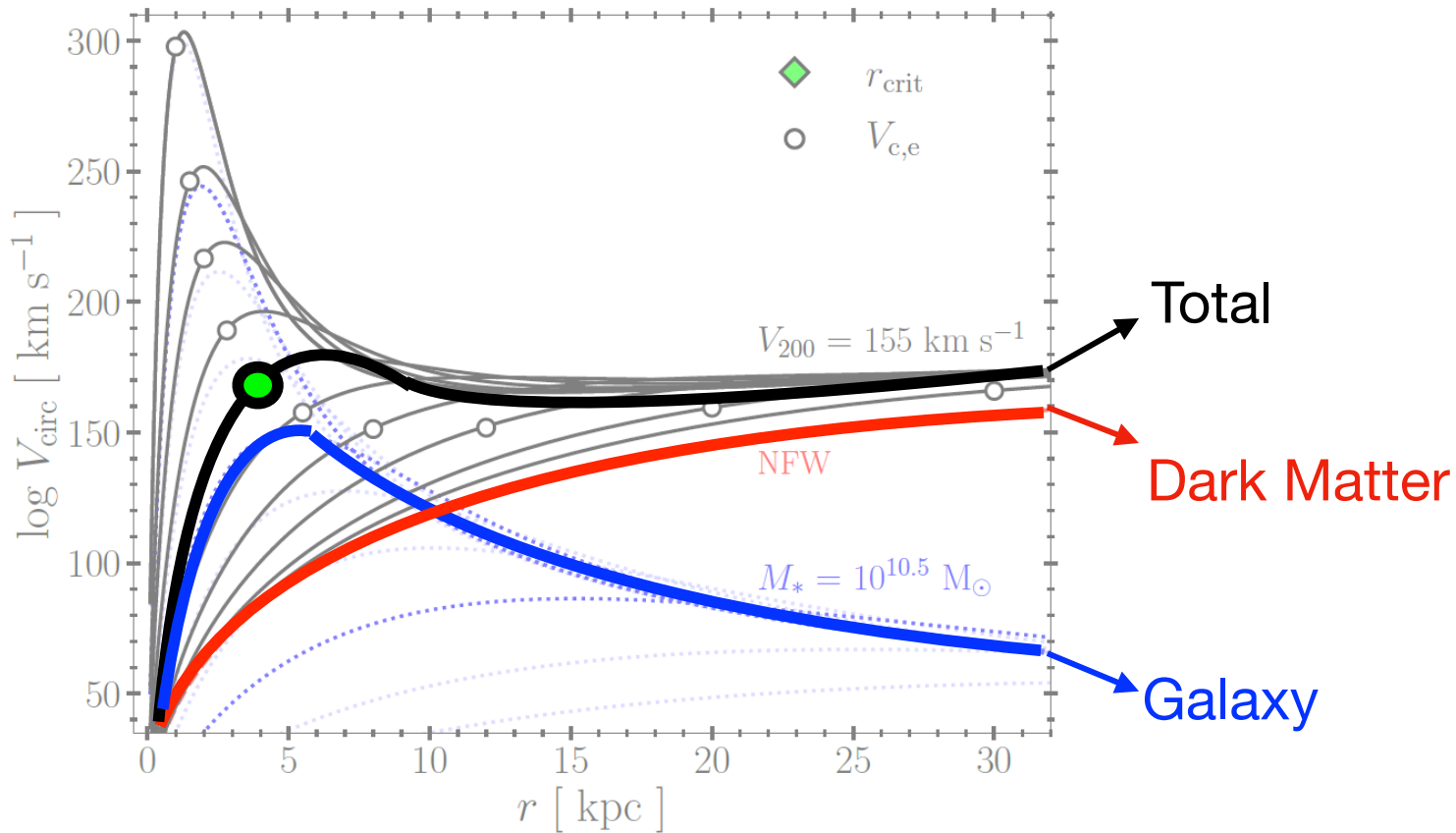
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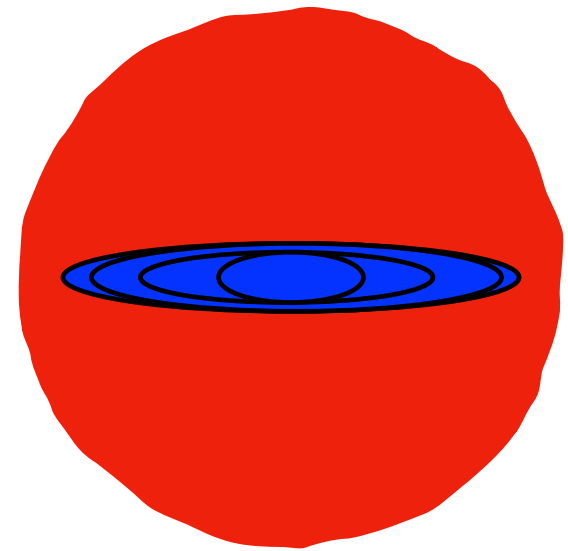
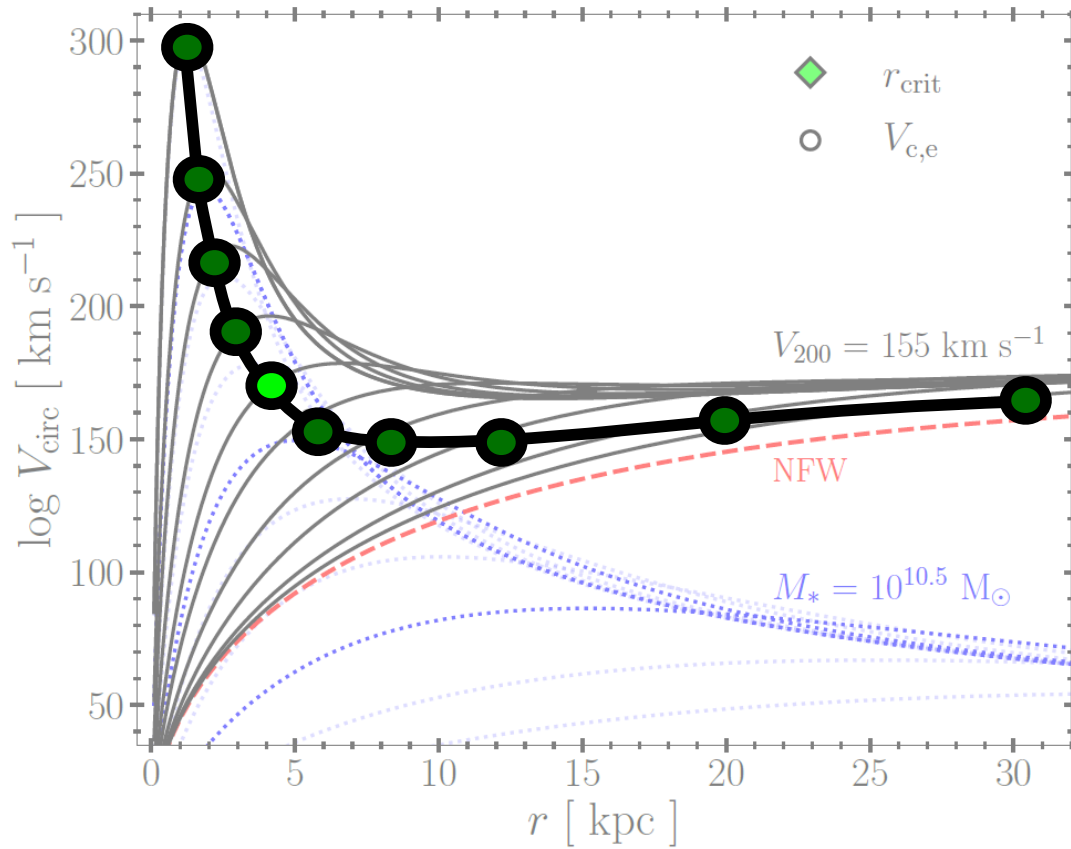
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Dark matter halo  
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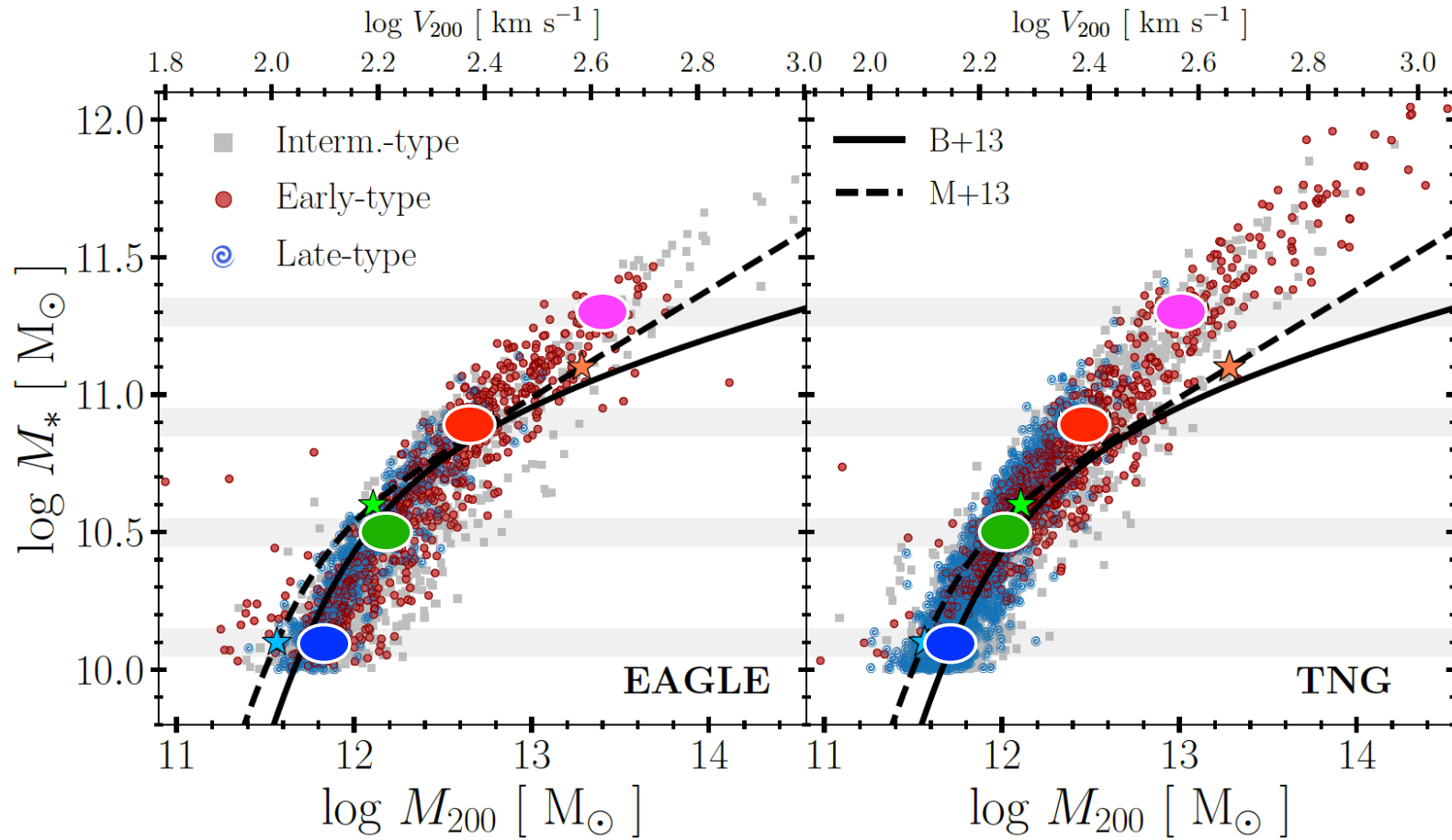
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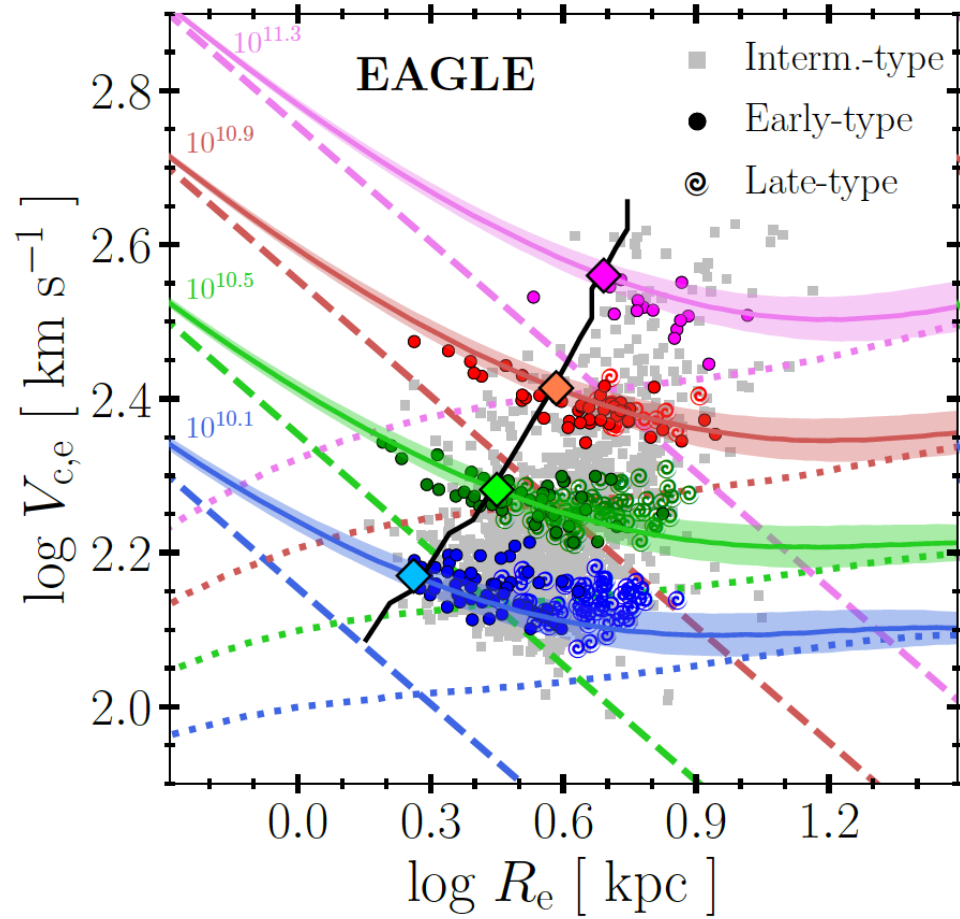
Dark matter halo  
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Galaxy  
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# Stellar Halo Mass Relation

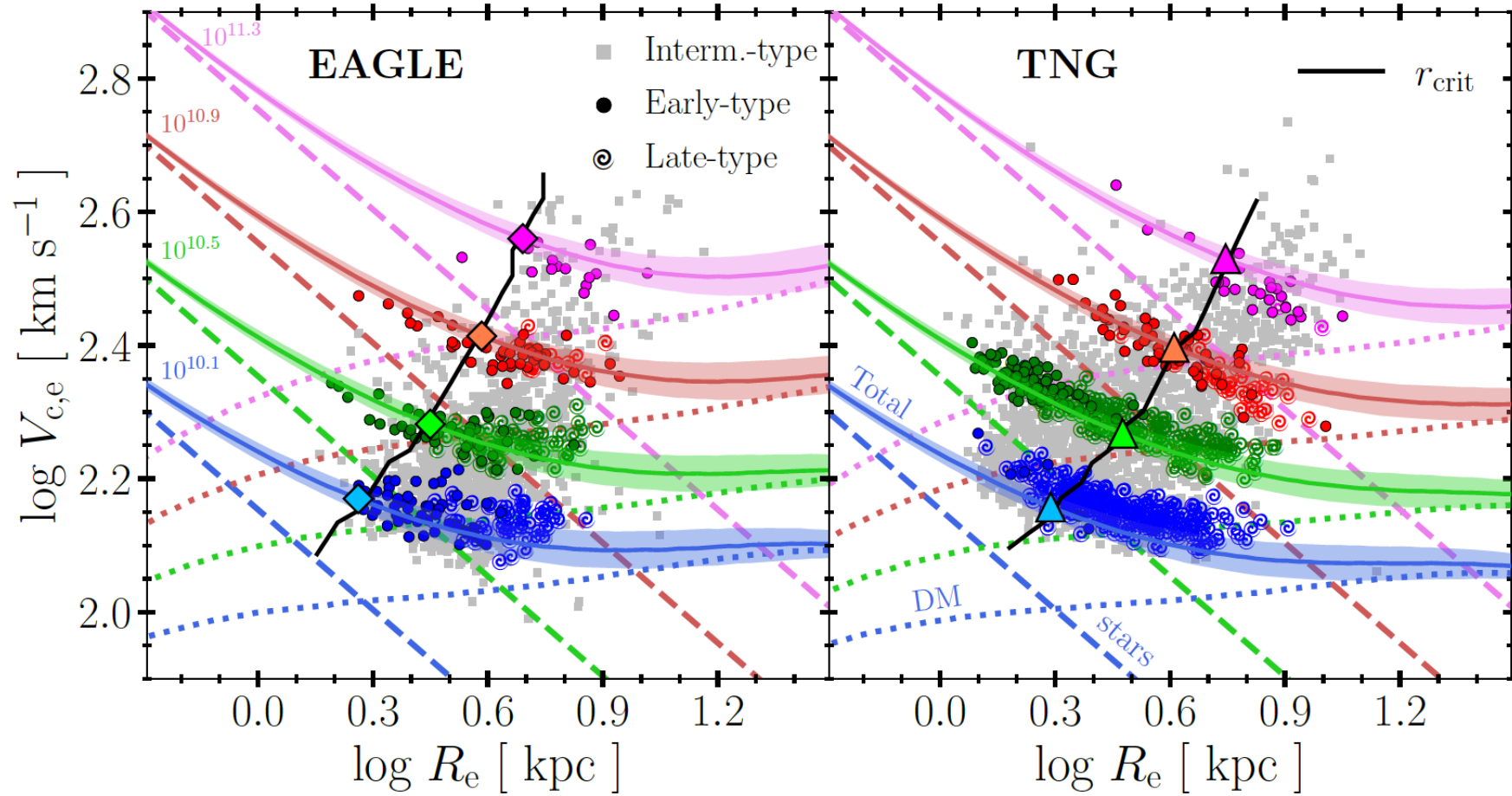


# Velocity-Radius Relation





# Velocity-Radius Relation

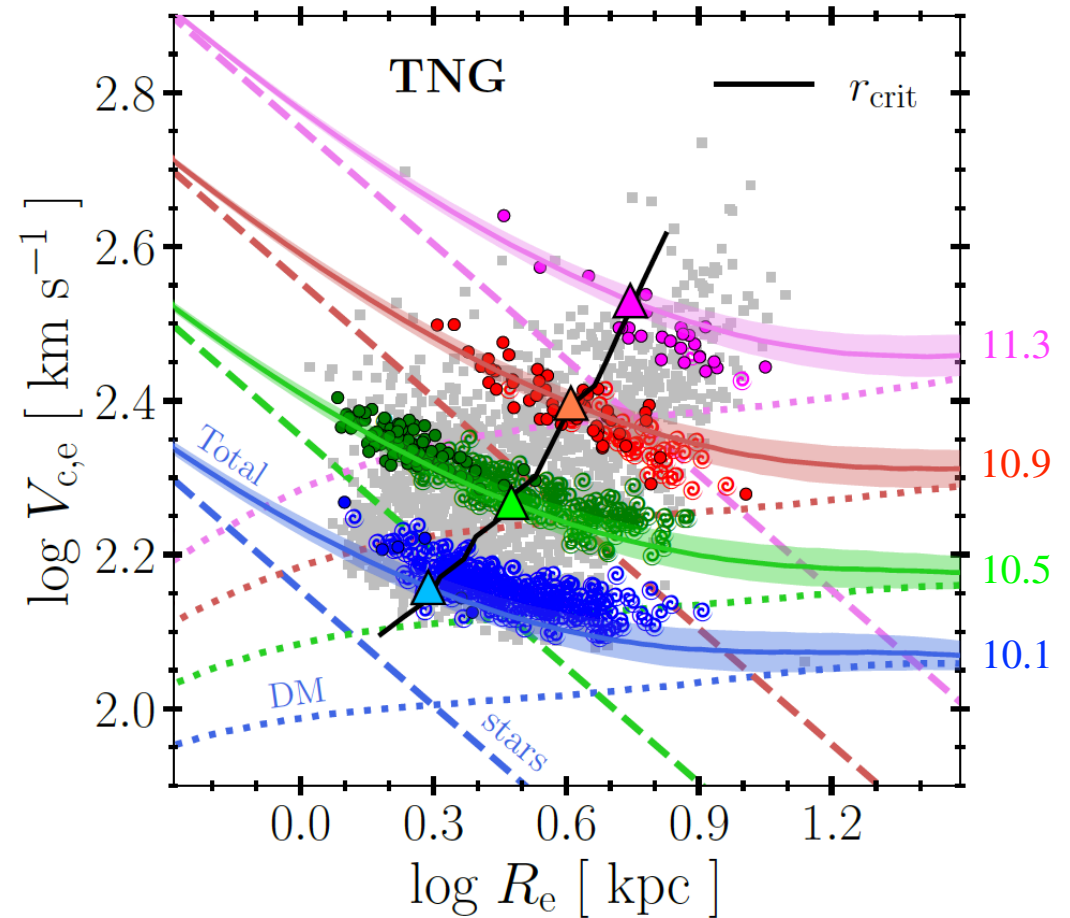


# Velocity-Radius Relation

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.2}{0.6}$$

$$M \sim V^2$$

0.6



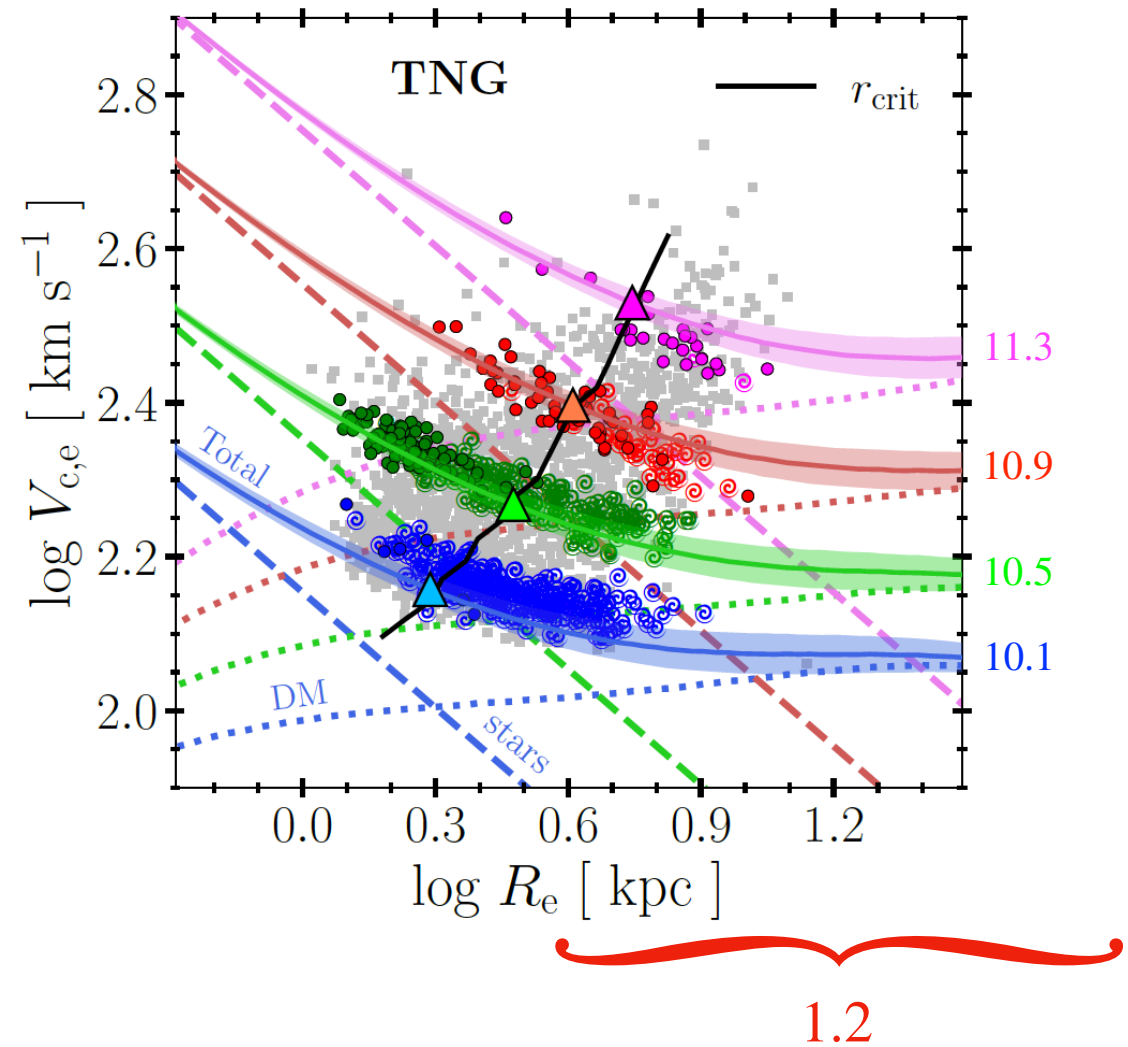
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$$M \sim R$$



# Velocity-Radius Relation

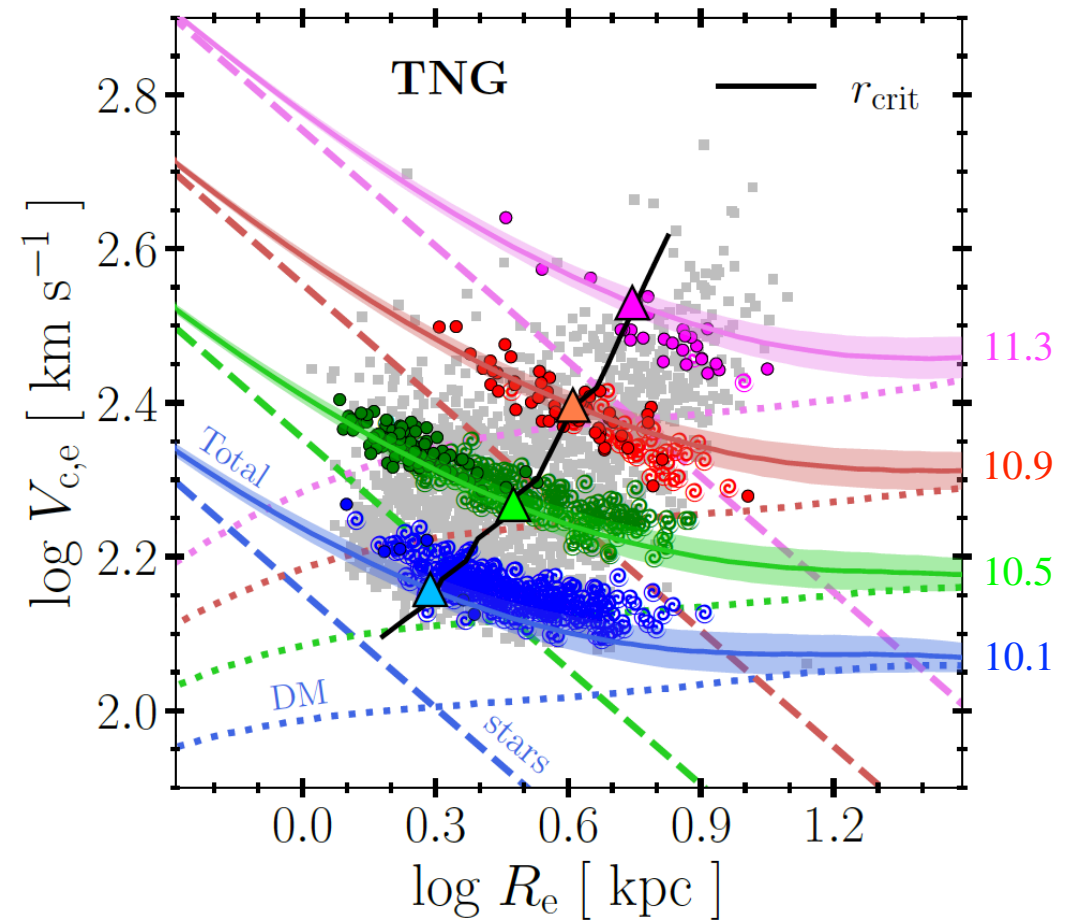
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$$M \sim R$$

$M \sim V^2 R$  Fundamental Plane



# Velocity-Radius Relation

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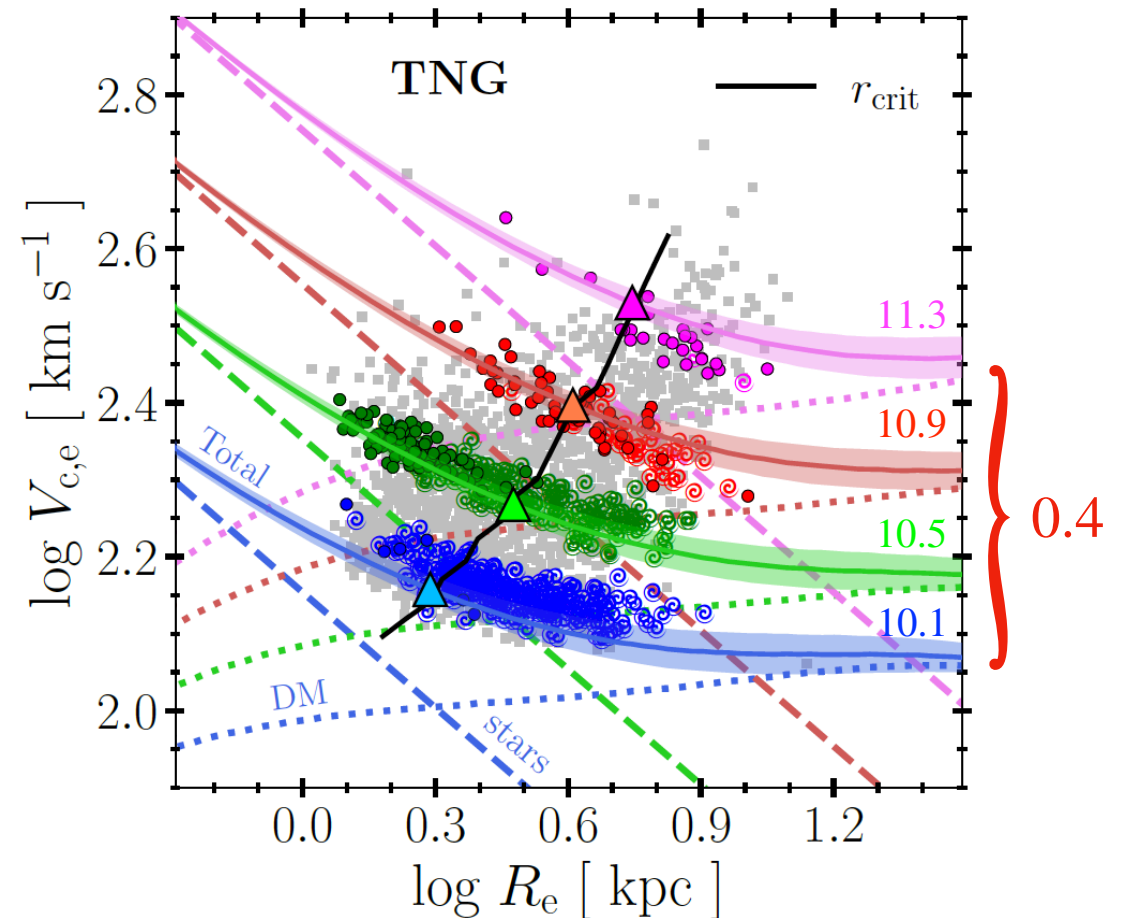
$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.2}{1.2}$$

$$M \sim R$$

$$M \sim V^2 R \quad \text{Fundamental Plane}$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.2}{0.4}$$

$$M \sim V^3$$





# Velocity-Radius Relation

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.2}{0.6}$$

$$M \sim V^2$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.2}{1.2}$$

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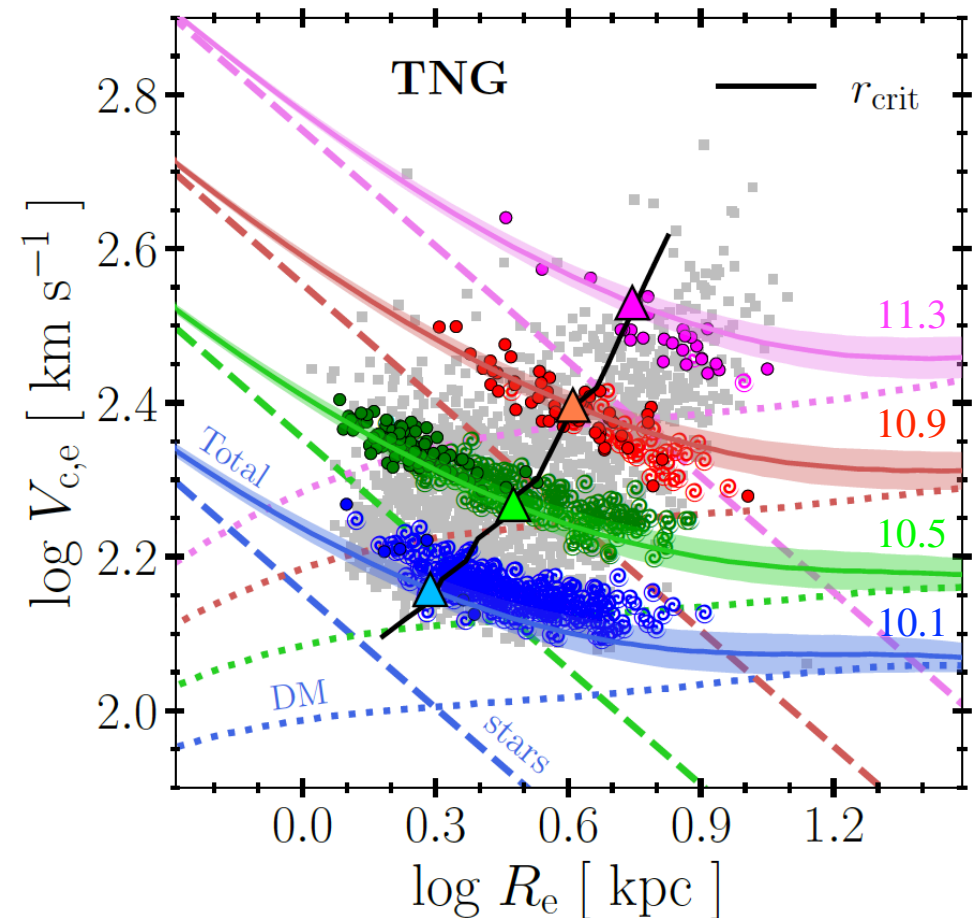
$M \sim V^2 R$  Fundamental Plane

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.2}{0.4}$$

$$M \sim V^3$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.2}{\infty}$$

$$M \sim R^0$$



# Velocity-Radius Relation

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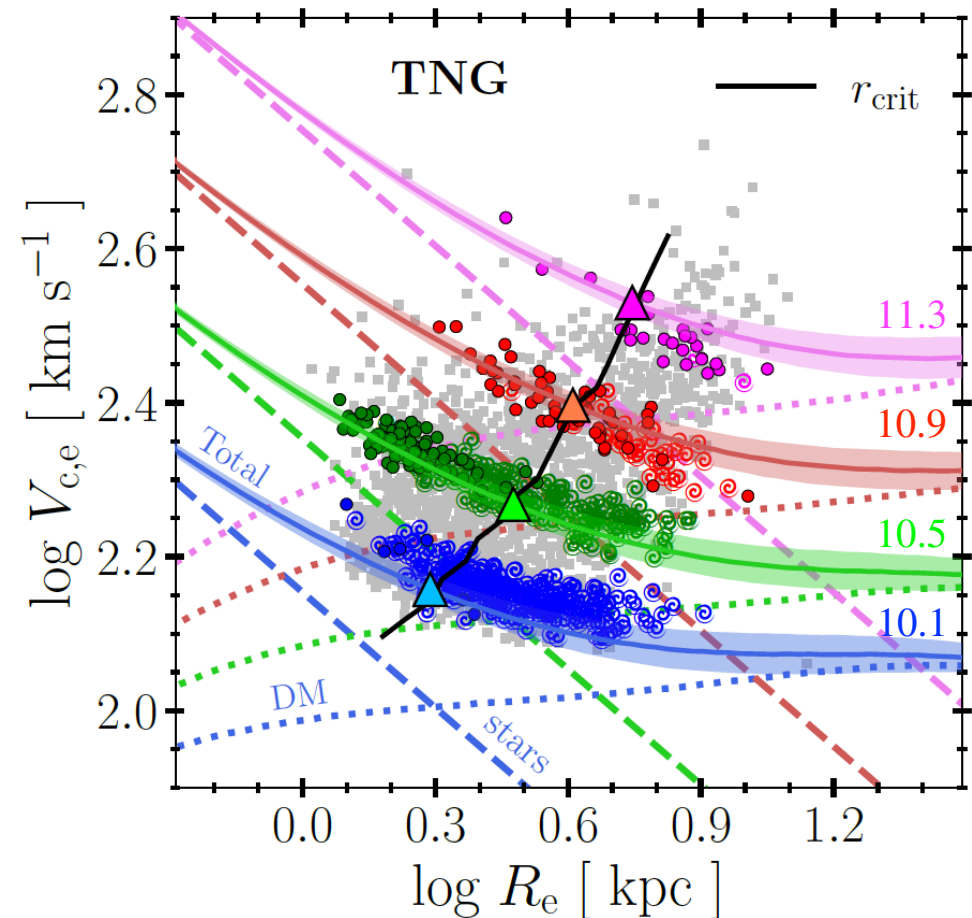
$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.2}{0.4}$$

$$M \sim V^3$$

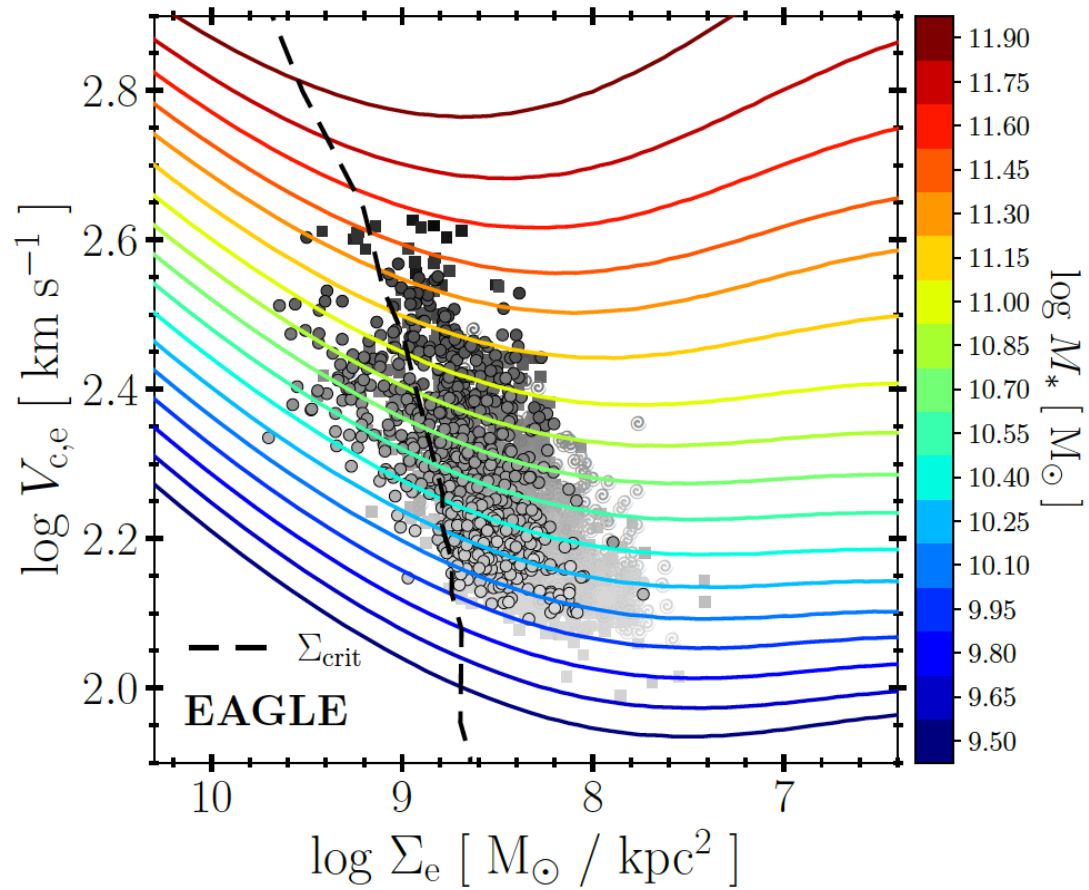
$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.2}{\infty}$$

$$M \sim R^0$$

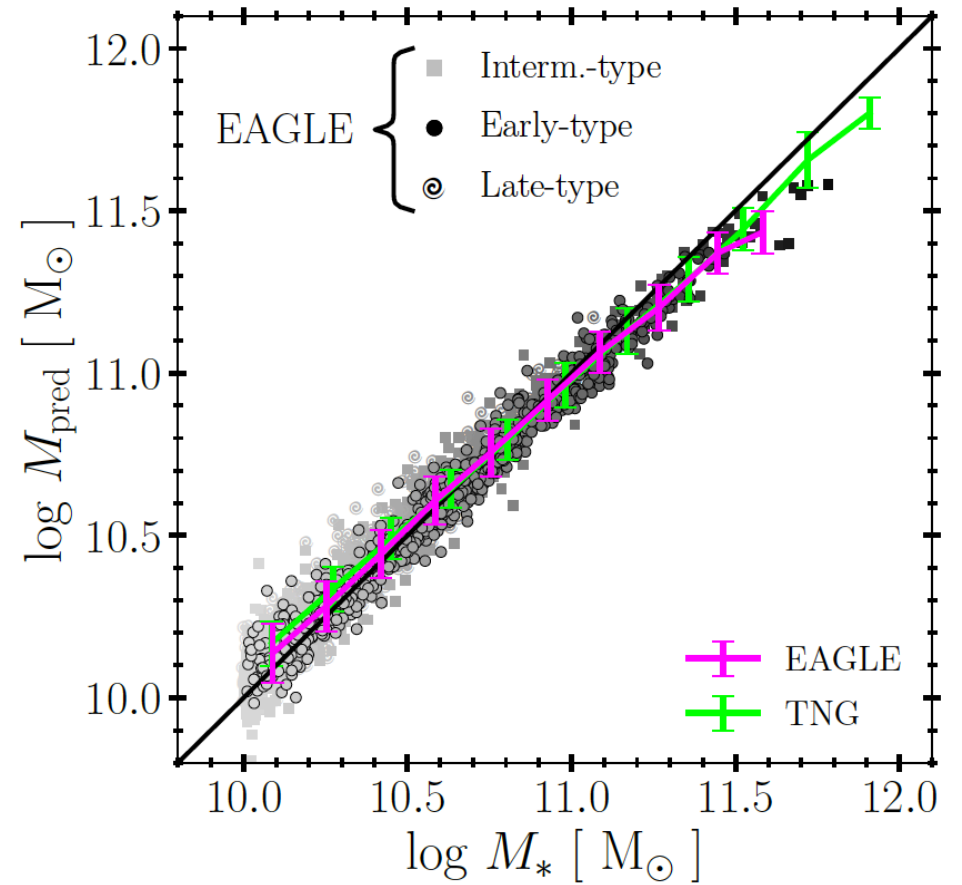
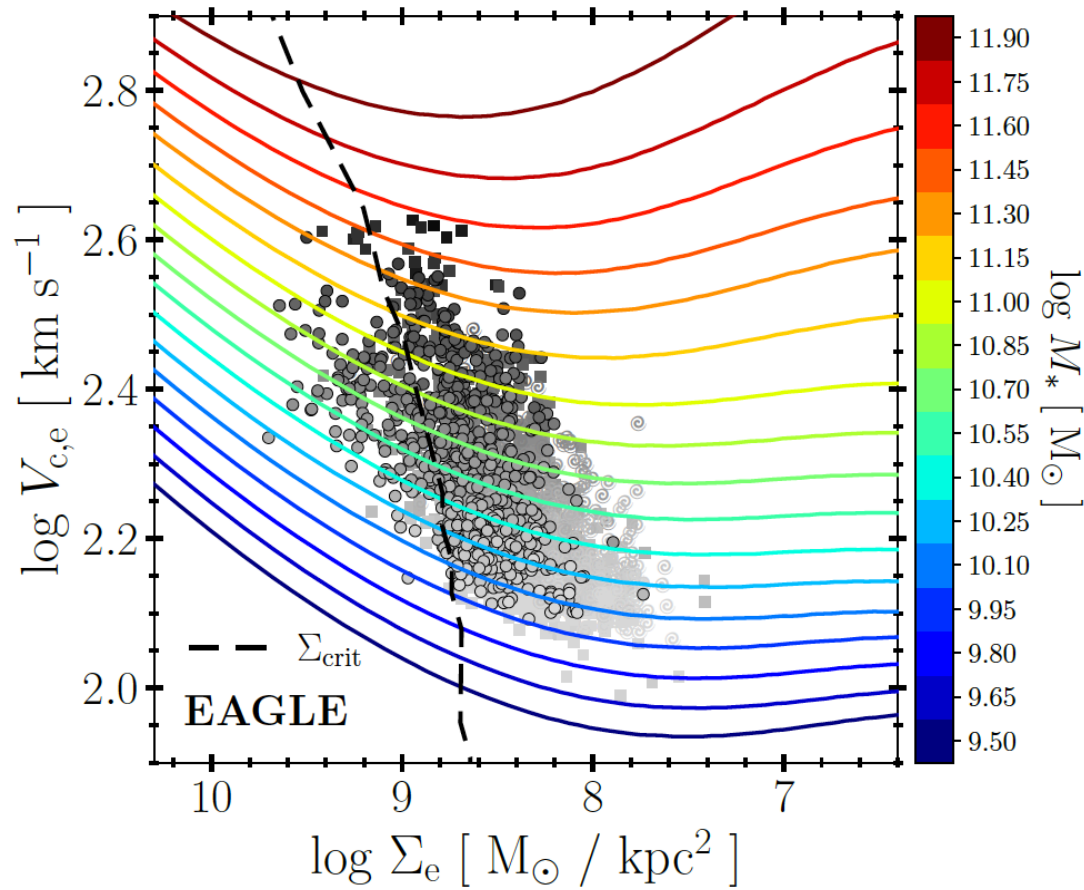
$$M \sim V^3 \quad \text{Tully-Fisher}$$



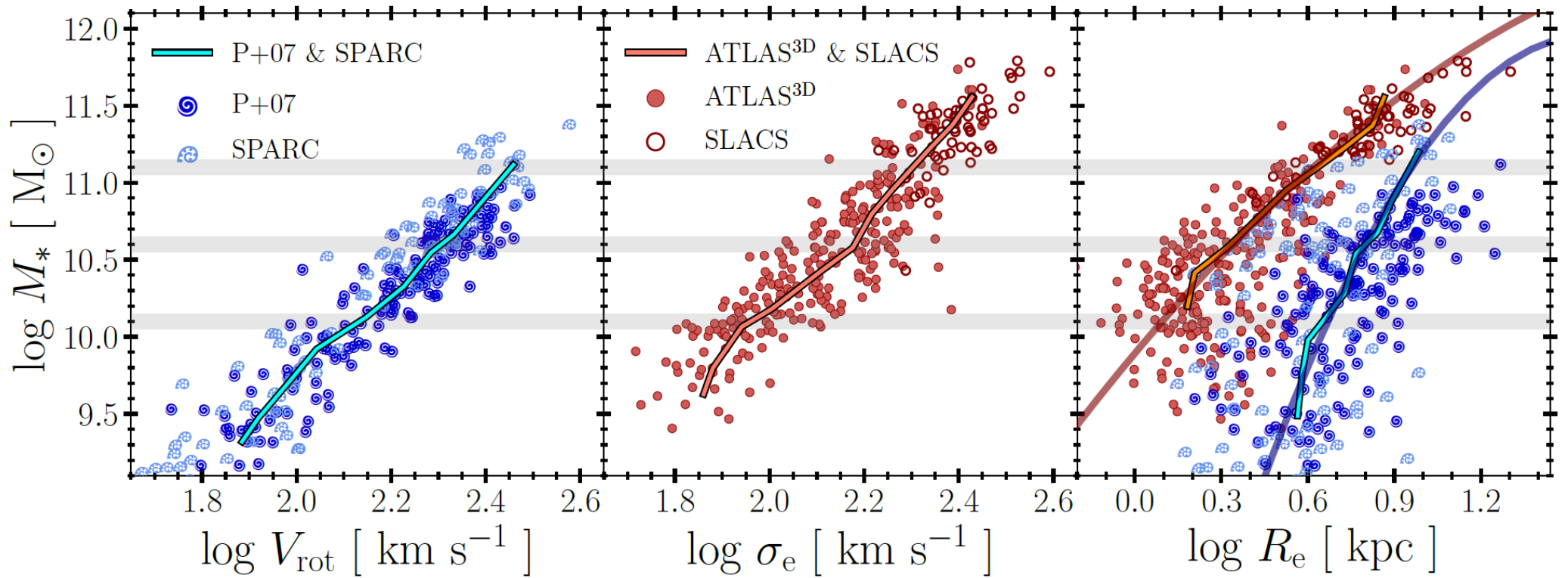
# Secondary Distance Indicator



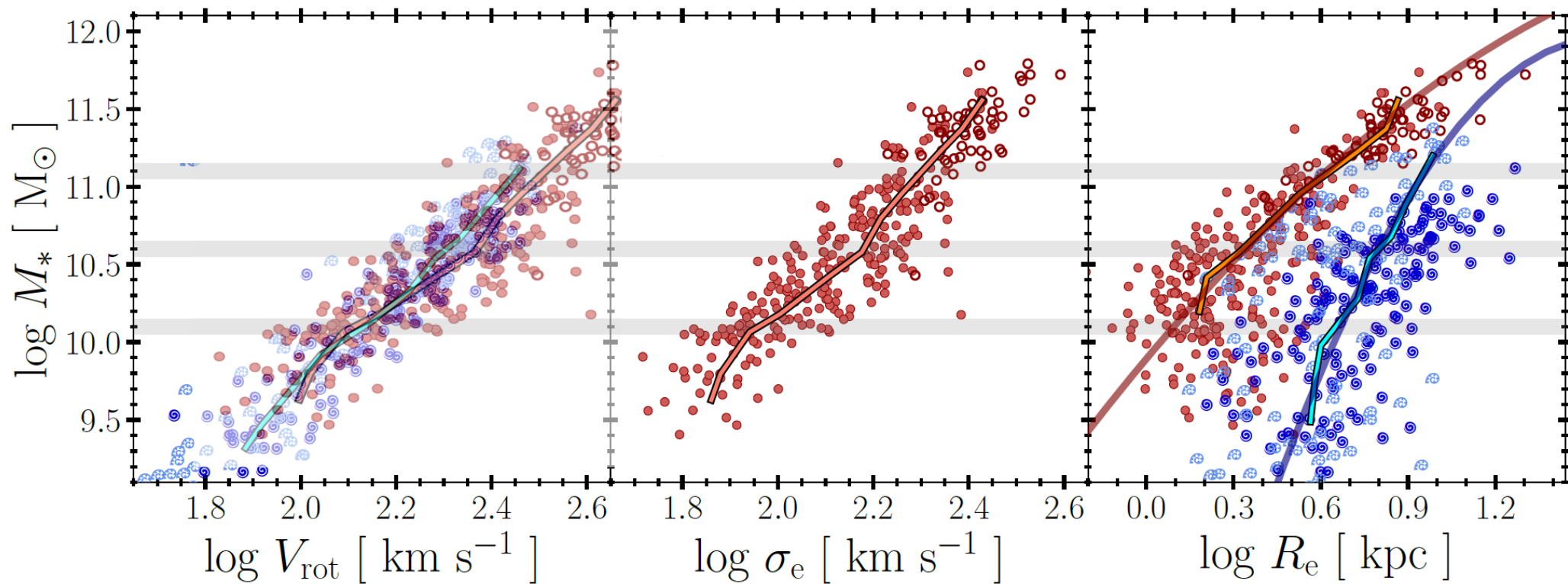
# Secondary Distance Indicator



# Observed Scaling Relations

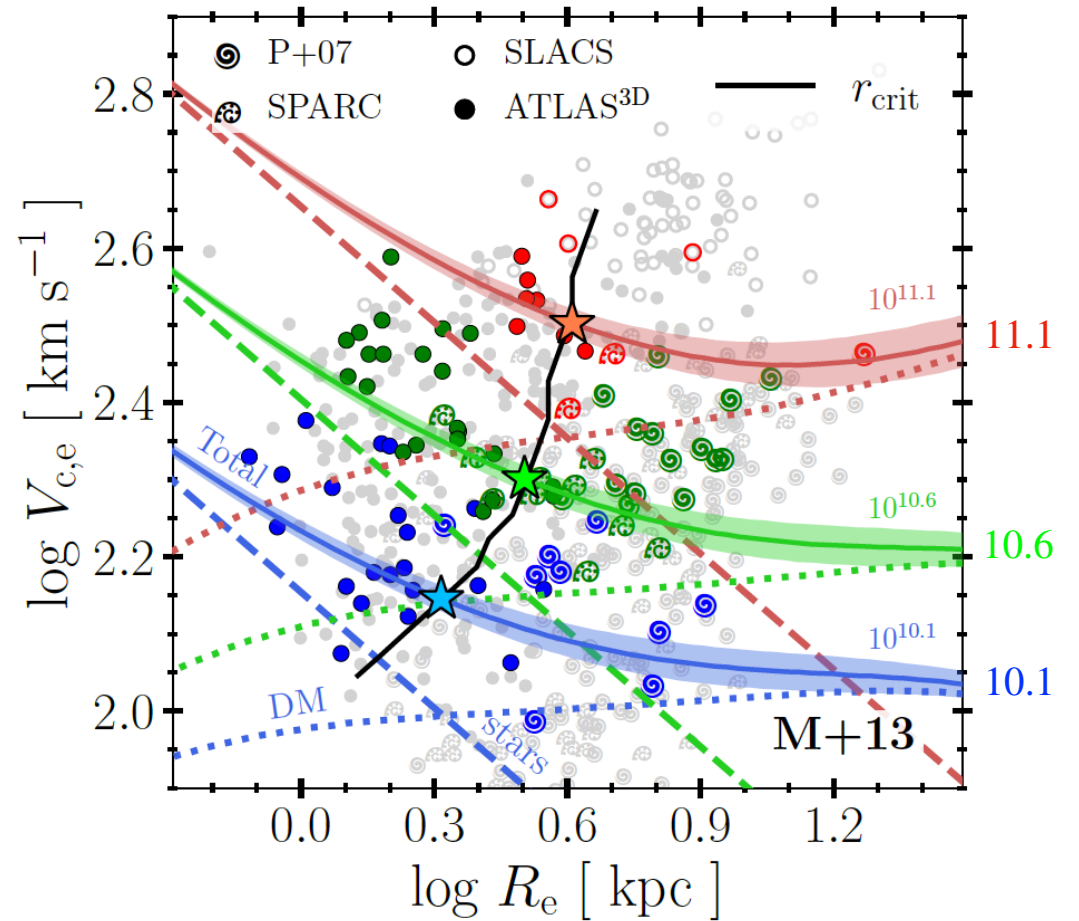


# Observed Scaling Relations





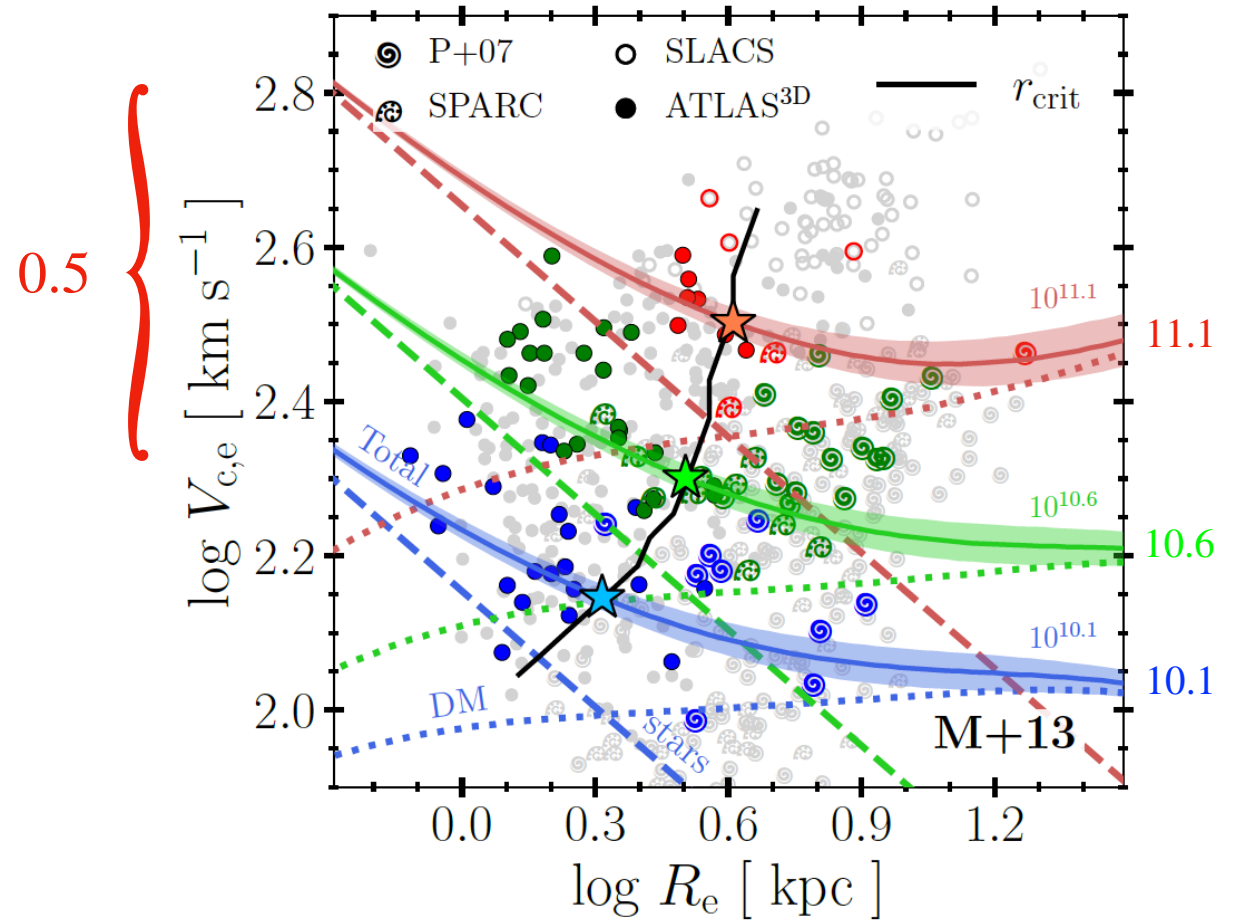
# Velocity-Radius Relation



# Velocity-Radius Relation

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.5}$$

$$M \sim V^2$$



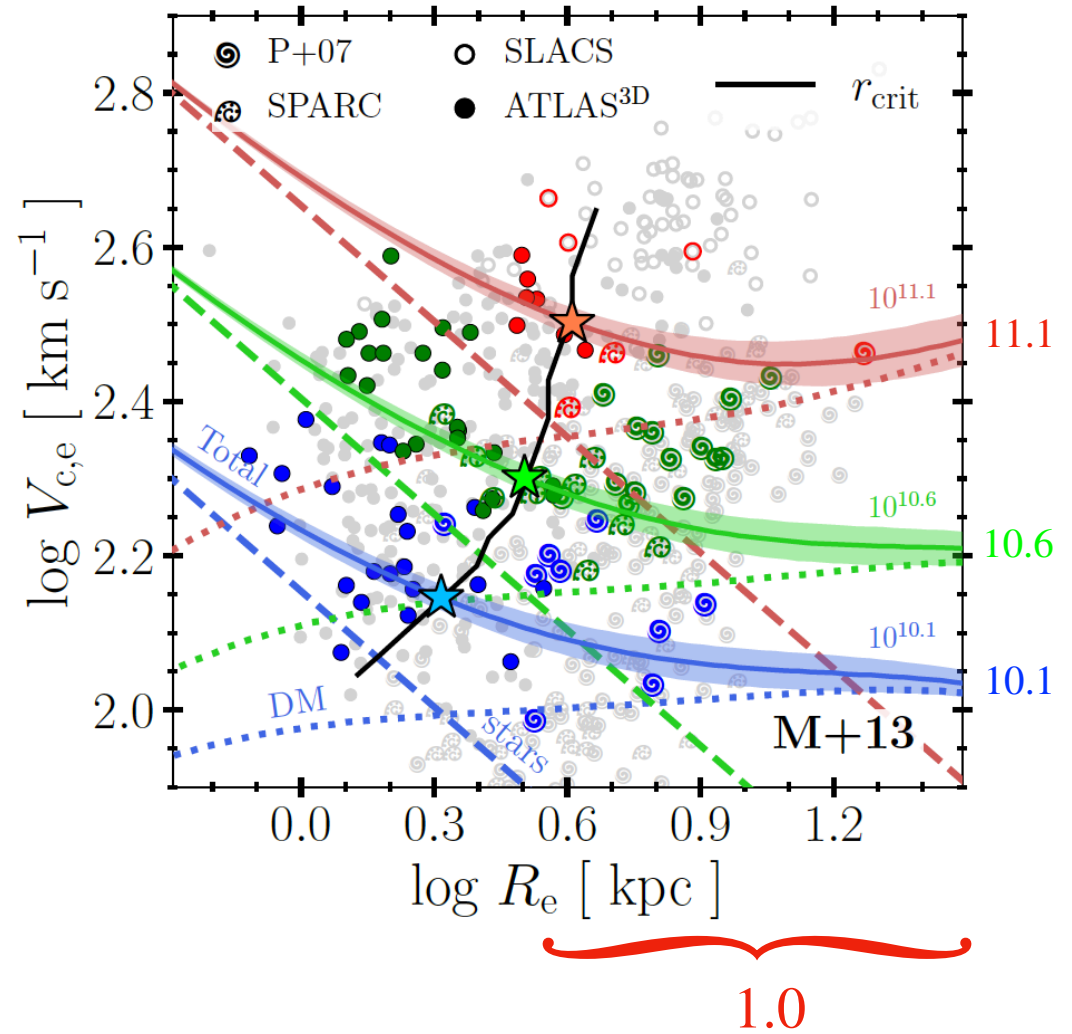
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$$M \sim V^2$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.0}{1.0}$$

$$M \sim R$$



# Velocity-Radius Relation

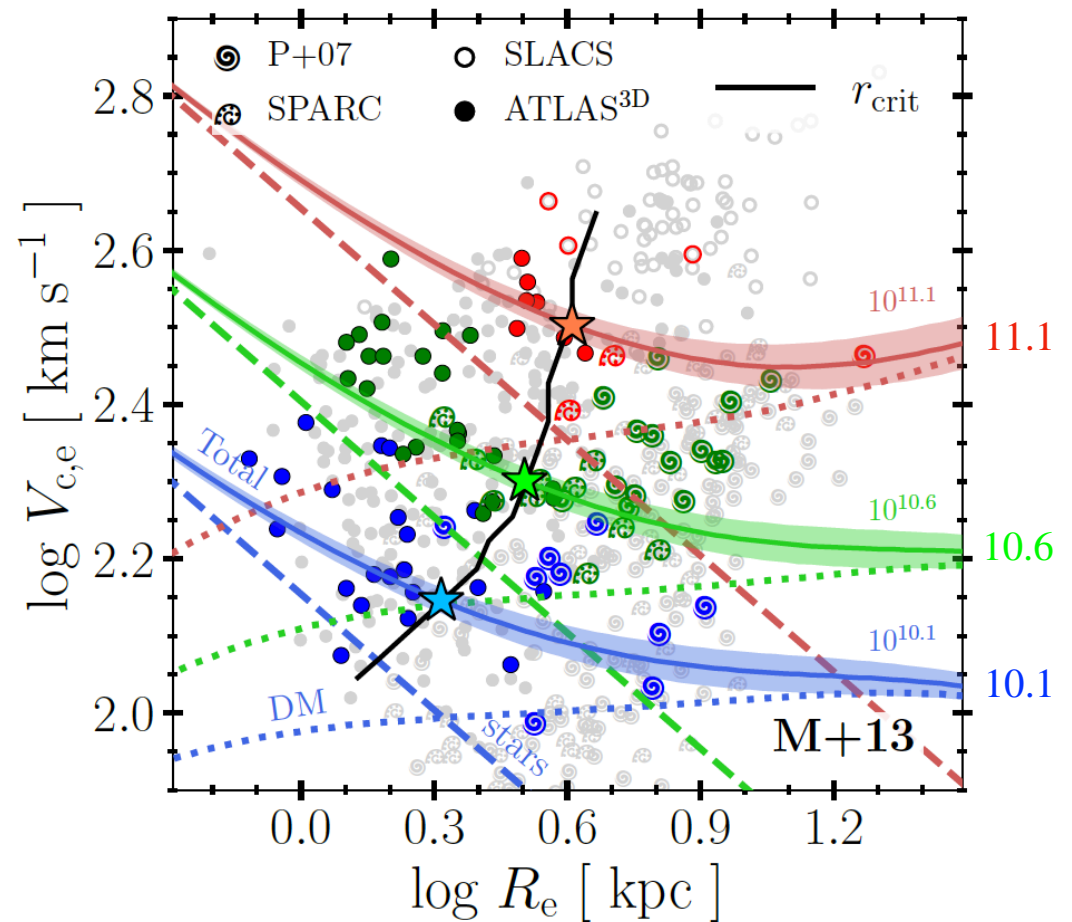
$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.5}$$

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$M \sim V^2 R$  Fundamental Plane



# Velocity-Radius Relation

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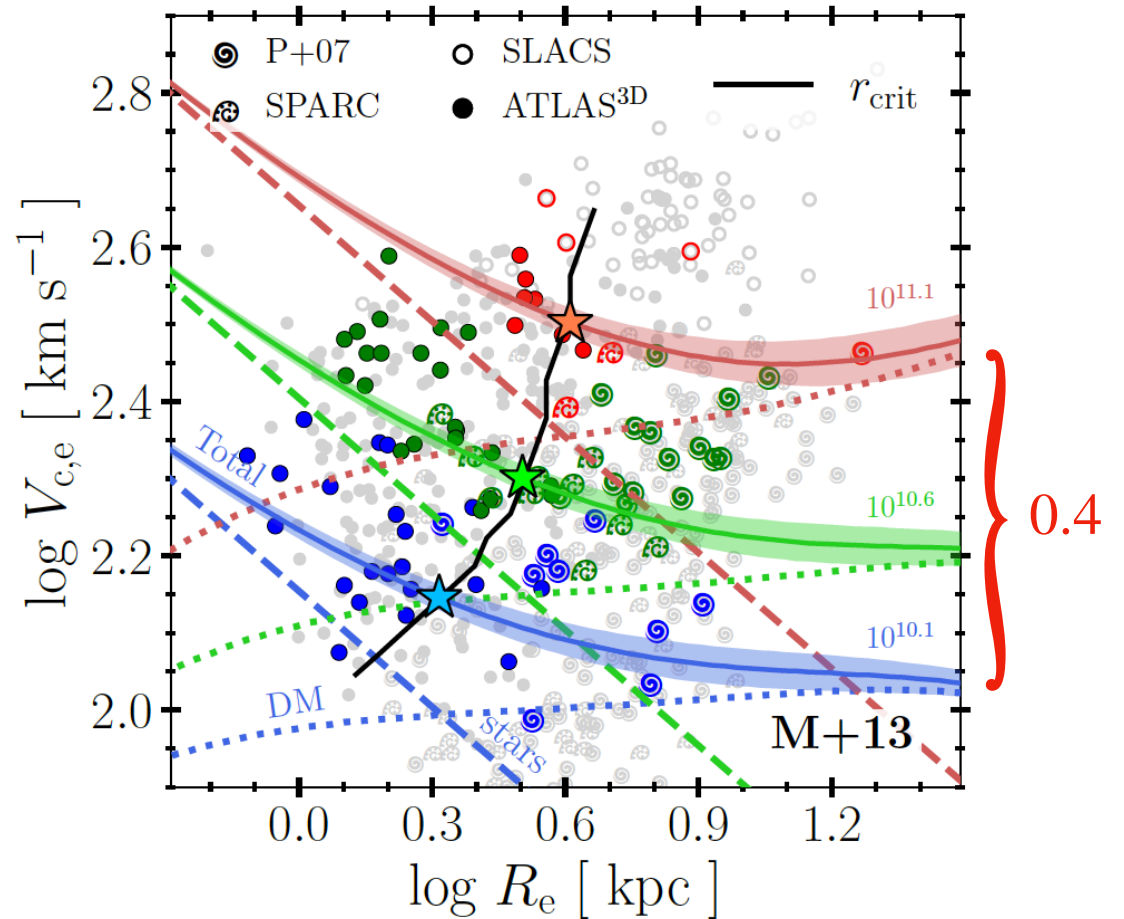
$$M \sim R$$

$M \sim V^2 R$  Fundamental Plane

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$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.4}$$

$$M \sim V^{2.5}$$



# Velocity-Radius Relation

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.5}$$

$$M \sim V^2$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.0}{1.0}$$

$$M \sim R$$

$M \sim V^2 R$  Fundamental Plane

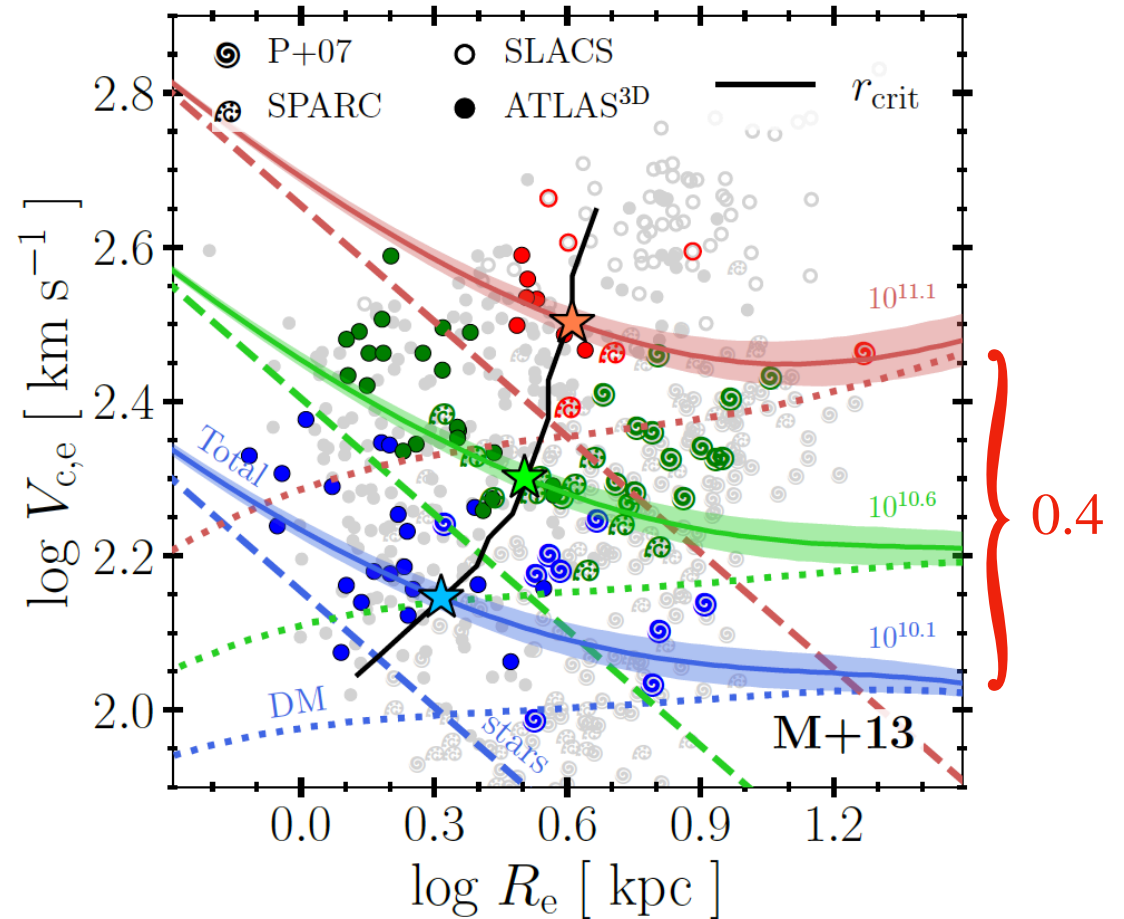
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$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.4}$$

$$M \sim V^{2.5}$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.0}{\infty}$$

$$M \sim R^0$$





# Velocity-Radius Relation

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.5}$$

$$M \sim V^2$$

$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.0}{1.0}$$

$$M \sim R$$

$$M \sim V^2 R \quad \text{Fundamental Plane}$$

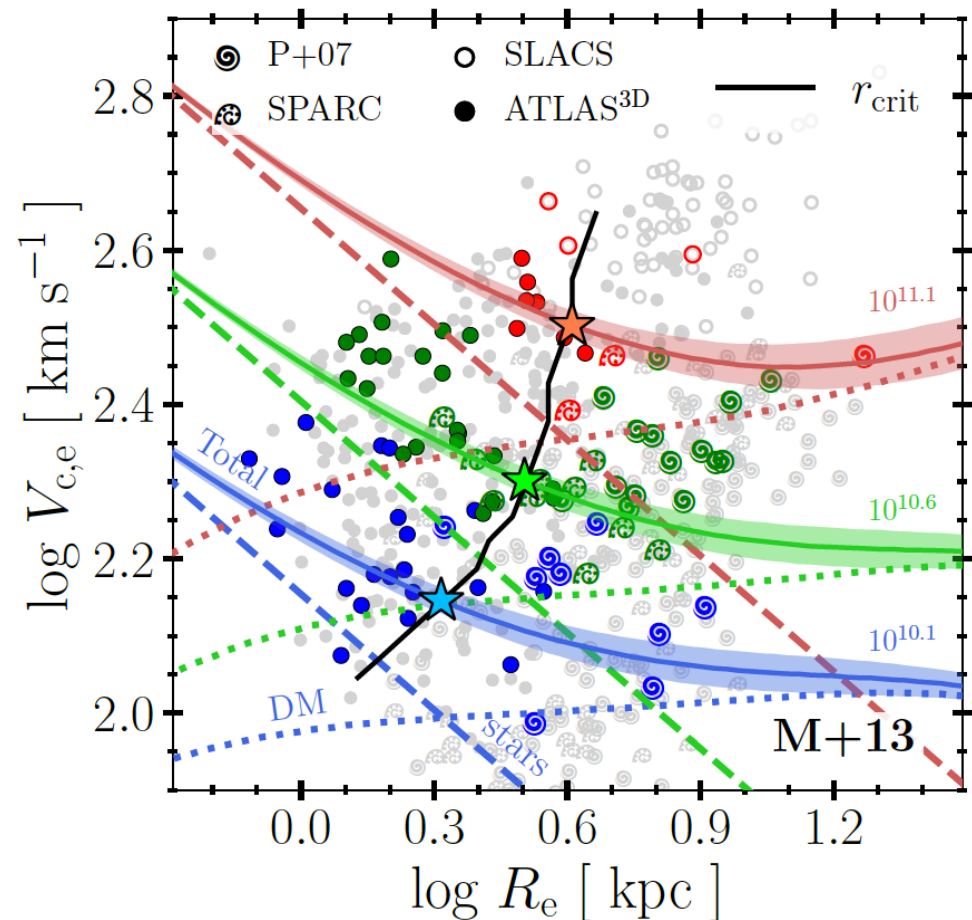
$$\frac{\Delta \text{Log} M}{\Delta \text{Log} V} \sim \frac{1.0}{0.4}$$

$$M \sim V^{2.5}$$

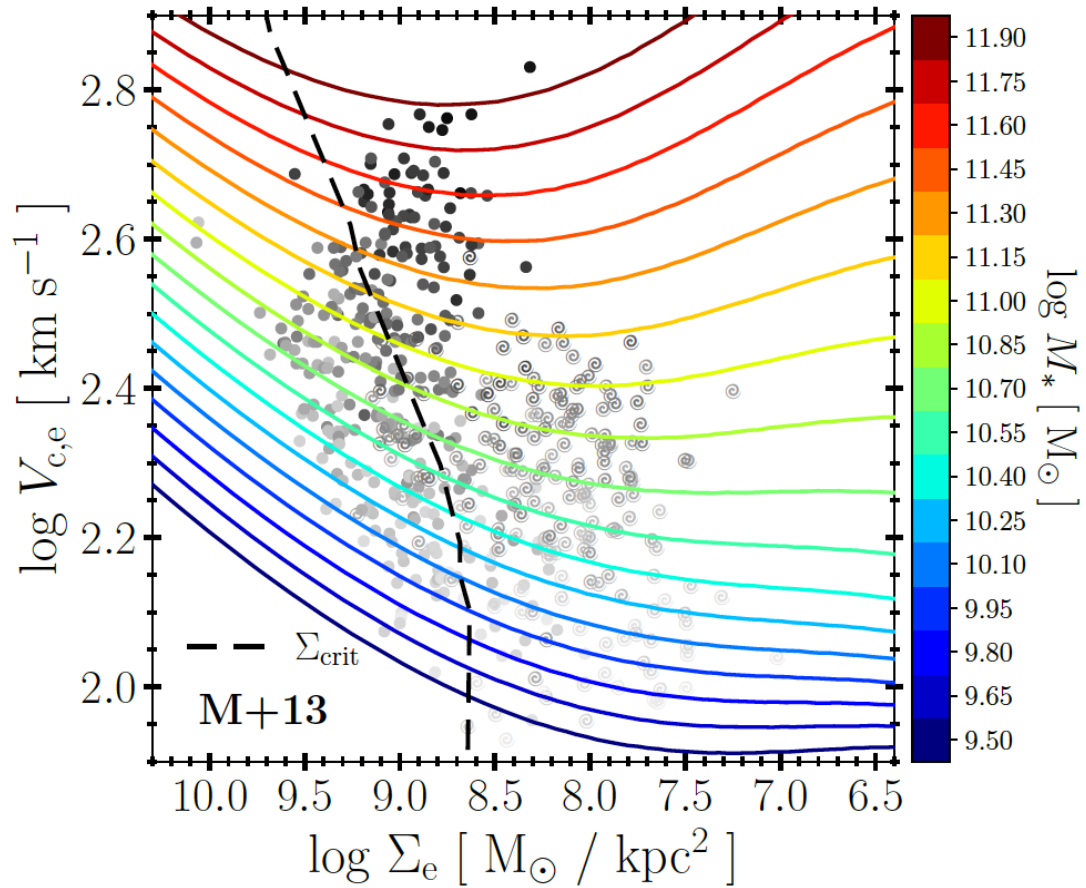
$$\frac{\Delta \text{Log} M}{\Delta \text{Log} R} \sim \frac{1.0}{\infty}$$

$$M \sim R^0$$

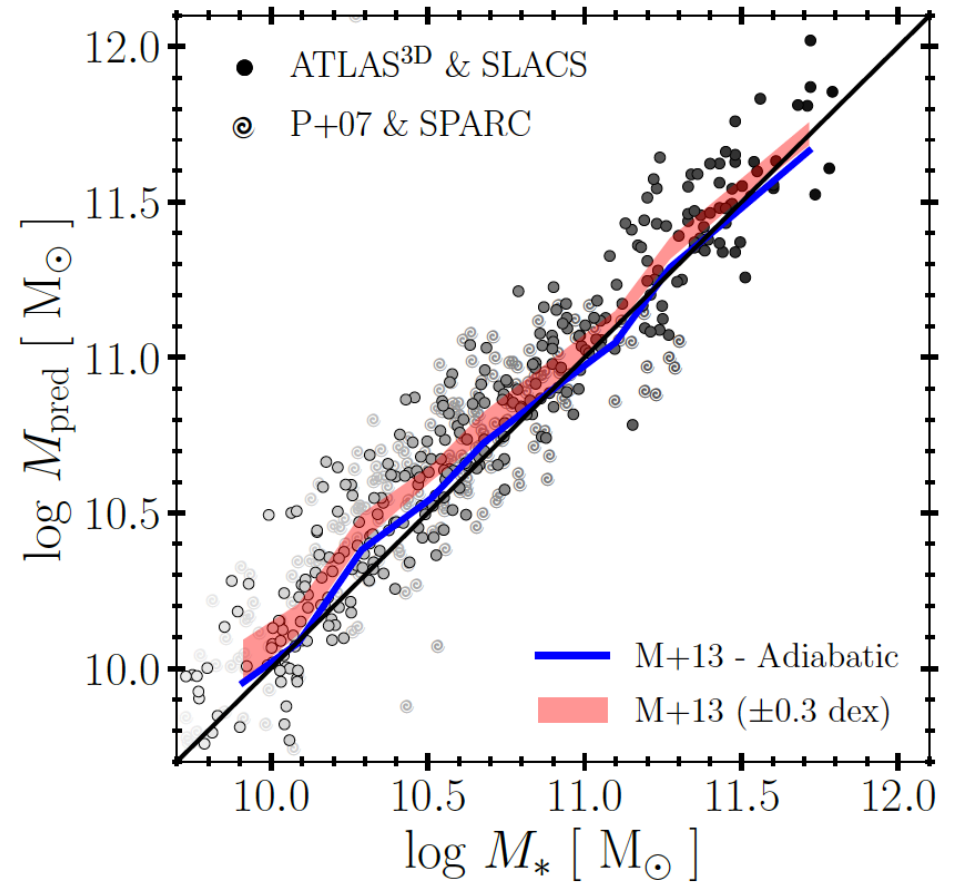
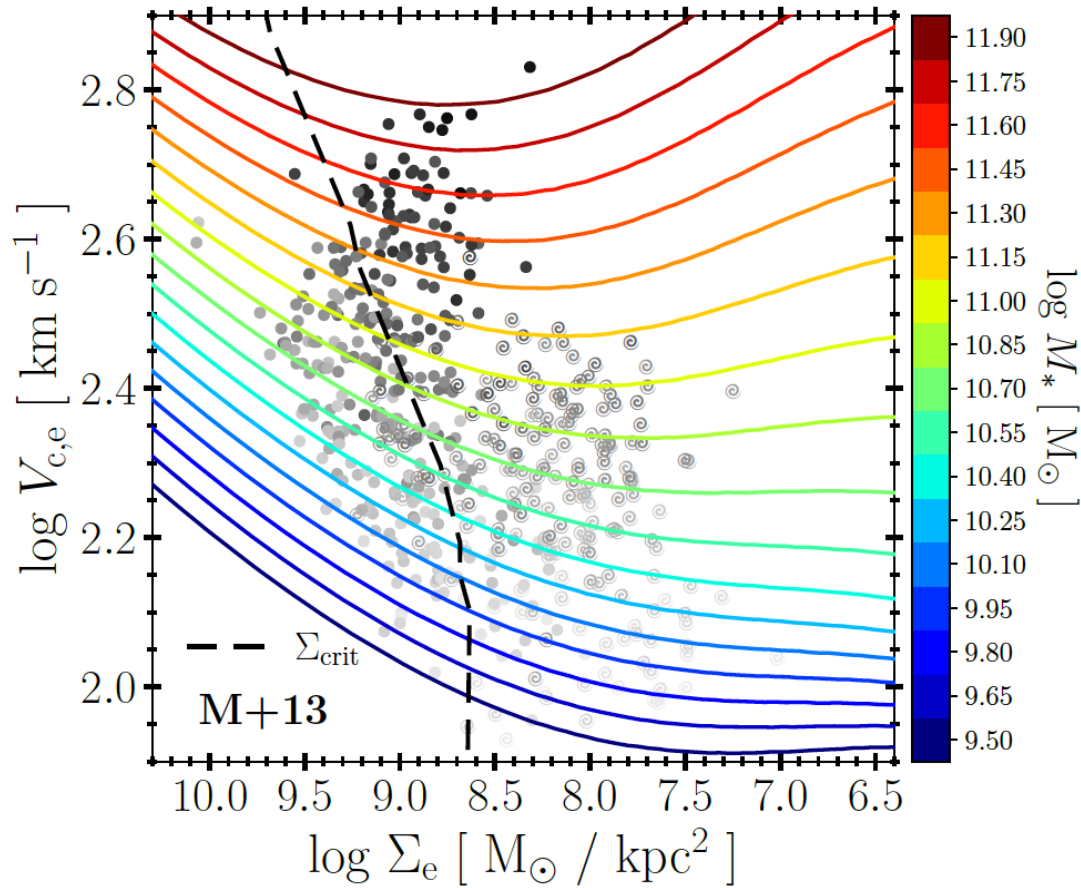
$$M \sim V^{2.5} \quad \text{Tully-Fisher}$$



# Secondary Distance Indicator



# Secondary Distance Indicator



# Take-home Message

- 1) Tully-Fisher and Fundamental Plane scaling relations emerge as a consequence of the different dark matter content inside the effective radius
- 2) A unified distance indicator gives competitive results for all galaxies independent of the morphological type.