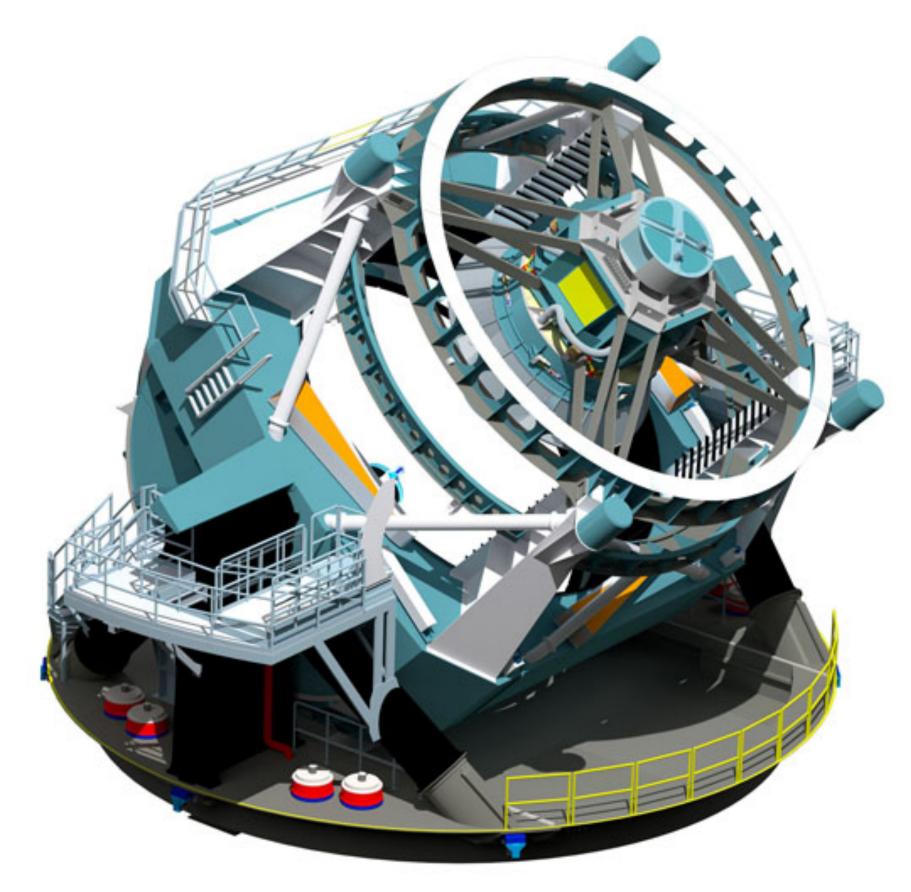
LSST and DESC Data Access and Analysis

Michael Wood-Vasey University of Pittsburgh



LSST in a Nutshell

The LSST is an integrated survey system designed to conduct a decade-long, 'wide-fast-deep' time-domain survey of the optical sky.



Key Characteristics

- 8.4 m primary mirror (6.5 m effective)
- Wide 9.6 deg² field-of-view
- High étendue (A Ω) of 319 m² deg²
- 3.2 Gpix camera, 6 filters: ugrizy
- Fully automated data processing system
- 20 TB data and up to 10 million alerts every night
- Interactive data access and analysis environment

2022 - 2032 LSST will enable a wide variety of complementary scientific investigations





High Level Requirements

Requirements determined by constraints from the four science drivers

Survey Property

Survey Area & Median Number of Visits/Point

Filter set

Single (Alternate) visit

Single Visit Depth (SNR > 5)

Astrometry (single visit)

Photometry

Proper Motion

Median delivered image quality (seeing)

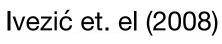




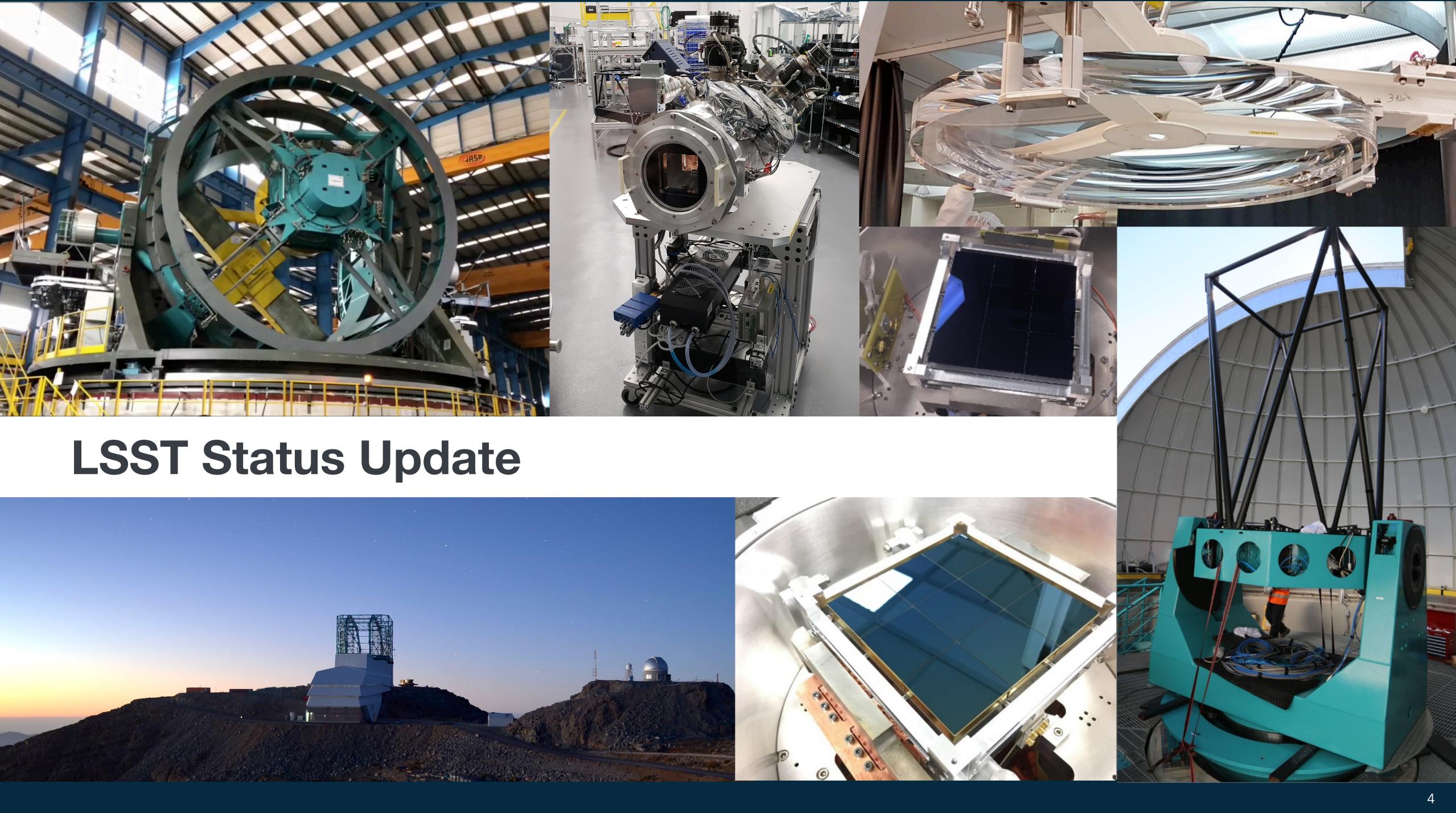


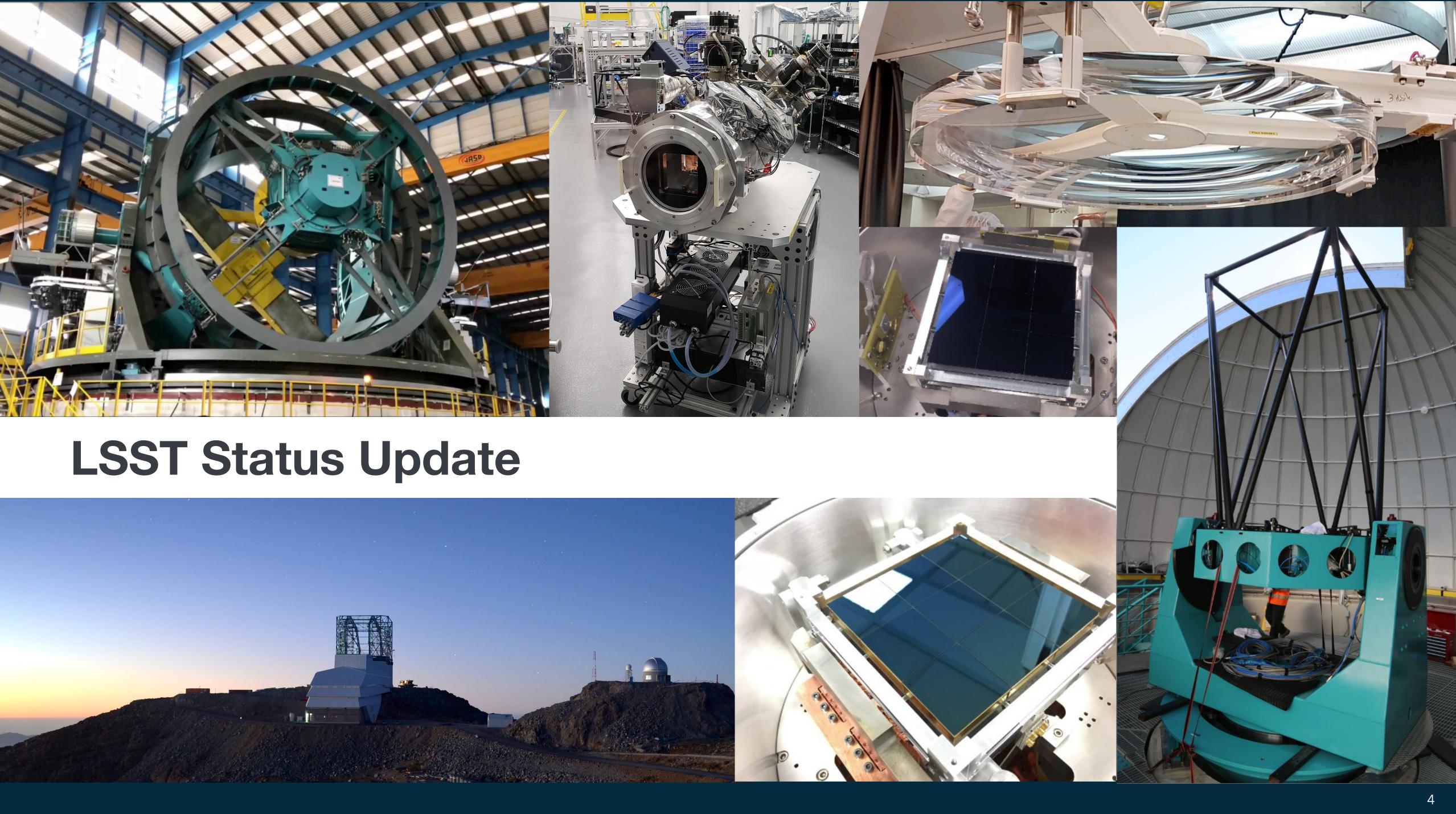
	Specification (Design Value)
ting	18,000 deg ² with 825 visits
	6 filters (ugrizy) from 320 - 1050 nm
	2x15s (1x30s) exposures
	u=23.5, g=24.8, r=24.4, i=23.3, y=22.1
	10 mas relative, 50 mas absolute
	0.5% relative, 2% absolute
	0.2 mas/yr @ r = 20.5, 1.0 mas/yr @ r = 24.0
	~0.7 arcsec FWHM

.... also drive the design of the data processing system









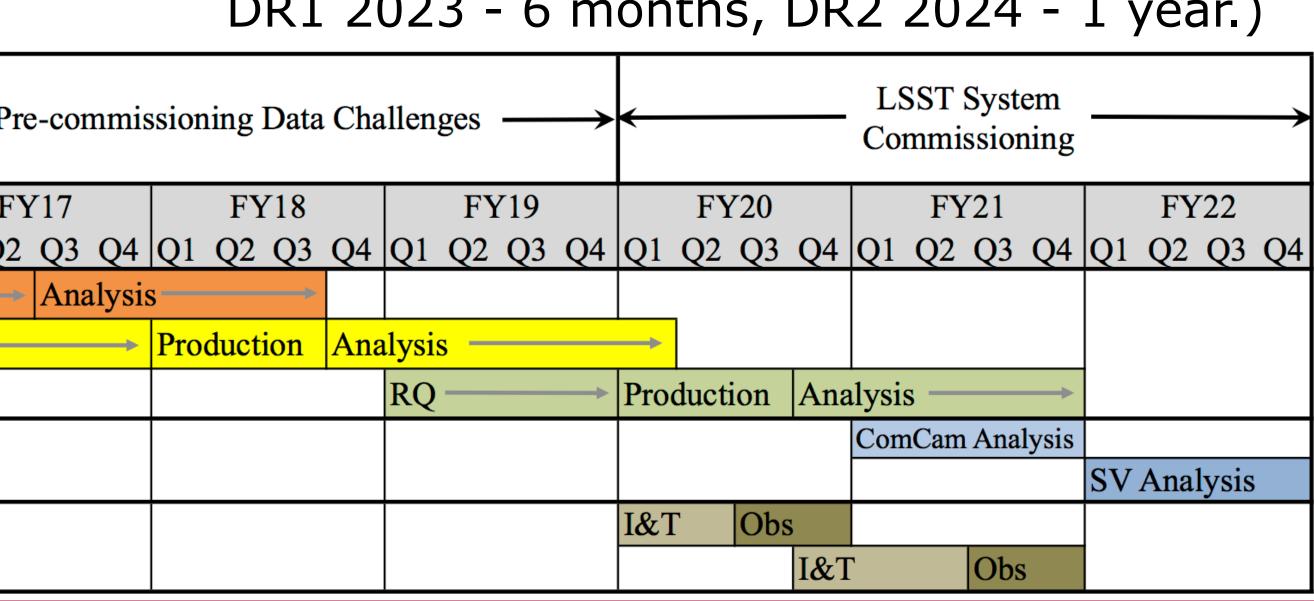
Dark Energy Science Collaboration

- Data Challenges to prepare for LSST 200 Full Members+ 600 interested Members
- https://lsstdesc.org
- Focused on Dark Energy Science with LSST DC2 (2017-2020)
- **DESC Science Requirements Document** https://arxiv.org/pdf/1809.01669.pdf
- DESC Science Roadmap https://lsstdesc.org/sites/default/files/ DESC_SRM_V1_4.pdf

		←	— LSST I	DESC P
			FY16	F
		Q1	Q2 Q3 Q4	Q1 Q2
LSST DESC	Data Challenge 1 (DC1)	RQ	Production -	
Pre-commissioning	Data Challenge 2 (DC2)			<mark>RQ —</mark>
Data Challenges	Data Challenge 3 (DC3)			
LSST DESC	ComCam Data Challenge			
Commissioning	SV Data Challenge			
LSST Facility	Early Commissioning, ComCam			
Commissioning	LSSTCam Commissioning			

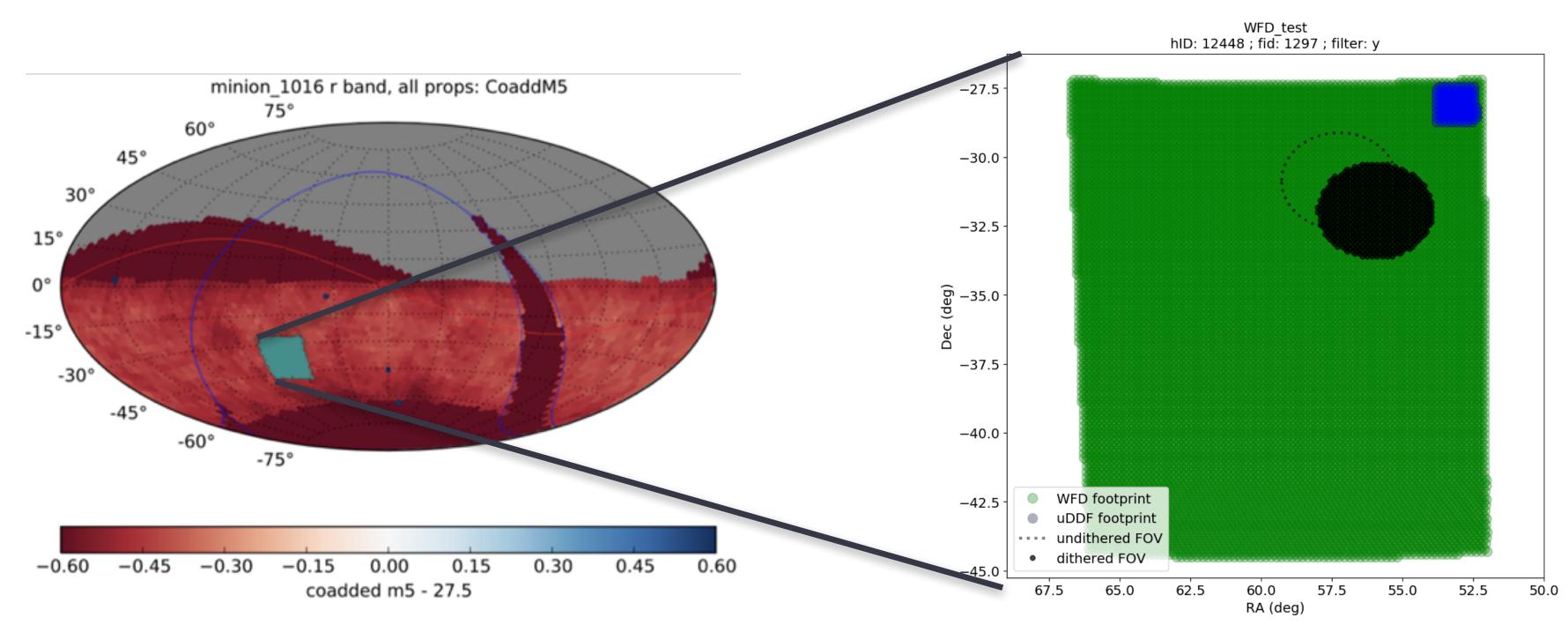


- DC1 (2016-2018)
- DC3 (2019-2021)
- LSST ComCam (2021)
- LSST Science Verification (2022)
- LSST Full Operational Data 2023.
 - DR1 2023 6 months, DR2 2024 1 year.)



DESC Data Challenge 2

- Overall: 75,000 visits in DC2 = \sim 7.5M sensor visits
 - Similar in scale to planned LSSTCam Science Validation survey
 - Long-lasting legacy value, e.g., possible DC3 reprocessing activities
- Major supercomputing and human resources required!



DC2 patch (from DESCQA, Mao et al. 2018)



M sensor visits m Science Validation survey ssible DC3 reprocessing activities esources required!

WFD and DDF dithers, Run 2, Image Credit: H. Awan

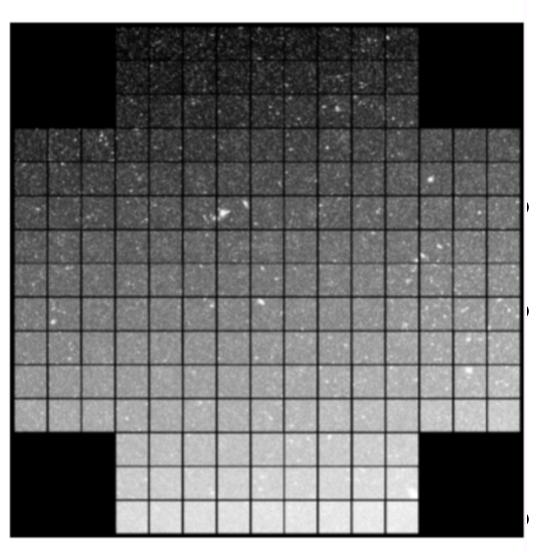
DESC Data Challenge 2

- Full end-to-end simulation and processing to enable testing and development of
 - Analysis and processing pipelines
 - Mitigation strategies for systematic effects
 - E.g., sensor defects, blending, atmospheric effects,
 - Data access strategies
- Design
 - Extragalactic sky: 5,000 sq deg input catalog
 - Static sky (WL, CL, LSS, PZ) with images:
 - 300 sq deg "main survey" area
 - 10 years ugrizy Wide-Fast-Deep (WFD) cadence
 - Time domain (SN, SL) and deep analyses:
 - 1 sq deg "Deep Drilling Field (DDF)" embedded in corner of main survey
 - 10 years ugrizy WFD + DDF visits
- Start with 3-year simulated survey, which deep enough for useful tests of many science cases!



cts pheric effects

Image Simulation Properties



Simulated LSST focal plane

- Objects
 - DC2 galaxies (clustering, WL) Ο
 - AGNs Ο
 - Stars from CatSim Ο
 - Sprinkled time-varying objects Ο (lensed AGNs, SNe)

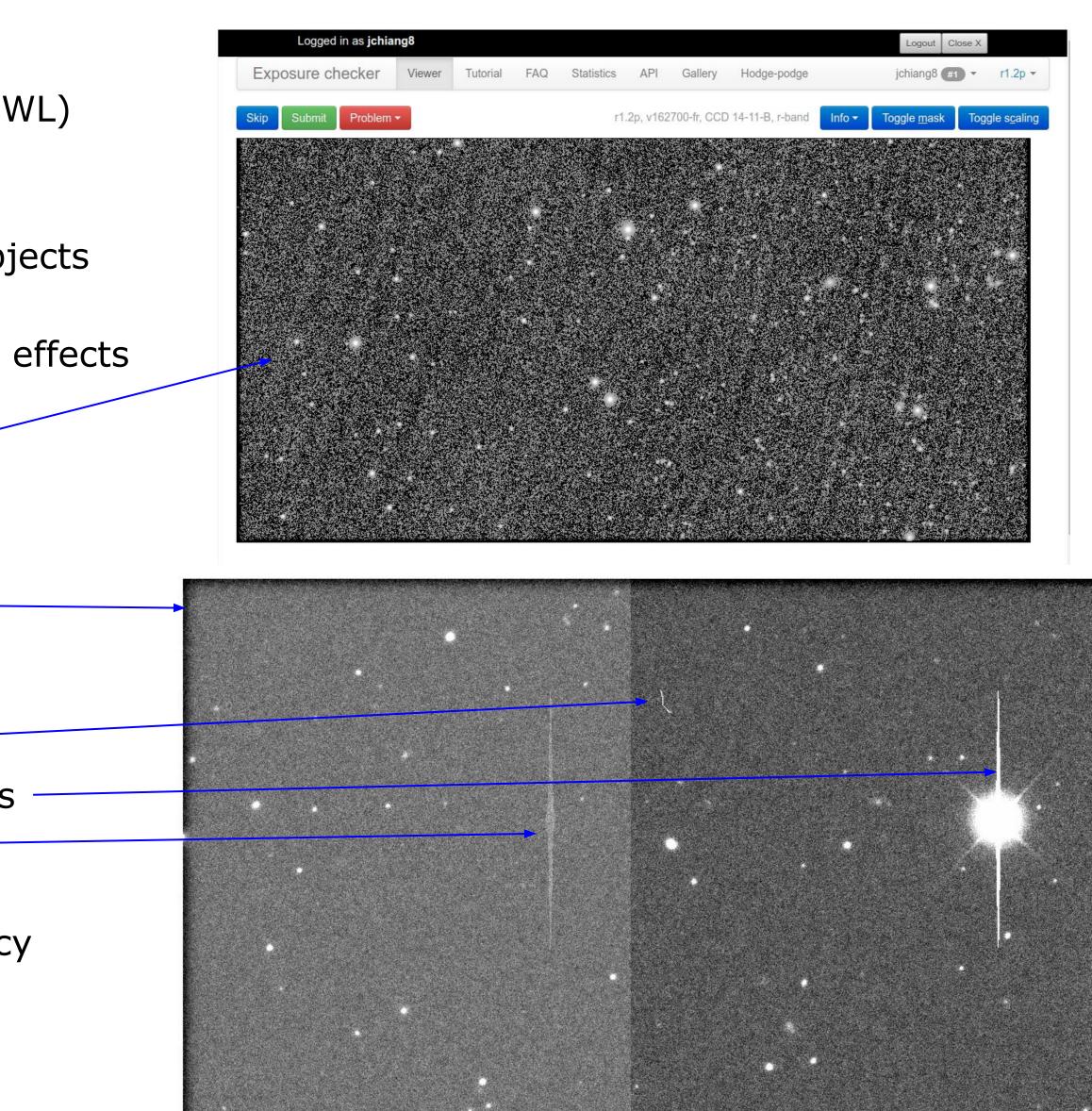
PSF modeled with atmospheric effects and optical distortions Sensor effects:

- Treerings Ο
- Brighter-fatter Ο
- Edge rolloff Ο

Instrument signature features

- Vignetting Ο
- Cosmic Rays Ο
- Saturation and bleed trails Ο
- Crosstalk Ο
- Bright/dark defects Ο
- Charge transfer inefficiency Ο





Adapted DES Exposure Checker for DC2

- Most common issues with background estimation in crowded regions and spurious cosmic ray masking (not major issues for DC2 science)
- Useful feedback on the tools and the data
- Excellent for quick diagnosis in early image simulation productions.

(Slide courtesy of Alex Drlica-Wagner)



Exposure checker

Look at simulated LSST images. Discover flaws we would otherwise have missed.

Make our simulations and data processing better!





9339

What's this good for?

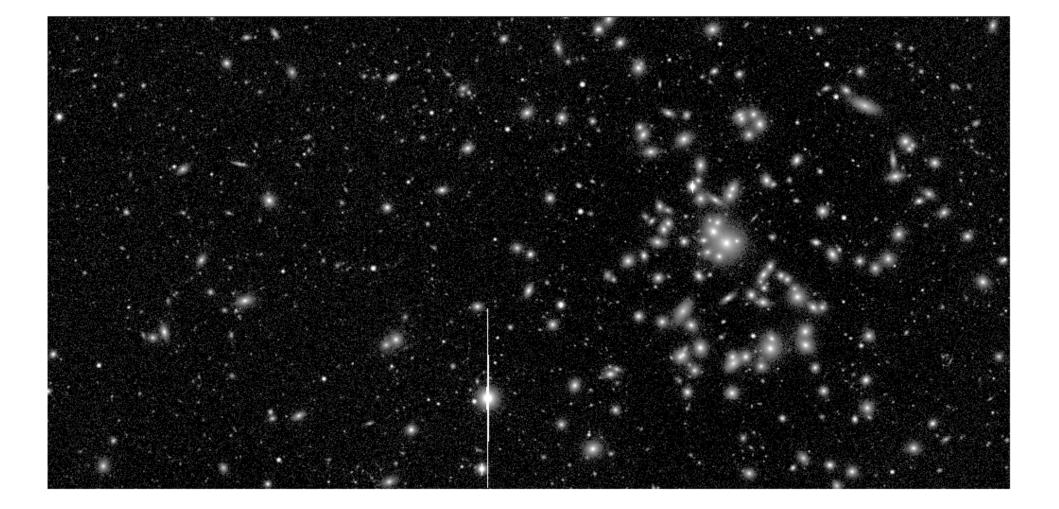
We all want to do science with LSST images. However, no data set is perfect, and we need a way to identify artifacts in our data. The Exposure Checker helps in two ways:

- It gives you convenient access to the images and demonstrates the performance of the algorithms that are already in place to identify and correct known problems.
- It enables everyone to search for undetected artifacts. We gather your submissions, analyze them, and feed them back to the simulation and data processing teams.

How can I start?



# 1	gplynch	1400
# 2	jneveu	1344
# 3	heitmann	1024
# 4	rcecile	1000
# 5	ccombet	700
#6	bleeml	600

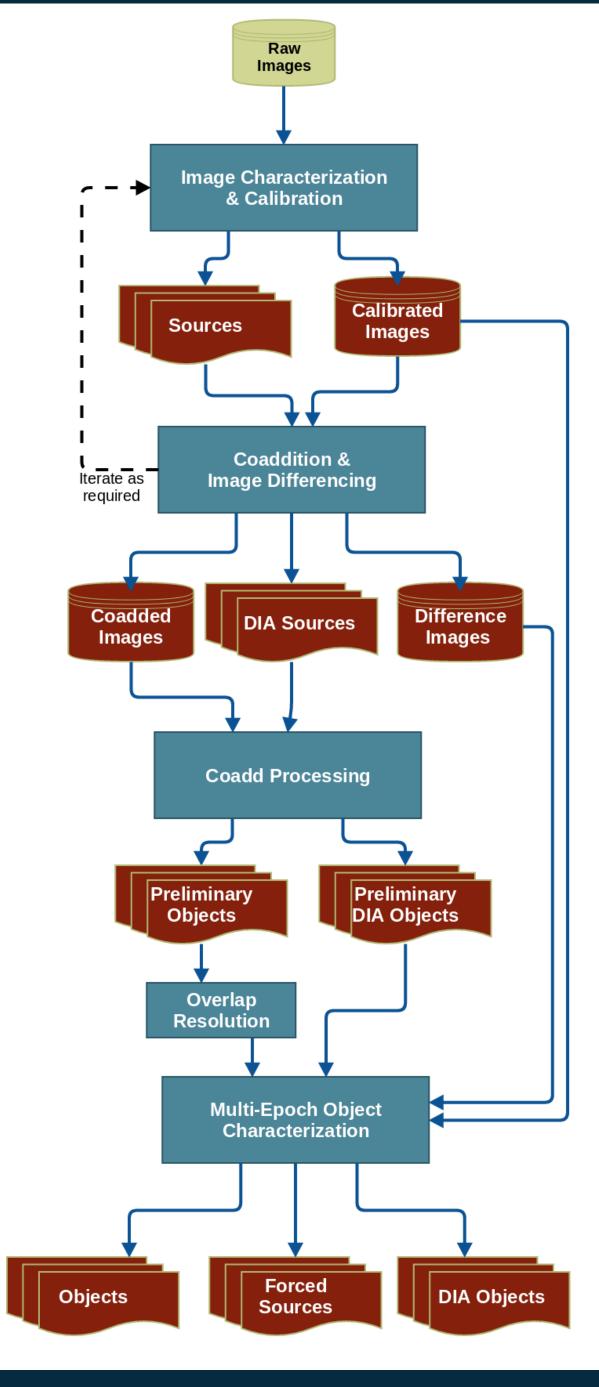


Data Release Data Products

LSST will produce a 'science-ready' database of measurements can be analysed without the need for complex image processing

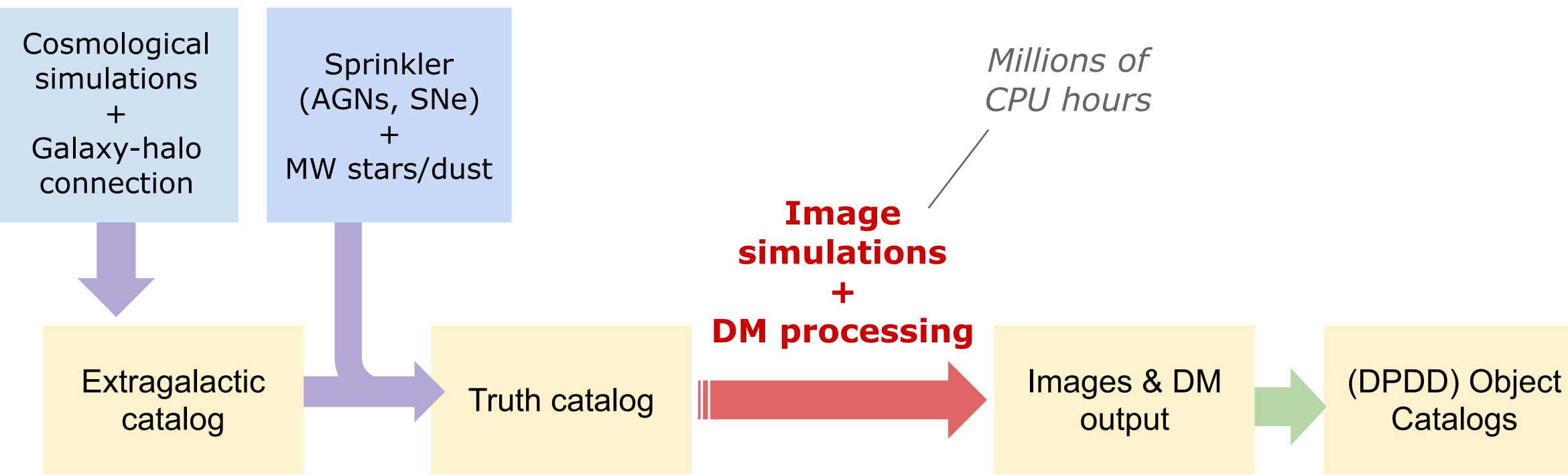
	Single-Visit	fully process processed
Images	Co-add	short-period best seeing & d one per filter (u transient-fi
	Source (single-epoch)	Detections i positions, aperture de-bl
Catalogs	Forced Source (single-epoch)	<i>In all single-vis</i> point-source flu
	Object (associations of sources representing astrophysical phenomena)	All sources as well as positions, aperte de-blendin seeing-indep

sed 2 x 15s exposures d visit images (PVIs) d (yearly & full survey) deepest (unless the same) (ugrizy), and 'multi-color' Free template images in single-visit images: e fluxes, point source fluxes, lending results sit images for all sources: uxes, forced photometry detections in co-added images: ure & point-source fluxes, ng results, model fits pendent galaxy colours





DC2 Key Data Products



Value-added data products





Overview of DC2 Data Products

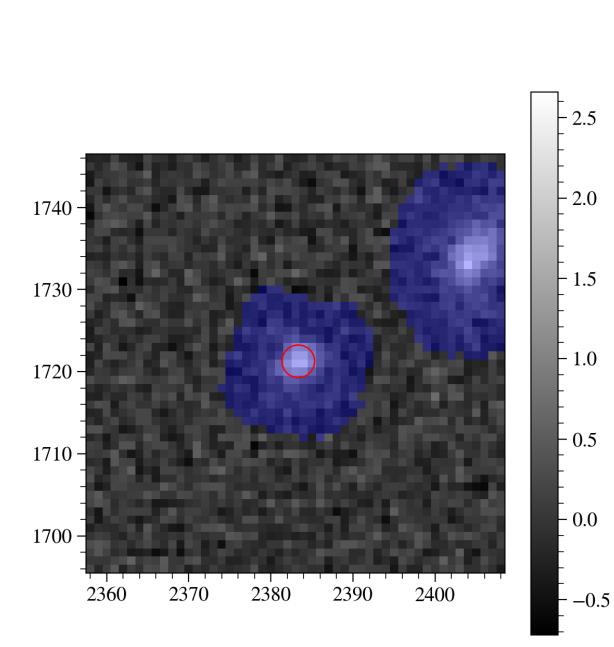
- CosmoSim team prepares extragalactic catalog
- ImSim team simulations images at NERSC and Argonne
- DM DC2 team processes images at IN2P3 and NERSC
- Data Access Task Force prepares data products for use by DESC members

Name	Description	Access
Extragalactic catalogs	Mock catalogs from cosmological simulations, including galaxies, shears,	GCR
Truth catalogs	Truth information about all objects in the simulated images, including stars, AGNs, SNe,	GCR
DM science pipeline outputs	Range of data products from the DM pipeline, including raw images, calibrated exposures,	Butler
DPDD(-like) object catalogs	Merge static-sky coadd catalogs with forced-position photometry at fixed RA, Dec across bands	GCR, pandas, spark



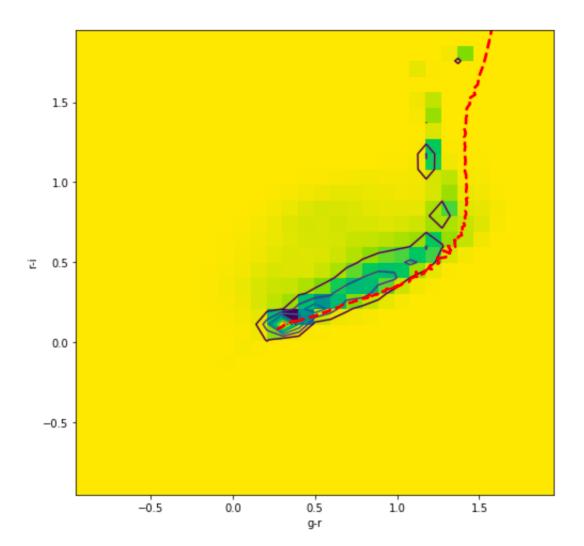
DESC Data Access

- MWV and Dominique Boutigny **Co-Lead Data Access Task Force**
- Provide data access for DESC
 - Different data access methods are Ο being explored:
 - GCR: Generic catalog reader
 - Qserv (project): Inject data at IN2P3 (SLAC instance)
 - PostgreSQL mirroring HSC set-up, spark
 - hdf5/parquet files





Postage stamp for set RA, Dec, Image Credit: Michael Wood-Vasey



Color-color plot, Image Credit: Joanne Bogart





Example Jupyter Notebooks on DC2

- <u>https://github.com/LSSTDESC/DC2-analysis/blob/rendered/tutorials/</u>
 <u>object_pandas_stellar_locus.nbconvert.ipynb</u>
- <u>https://github.com/LSSTDESC/DC2-analysis/blob/rendered/tutorials/</u>
 <u>object_spark_1_intro.nbconvert.ipynb</u>
- <u>https://github.com/LSSTDESC/DC2-production/blob/master/Notebooks/</u> <u>object_catalog_performance_dask_parquet.ipynb</u>



DESC DC2 Data Volumes

		Run 1.2 25 sq. deg -	2,000 visits	Run 2.1 300 sq. deg - 75,000 visits					
Туре	Product	Size	# Files	Size	# Files				
Images	Raw	5 TB	300,000	125 TB	7,500,000				
Images	Processed	30 TB	1,000,000	750 TB	21,500,000				
Tables	Object	4 GB	1	100 GB	1				
Tables	Source	30 GB	2,000	1,125 GB	75,000				
Tables	Forced Source	10 GB							

I quote images volume in TB and catalog in GB both because those are they natural scales But also because image volume determines disk storage, which is a TB-scale question While catalog volume determines memory size, which is a GB-scale question.





DESC DC2 Data Access

- Centered at NERSC
- Plan to extend to IN2P3.
 - Once that's done, likely easy extend to additional DESC partners.
- Jupyter Notebooks main user and exploration interaction
- Intensive pipelines, such as galaxy-galaxy, galaxy-shear, shear-shear correlations base on more programatic access.



end to additional DESC partners. exploration interaction -galaxy, galaxy-shear, shear-shear atic access.

DESC DC2 Data Volumes

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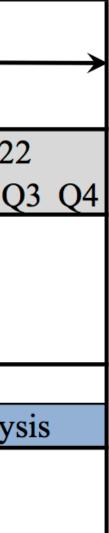




DESC Data Challenge Timeline Toward LSST Start

		\leftarrow LSST DESC Pre-commissioning Data Challenges \longrightarrow									\rightarrow					ST Syst missio				
			FY16		FY1	7		FY18]]	FY19			FY20			FY21			FY22
		Q1	Q2 Q3 Q4	Q1	Q2 (23 Q4	Q	Q1 Q2 Q3	Q4	Q1 Q	2 Q3	Q4	Q1 (Q2 Q3	Q4	Q1 (Q2 Q3	Q4	Q1	Q2 Q
LSST DESC	Data Challenge 1 (DC1)	RQ	Production		> A	Analysi	S	•												
Pre-commissioning	Data Challenge 2 (DC2)			RQ			P	roduction	Ana	alysis			•							
Data Challenges	Data Challenge 3 (DC3)									RQ —		→]	Prod	uction	Ana	alysis				
LSST DESC	ComCam Data Challenge															Com	Cam Ana	lysis		
Commissioning	SV Data Challenge																		SV A	Analy
LSST Facility	Early Commissioning, ComCam]	I&T	Obs	5					
Commissioning	LSSTCam Commissioning														I&7	Γ	Ob	S		





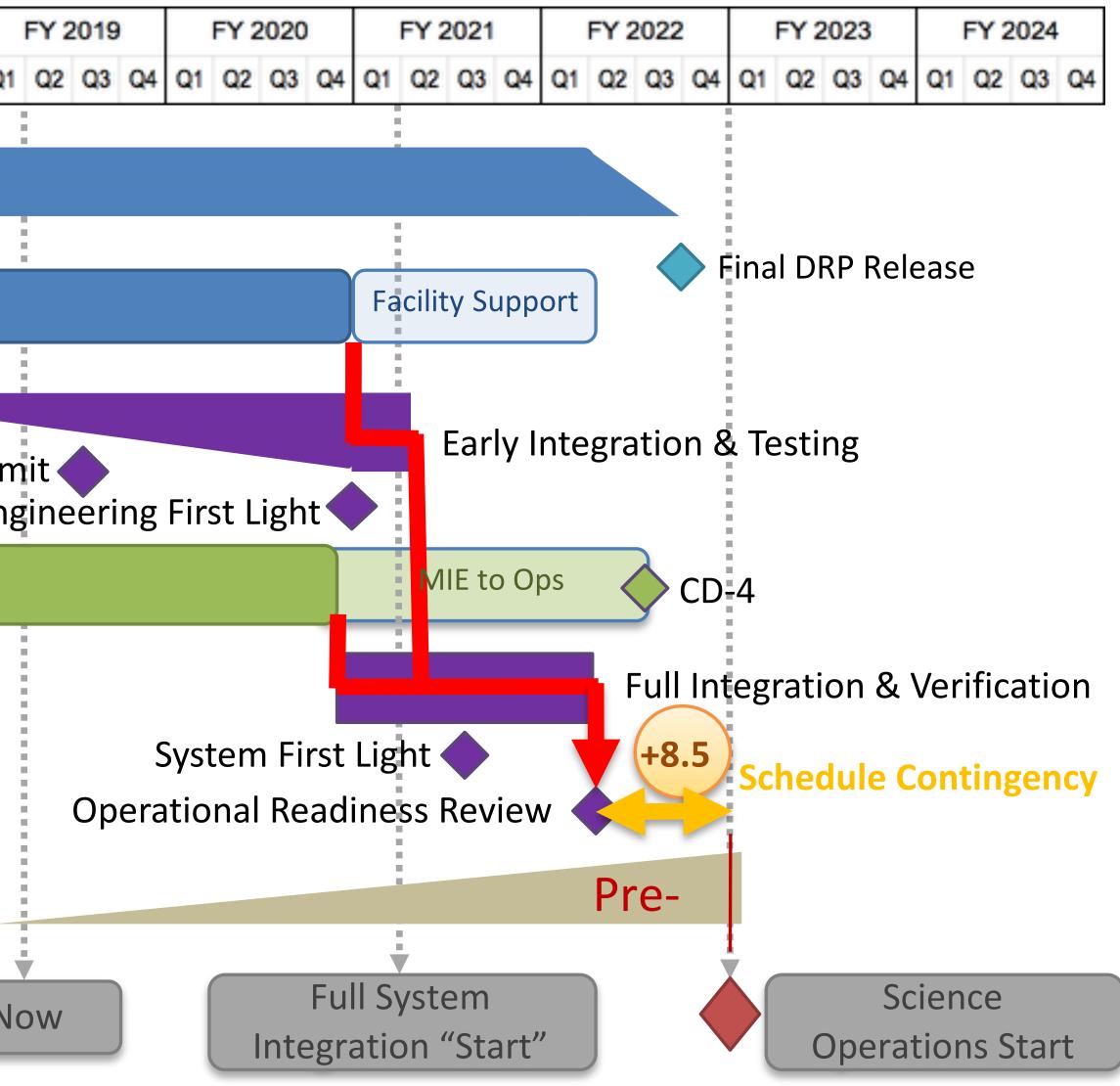
Extra Slides





FY 2	2014		FY 2015			Y 2015 FY 2016 FY								Y 2017 FY 201					
Q1 Q2	Q3	Q4	Q1	Q2	Q3	Q4	Q1	Q2	Q3 (Q 4	Q1	Q2	Q3	Q4	Q1	Q2	Q3	Q4	Q1
					Da	ta	N	lar	lag	e	me	en	t						
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LSST Key Dates

Milestone

Start of On-Sky Data from Auxiliary Telesco

Start of On-Sky & Calibration Data with Con

Sustained Observing with ComCam

Start of On-Sky & Calibration Data with LSS

Sustained Observing with LSSTCam

Start of Science Verification Surveys

Operations Readiness Review



Date
Aug 2019
Oct. 2020
Feb. 2021
July 2021
Oct. 2021
Dec. 2021
Mar 2022





Data Product Categories

The LSST data products are organized into three main categories.



Prompt Data Products

Real Time Difference Image Analysis (DIA)

- A stream of ~10 million time-domain events per night (Alerts), transmitted to event distribution networks within 60s of camera readout.
- Images, Object and Source catalogs derived from DIA, and an orbit catalog for ~6 million Solar System bodies within 24h.
- Enables discovery and rapid follow-up of time domain events



User Generated Data Products

User-produced derived, added-value data products

- Deep KBO/NEO, variable star classifications, shear maps, etc ...





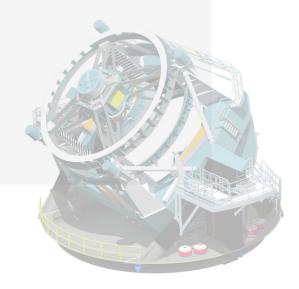
Data Release Data Products

Reduced single-epoch & deep co-added images, catalogs, reprocessed DIA products

- Catalogs of ~37 billion objects (20 billion galaxies, 17 billion stars), \sim 7 trillion sources and \sim 30 trillion forced source measurements.
- 11 Data Releases, produced ~annually over 10 years of operation
- Accessible via the LSST Science Platform & LSST Data Access Centers.

• Enabled by services & computing resources at the LSST DACs and via the LSST Science Platform (LSP).

10% of LSST computing resources will be allocated for User Generated data product storage & processing.





Nomenclature

Processed Visit Image: LSST image from a single visit with the background subtracted.

Source: single detection of an astrophysical object in an image, **Forced Source:** measurement a source, with one or more parameters held fixed **Object:** association of Sources by coordinate, an astronomical object (star, galaxy)

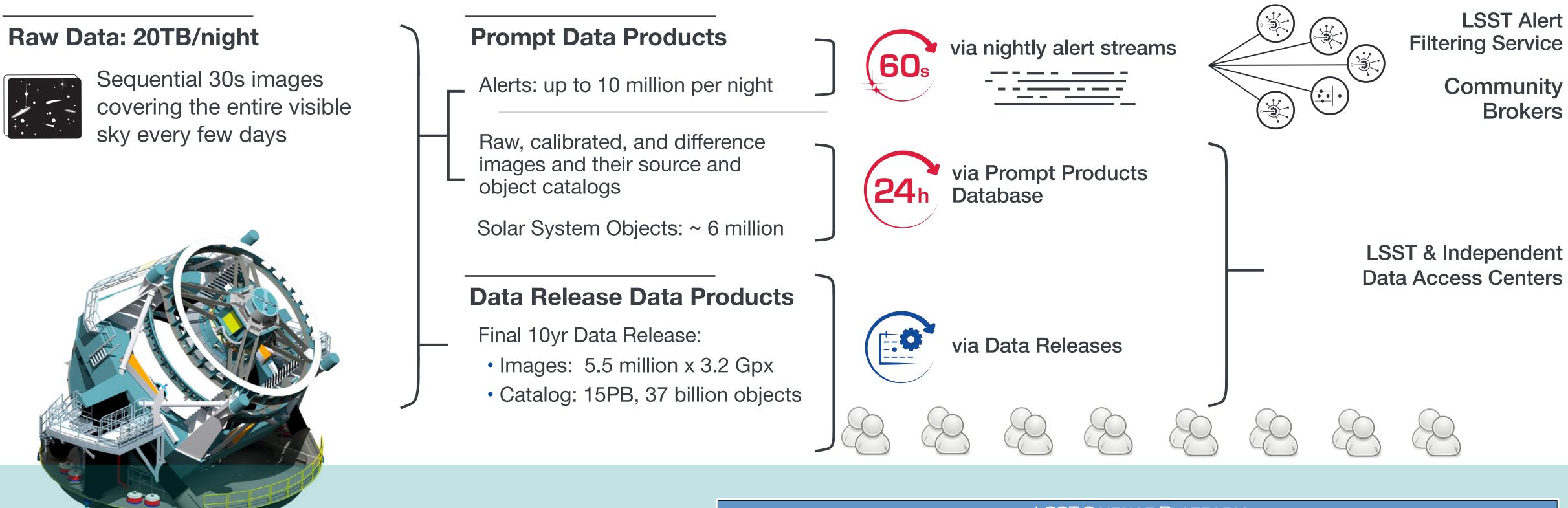
Analysis, Alert Production. This processing generates **Prompt Data Products**. identify DIASources and DIAObject. This processing generates *Alerts*. Data Release Production (DRP): A (re)processing all of the accumulated LSST images. This processing generated the **Data Release data products**.



- **Source Association:** association of sources at different epochs, or passbands, with an Object
- **Prompt Processing:** processing of the nightly stream of raw images, including Difference Imaging
- Alert Production (AP): component of Prompt Processing that processes and calibrates images to

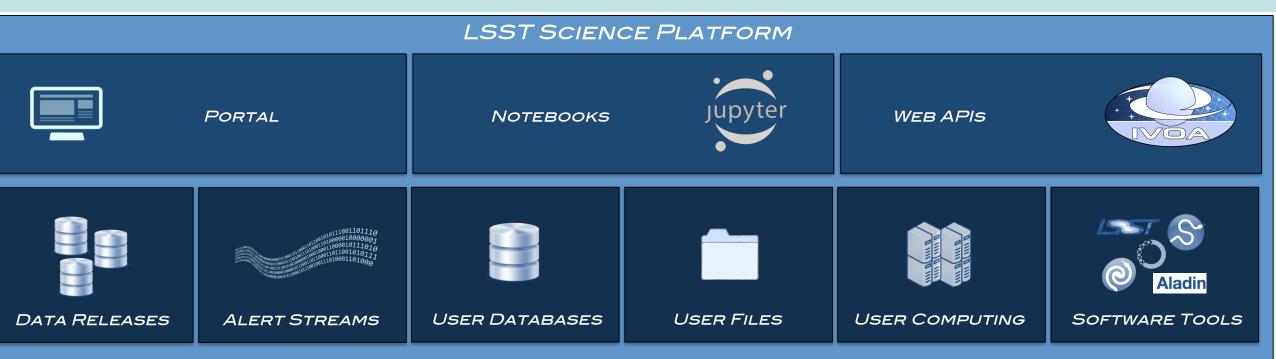
23

LSST Data Management System



LSST Science Platform

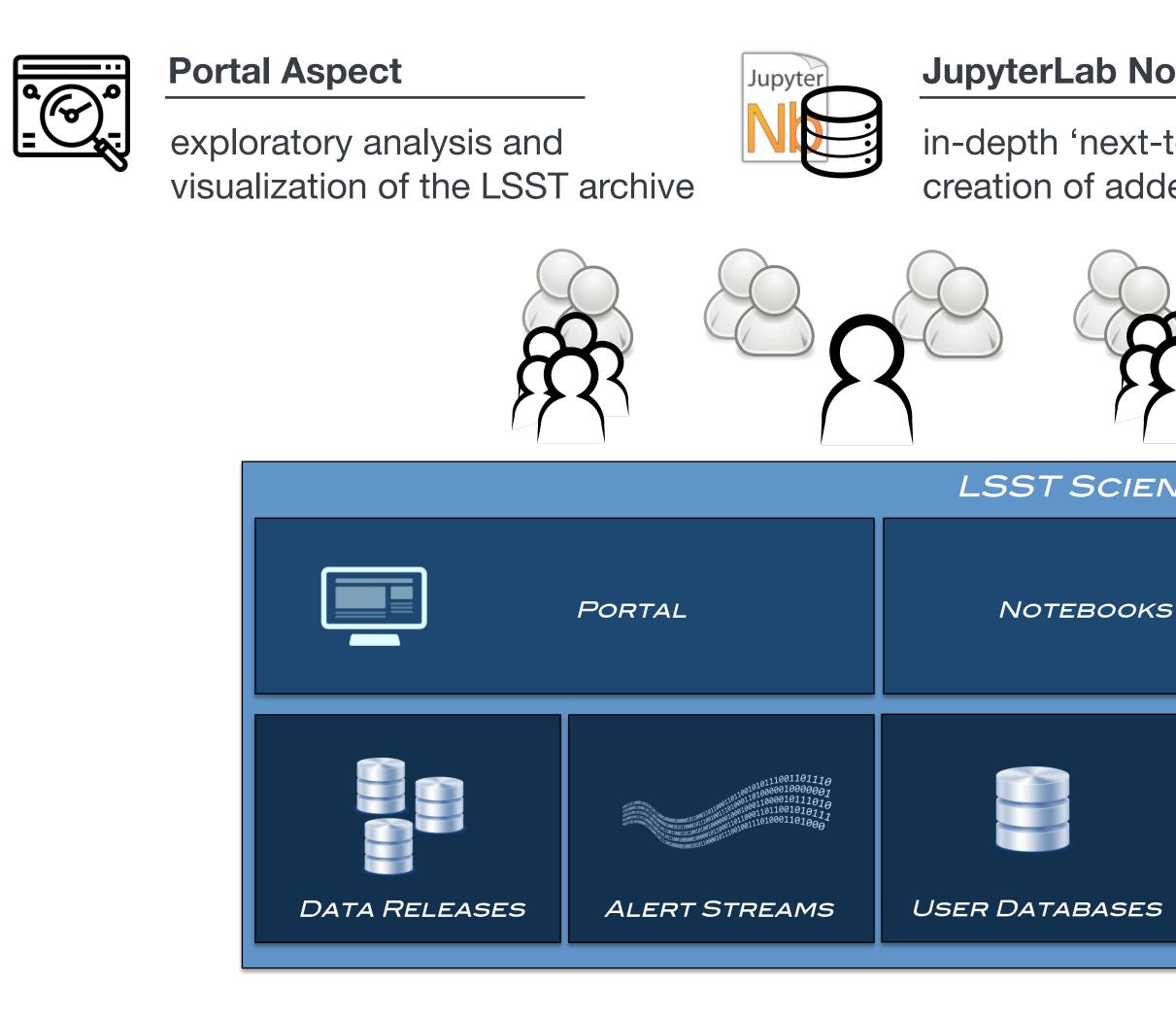
Provides access to LSST Data Products and services for all science users and project staff





The LSST Science Platform

A set of integrated web applications and services deployed at LSST Data Access Centers (DACs) through which the scientific community will access, visualize, subset and perform next-to-the-data analysis of the LSST Data products.



JupyterLab Notebook Aspect

in-depth 'next-to-the-data' analysis & creation of added-value data products Web API Aspect

remote access to the LSST archive via industry-standard APIs

